# Gypsic Soils on the Kane Alluvial Fans, Big Horn County, Wyoming

## U.S. GEOLOGICAL SURVEY BULLETIN 1590-C





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By MARITH C. REHEIS

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SOIL CHRONOSEQUENCES IN THE WESTERN UNITED STATES

# DEPARTMENT OF THE INTERIOR DONALD PAUL HODEL, Secretary

U.S. GEOLOGICAL SURVEY Dallas L. Peck, Director



#### UNITED STATES GOVERNMENT PRINTING OFFICE, WASHINGTON: 1987

For sale by the Books and Open-File Reports Section U.S. Geological Survey Federal Center, Box 25425 Denver, CO 80225

#### Library of Congress Cataloging-in-Publication Data

Reheis, Marith C. Gypsic soils on the Kane alluvial fans, Big Horn County, Wyoming.

(Soil chronosequences in the western United States) (U.S. Geological Survey Bulletin 1590–C) Bibliography Supt. of Docs. No.: I 19.3:1590–C

1. Soils—Wyoming—Big Horn County—Gypsum content.
2. Soils—Gypsum content.
3. Soil formation—Wyoming—Big Horn County.
4. Soil chronosequences—Wyoming—Big Horn County.
5. Geology—Wyoming—Big Horn County.
6. Title.
7. Title.
8. Title: Kane alluvial fans, Big Horn County, Wyoming.
8. Series.
8. Series: U.S. Geological Survey Bulletin 1590—C.

QE75.B9 No. 1590-C

557.3

86-600172

[S599,W8]

[631.4′7′78733]

#### **FOREWORD**

This series of reports, "Soil Chronosequences in the Western United States," attempts to integrate studies of different earth-science disciplines, including pedology, geomorphology, stratigraphy, and Quaternary geology in general. Each discipline provides information important to the others. From geomorphic relations we can determine the relative ages of deposits and soils; from stratigraphy we can place age constraints on the soils. Field investigations and mineralogic and sedimentologic studies provide information on the nature and types of deposits in which soils form. As a result of our work, we have estimated rates of soil formation, inferred processes of soil formation from trends in soil development with increasing age, and obtained information on the types of weathering that occur in various areas. In return, soil development and soil genesis have provided data on the age of landforms, the timing and duration of sedimentation, and, in some cases, the history of climatic fluctuations.

Between 1978 and 1983, a coordinated and systematic study was conducted on soil development in different types of geologic deposits in the Western United States. The goals of this project, led by the late D.E. Marchand and subsequently by M.N. Machette, were to learn whether rates of chemical, physical, and mineralogic transformations could be determined from soil chronosequences; how these rates vary in different mineralogic and climatic environments; and how accurately soils can be used for such problems as estimating the ages of deposits, periods of landscape stability, and timing of fault movements. This series of reports presents data from several soil chronosequences of that project.

More than 100 analyses on more than 1,000 samples were performed on soils collected in the Western United States. Some results have appeared in various books, journals, and maps (for example, Harden and Marchand, 1977, 1980; Burke and Birkeland, 1979; Dethier and Bethel, 1981; Marchand and Allwardt, 1981; Meixner and Singer, 1981; Busacca, 1982; Harden, 1982a, b; Harden and Taylor, 1983; Machette, 1983; Machette and Steven, 1983; Busacca and others, 1984; Machette and others, 1984; Reheis, 1984). In the reports in this series, the basic field information, geologic background, and analytical data are presented for each chronosequence, as well as some results additional to the previous publications.

One of the most significant aspects of these chronosequence studies is that in every study area, many soil parameters change systematically over time, or with the age of deposits. As Deming (1943) emphasized, it is this recurrence of correlation in such different conditions that is most significant to geologic and pedologic studies. In relatively moist areas, such as coastal and central California, such soil properties as percent clay or reddening of soil colors change most systematically over time. In more arid regions, such as in the Bighorn Basin of Wyoming, calcium carbonate and gypsum contents best reflect the relative ages of the deposits. A few parameters—for example, elemental composition of sands or clays—appear to be comparable between areas so diverse in climatic setting.

Numeric age control has enabled us to estimate rates of soil development. In some places, we have been able to compare rates between different areas. For example, in central California, rates of clay accumulation were found to be most rapid during the initial stages of soil development; these rates declined with increasing age. The straightest lines for regression were on a loglog scale. In coastal California, rates of clay accumulation appeared to be much higher than in central California. This difference in rates could be due to parent material (the coastal soils that we studied were formed on reworked shale and sandstone, whereas central California soils were developed in granitic alluvium), and (or) the differences in rates could be due to colian additions of clay. In the Bighorn Basin of Wyoming, rates of clay accumulation, as well as most other soil properties, increased linearly over time, with no apparent decrease in initial rates.

The data we present here suggest many opportunities for further interpretation. For example, we may learn how climate, vegetation, and mineralogy affect the rates of clay formation or organic-matter accumulation. In some study areas, we present data for rare-earth elements, which could be used to examine how each element reacts in different weathering environments. These examples are only a fraction of the possible future studies that could be conducted on the data presented here.

J.W. Harden Editor

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### Gypsic Soils on the Kane Alluvial Fans, Big Horn County, Wyoming

By Marith C. Reheis

ABSTRACT

Soils on the Kane alluvial fans, in the northeastern Bighorn Basin near Lovell, Wyo., accumulate gypsum (CaSO<sub>4</sub>·2H<sub>2</sub>O) over time. This gypsum is added chiefly as eolian dust derived from both local and long-range sources. The arid but cool climate has permitted gypsum to accumulate continuously for the past 600 ka (thousands of years). The linear increase in pedogenic gypsum with time suggests that effective moisture in the Kane area has never increased substantially in the past 600 ka. The dates when the fan surfaces stabilized and, thus, the times when the soils began to form are estimated by various dating methods, including tephrochronology, paleontology, and stream-incision rates.

The gypsum content influences the development of other soil properties. At the Kane fans, pedogenic or secondary gypsum has accumulated at annual rates of 1.1 X 10<sup>-6</sup> to 120 X 10<sup>-6</sup> g/cm<sup>2</sup>; modern aerosolic gypsum measured in dust traps is being added at annual rates of 26 X  $10^{-6}$  to 60 X  $10^{-6}$  g/cm<sup>2</sup>. Secondary calcium carbonate (CaCO3) does not increase in soils older than 100 ka, despite the availability of CaCO3 in eolian dust and in the parent material, probably because gypsum suppresses the solubility of CaCO3. Total texture (a measure of pedogenic silt and clay) and pH increase over time in soils younger than 300 ka, but decrease or show irregular trends in older soils because of large accumulations of pedogenic gypsum. In contrast to soils in most environments, rubification (reddening and brightening of soil colors) in the study area increases with depth. Gypsum also causes physical disintegration of gravel clasts in the soils by crystallizing in cracks and along grain boundaries.

Depth to gypsum in soils on the Kane fans is controlled by climate, soil age, and the hygroscopic properties of gypsum. The youngest soil containing visible gypsum and those soils older than 100 ka have their maximum gypsum accumulations at the average depth (2.6 cm) to which the modern precipitation moistens the soil. The 100-ka-old soils have their maximum gypsum accumulations at a greater depth (more than 5.5 cm). This disparity between gypsum depths probably reflects both differences in climate and the near-surface accumulation of gypsum during interglacial periods. Buildup of pedogenic gypsum at shallower depth during interglacials may have hindered subsequent leaching and reprecipitation of gypsum at greater depth during glacial periods of increased soil moisture, because the gypsum can hygroscopically hold the increased moisture. Thus, the near-surface accumulations of gypsum in soils older than 100 ka mainly reflect interglacial wetting depths.

#### INTRODUCTION

Studies of soils formed on the Kane alluvial fans<sup>1</sup>, on the northeastern margin of the Bighorn Basin, yield valuable information on the processes and rates of soil development in a cool, arid climate. This study assesses the effects of time on soil development in a chronologic sequence of soils. To ascribe the characteristics of soils to their ages, it is necessary to demonstrate conclusively that the other four important soil-forming factors—parent material, topography, vegetation, and modern climate, as defined by Jenny (1941)—are constant or have little effect on these characteristics.

Many studies of aridic soils in the southwestern United States (for example, Gile and others, 1966; Gardner, 1972; Bachman and Machette, 1977; Shlemon, 1978; Machette, 1985) have demonstrated that the dominant pedogenic process in that region is the accumulation of calcium carbonate (CaCO3), most of which is derived from dust and rainfall. In contrast, gypsum (CaSO $_4$ \*2H $_2$ O) accumulates in soils in the arid environment of the Kane fans, and secondary CaCO $_3$  is negligible. Accumulation of soluble salts characterizes hot desert soils (Watson, 1979) and polar desert soils (Bockheim, 1980a), but gypsic soils have been little studied in the United States (Nettleton and others, 1982). There are few studies of soils in arid parts of the intermontane basins of the northern Rocky Mountains (for example, Reider and others, 1974, in the Laramie Basin, Wyo.). Information on the soils formed on the Kane fan may help to link soil-forming processes in warm deserts and polar deserts.

Interpretations in this paper rely chiefly on field data and gypsum analyses. Some particle-size and chemical data are used to discuss the parent material. Reheis (1984) discussed other laboratory data, including particle size and chemical and mineralogic analyses.

Acknowledgments.—My greatest appreciation goes to P. W. Birkeland, who critically read most of this manuscript, for discussion and thoughtful evaluation of many ideas. K. L. Pierce, D. W. Moore, and L. D. McFadden also contributed valuable perspectives. The late D. E. Marchand coerced me into studying gypsic soils. Able field assistance was provided by Kim Jones.

<sup>&</sup>lt;sup>1</sup>The Kane alluvial fans are informally named for an abandoned railway station across the Bighorn River, which also gives its name to the Kane, Wyo., 7 1/2-minute quadrangle within which the fans lie.

#### Geomorphic Setting

The Kane fans are alluvial fan deposits of Cottonwood Creek, an ephemeral stream draining a 45-km<sup>2</sup> catchment basin on the west front of the Bighorn Mountains. The fans are sandwiched between the mountain front and the confluence of the Bighorn and Shoshone Rivers (figs. 1, 2). Lovell, the nearest town, is situated on the Shoshone River about 20 km upstream from the fans.

The relation of fan deposition to fluvial-terrace deposition in the study area must be understood in order to constrain the ages of the fan surfaces. Only two dates have been obtained from materials within the fan deposits. However, the fan and underlying Bighorn River deposits are related to regional glacialinterglacial climatic fluctuations that provide a chronologic framework for these deposits. gradient and depositional regime of the unglaciated Cottonwood Creek has been controlled by the positions of the Bighorn and Shoshone Rivers. The Bighorn River is the master stream in the Bighorn basin. It drains the glaciated Wind River Range southwest of the study area, and is fed by streams draining the west flank of the Bighorn Mountains and glaciated areas of the Absaroka Range. The Shoshone River heads in the Absaroka Mountains to the west (fig. 1).

The formation of the Shoshone River terraces has long been argued. Mackin (1937) believed that the terraces were formed by lateral planation of the Shoshone River and were not of glaciofluvial origin. On the Shoshone River, John Moss (Moss and Bonini, 1961; Moss and Whitney, 1971; Moss, 1974) later demonstrated that deposits of the lowest major terrace (Cody terrace of Mackin, 1937), and possibly the next higher terrace (Powell terrace of Mackin, 1937), can be traced upvalley into moraines and constitute fill terraces near the front of the Absaroka Downstream, these are strath terraces Range. comprised of about 5 m of gravel overlying bedrock surfaces. Moss (1974) believed that the spreading of gravel into the basin, whether deposited as fill or as strath terraces, accompanied or slightly postdated a maximum glacial advance in the Absaroka Range, and subsequent incision was associated with deglaciation and interglacial conditions. Palmquist (1983) suggested, however, that incision episodes of the Bighorn River are out of phase with episodes on its tributaries; for example, the Bighorn may incise during glaciations even as the Shoshone River aggrades. Some terraces may have formed by Recently, Ritter and Kaufman (1983) demonstrated that the Powell terrace as previously mapped below the town of Powell is younger than the Powell terrace above the town, and was formed by capture.

The regime of the Bighorn River is probably controlled by climate. Near the headwaters area of the Bighorn (Wind) River, an outwash terrace sequence suggests eight major glacial advances (Richmond, 1976).

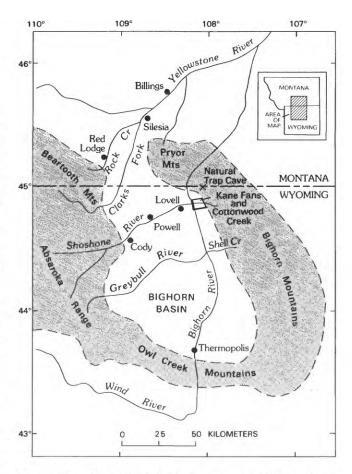


Figure 1. Location of study area in the northern Bighorn Basin (modified from Mackin, 1937).

#### Modern Climate and Vegetation

Climate is one of the most important factors that govern soil development. Weather data from Lovell, Wyo. (National Climatic Center, 1976), and data from soil temperature probes planted at the study site, represent the present climatic regime near the Kane fans (table 1). Although the fans are about 20 km from Lovell, the climate of the study area is probably similar to that of Lovell because the vegetation is similar. At the Kane fans, sagebrush, pricklypear cactus, and bunchgrass are the dominant species. Scrub juniper occurs on the mountain footslopes above the fans, and cottonwoods grow along the Bighorn and Shoshone Rivers.

Judging from the vegetation, all soil sites have about the same climate. Rainfall generally increases with elevation (orographic effect), but the distribution of deposits and limited land access did not permit all sample sites to be at the same elevation. Soil sites on the younger fans (fig. 2) are lower and farther from the mountain front, and these sites, therefore, may be slightly drier than those on the older, higher fans (lowest soil site is B-5 on fan 2 at 1,117 m; highest site is B-19 on fan 7 at 1,275 m).

The Lovell area (table 1) has an aridic soilmoisture regime and a mesic soil-temperature

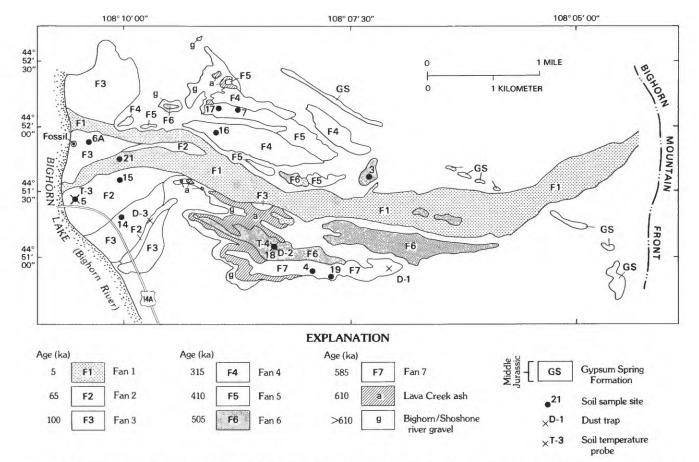


Figure 2. Generalized map of the Kane alluvial fan gravels of Cottonwood Creek and fluvial deposits of the Bighorn and Shoshone Rivers. Outcrop of gypsum beds of the Middle Jurassic Gypsum Spring Formation is also shown. Map unit F1 includes the modern channel and flood plain of Cottonwood Creek. Fluvial gravels under fan 3 not shown because they are exposed in vertical cliffs. Soil site numbers correspond to those preceded by "B-" in text.

Soil temperature was estimated from the regime. mean annual temperature (Soil Survey Staff, 1975, p. 61) and also measured by soil temperature probes. The estimated soil temperature of 8.20 C is near the boundary between the mesic and frigid temperature The soil temperature probes, probably regimes. affected by local conditions, 2 recorded temperatures about 40 C warmer than the estimated soil temperature (Trembour and others, 1986).

Near Lovell, much of the moisture falls in late spring and early summer and is ineffective in because weathering or solution of salts, evapotranspiration rates are high at these times. Using Arkley's (1963) leaching index for Lovell, only 26 mm of water is available to wet the soil from November to March (fig. 3). Thus, in the average year, only the surface of the soil is ever wet; insufficient water percolates through the soils to remove soluble salts and other weathering products.

#### Paleoclimate

Characterization of the modern adequately describes the general climate of the Holocene but fails to describe the climate during glacial periods. More moist glacial periods probably had important effects on soil development at the Kane fans, so it is important to discuss the available evidence concerning paleoclimate.

Sediments at Natural Trap Cave, 10 km north of and 350 m higher than the Kane fans (fig. 1), contain opal phytoliths of tall prairie grass (Gilbert and others, 1980). The modern vegetation around Natural Trap Cave consists of short grasses and cactus; tall prairie grass commonly occurs at elevations 300 m higher than the cave. The tall-grass phytoliths span the period from 110 thousand to 12 thousand years before present (ka) and indicate increased moisture during that time. Tundra phytoliths near the top of the sedimentary section suggest an episode of cold dry climate from 17 Environmental reconstructions based on faunas from several Wisconsin-aged sites in Wyoming (summarized in Walker, 1982) and the presence of a musk-ox vertebra (discussed below) in the Kane fans indicate that a cold, dry arctic steppe/savannah biome existed on the basin floor.

<sup>&</sup>lt;sup>2</sup>Temperatures measured by soil-temperature probes can be affected by local conditions, such as the yearly range of temperatures, or air pockets left along the probe when refilling the probe hole.

Table 1. Climatic data for Lovell, Wyo., for 1951 to 1974, and postulated climatic parameters for the last glaciation

	Modern	Glacial
Mean annual precipitation	16.5 cm	16.5 cm
Mean annual air temperature	7.2°C	-2.8°C
Mean July air temperature	22.1°C	9.1°C
Mean January air temperature	-7.9°C	-12.9°C
Mean annual soil temperature		
from probes at Kane fans, 1980-82:		
at 0.5 m depth	12.1°C	
at 1.0 m depth	12.6°C	
Mean annual soil temperature,		
Soil Survey Staff calculation:	8.2°C	
Leaching index	2.6 cm	8.2 cm

This climatic reconstruction is supported by other work. Recent studies of relict ice and sand wedges in Quaternary deposits in the intermontane basins of Wyoming (Mears, 1981) indicate that the climate during the last glaciation was dry and cold; mean annual temperatures may have been 10-13° C colder than at present. This estimate of cooler temperature is compatible with glacial to present snowline changes if the effect of altitudinal gradients of precipitation is considered (Pierce, 1982). Gates's (1976) model of global ice-age climate suggests that mean July temperature decreased by 13°C near the Bighorn Basin, and that precipitable moisture decreased 35 percent in the northern hemisphere.

A plausible reconstruction of climate during the last glaciation can be made for the Lovell area, assuming that (1) the mean annual temperature was 10° C lower than present, (2) the mean July temperature was 13° C lower, and (3) the mean January temperature was 5° C lower (table 1). No evidence is available concerning glacial winter temperature, but some lowering probably occurred. For the reconstruction, temperatures for other months are adjusted to fit the postulated changes of mean annual, July, and January temperatures. No moisture decrease is postulated in this reconstruction, in order to avoid extrapolation of hypothetical moisture changes in the northern hemisphere (Gates, 1976) to the study area. Evapotranspiration is calculated from monthly temperature data using van Hylckama's (1959) Diagrams of soil-temperature and soilmoisture balance compare glacial with modern conditions (fig. 3).

In the climate reconstruction (table 1 and fig. 3), the glacial soil-moisture regime probably was aridic and the soil-temperature regime pergelic (modern pergelic regimes correspond to permafrost zones; terms defined in Soil Survey Staff, 1975). The estimated glacial leaching index (method of Arkley, 1963) increases to 8.2 cm from the modern level of 2.6 cm because evapotranspiration is lower. Some of the additional available moisture may not have been utilized in leaching because more precipitation would have fallen as snow. If snowfall from November to March did not contribute to soil moisture, the glacial leaching index was 5.3 cm. Gates's (1976) climatic model also suggests that the glacial leaching index was

lower than 8.2 cm. The presence of a large stable high-pressure zone over the continental ice sheet 350 km northwest of the study area probably would have caused a year-round dominance of dry northerly air flow, whereas the present southerly air flow in spring and summer delivers moisture from the Gulf of Mexico.

#### GEOLOGIC SETTING AND CHRONOLOGY

#### **Previous Work**

Andrews and others (1947) were the first to map the Kane fans, but called them pediment surfaces. Birdseye (1980, 1983) described the alluvial fans of Five Springs Creek, 5 km south of the Kane fans, and has mapped the Kane fans as well (written commun., 1982). Palmquist (1978, 1983) used the Lava Creek ash bed (Izett and Wilcox, 1982) that underlies the Kane fan gravels to date terraces along the Bighorn River. Bedrock geology near the Kane fans was described by Darton (1906) and later by Rioux (1958) and Pierce (1978).

#### Source of Deposits

Bedrock units in the Cottonwood Creek drainage basin range in age from Cambrian to Cretaceous and include marine shale and sandstone, evaporites,

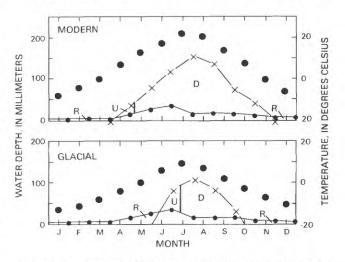


Figure 3. Modern (24-year record) and last-glacial climate and soil-water balance for Lovell, Wyo. Reconstruction of glacial temperature described in text. No surplus water available during either period. Large dots are air temperatures, small dots are precipitation, and X's are inferred evapotranspiration. Soil AWC (available water-holding capacity) of 110 mm is average for Kane fan soils; young soils have lower AWCs, older soils have higher AWCs. R denotes period of recharge of soil water when precipitation exceeds evapotranspiration; U denotes period of utilization of stored soil water when soil is dry in some part, because evapotranspiration exceeds precipitation; D denotes deficit period when soil-moisture control section is dry.

redbeds, limestone, and dolomite (Pierce, 1978). The Kane fan deposits overlie Cretaceous sandstone and shale but contain little debris from these formations. Continental redbed and evaporite deposits of Mesozoic age constitute 40 percent of the local sedimentary section. The Middle Jurassic Gypsum Spring Formation has up to 50 percent by volume of lenticular gypsum beds and crops out just upslope from the fan deposits (fig. 2). Little or no gypsum was deposited in the fans, however (table 2 and supplementary table 3). Paleozoic limestone and dolomite account for another 40 percent of the section. Because these latter rocks are resistant to weathering and mechanical abrasion, they constitute the majority of the fan material.

A second component of the fan deposits is andesitic sand and gravel, which was deposited by the Bighorn and (or) Shoshone Rivers and was incorporated into the fans by reworking.

#### Parent Material—Nature and Variability

The parent material is another important factor in soil development. If the parent materials of the soils on the Kane fans are different, then differences in the soils cannot be ascribed to either time or climate. Thus, it is critical to compare stable chemical and physical properties of the various alluvia.

The parent material among all soils examined on the Kane fans is compositionally similar (table 2). By visual estimate, the gravel consists of about 90 percent limestone and dolomite, 5 percent andesite, and 5 percent sandstone, siltstone, and chert. Carbonates usually compose about 50 percent of the unweathered <2-mm fraction (supplementary table 3). Soils on fan 5 deposits have slightly greater proportions of igneous detritus, because they contain more reworked river gravel and volcanic ash.

Fresh fan deposits in the Kane area typically have 80 percent gravel, but the upper 20-75 cm of the soils is less gravelly. Deposits in the youngest fan are very gravelly up to 5 cm below the surface. The thickness of the less gravelly layer roughly increases with fan age (supplementary tables 1 and 2). This trend suggests that the finer material at the surface is a result of soil-forming processes rather than deposition.

Chemical uniformity of parent material both within and between soil profiles was examined using titanium-oxide to zirconium-oxide ratios of the silt-plus-clay fraction (TiO<sub>2</sub>/ZrO<sub>2</sub>; data in supplementary table 5). Titanium and zirconium are primarily contained within the minerals rutile and zircon, which are very resistant to weathering (zircon is most resistant). Generally, the proportion of TiO<sub>2</sub> to ZrO<sub>2</sub> should reflect the composition of the parent material; thus abrupt changes in the TiO<sub>2</sub>/ZrO<sub>2</sub> ratio with depth probably reflect changes in the parent material chemistry rather than changes due to weathering (Chapman and Horn, 1968; Smeck and Wilding, 1980).

 ${
m TiO_2/ZrO_2}$  ratios for all horizons of four profiles (B-21, B-14, B-17, and B-19; fans 1, 3, 5, and 7 respectively) range from 5.1 to 11.9, with a mean (x) of 8.9 and standard deviation (s) of 2.0 (table 3). No ratios exceed 2s from the mean, and only six ratios

Relative variability among (N = 24) exceed 1s. different groups is examined by the coefficient of variability  $(\underline{s}/\underline{x})$ , which for  $\text{TiO}_2/\text{ZrO}_2$  ratios of all horizons is 23 percent. Because Ti does weather, some of this variation may be caused by soil-forming processes (Chapman and Horn, 1968; Wilding and others, 1971). If only C-horizon TiO<sub>2</sub>/ZrO<sub>2</sub> ratios are considered, the mean ratio is 10.1 and the standard deviation is 1.6, for a between-profile variability of 16 percent. Within-profile variability ranges from 11 to 25 percent, averaging 19 percent. These ratios show that the between-profile variability is similar to the within-profile variability. TiO<sub>2</sub>/ZrO<sub>2</sub> ratios tend to be larger at depth, but the increase is not uniform and the changes are gradual. Because ratios for all horizons fall within 2s of the mean ratio, the variation is considered normal for the population, which suggests that the soils are chemically similar.

Textural variability is estimated by calculating ratios of coarse to very coarse sand (C/VC) and medium to coarse plus very coarse sand (M/C+VC) for C horizons. Near-surface sand fractions of Kane fan soils tend to be distinctly coarser than sands in the B and upper C horizons, especially in older soils (data not shown). In contrast to the sand fractions, the gravel clasts do not increase in size in the near-surface Deflation of surface fines is an unlikely cause of the near-surface coarseness of the sand fractions, because the thickness of the less gravelly surface layer increases with time (supplementary tables 1 and 2), suggesting that fines accumulating. Because the inverse grading of sand grains seems most marked in older soils, the surface material may be more coarse than deeper sediments due to the breakdown of sand grains by gypsum accumulation at depth (discussed below). Because the B-horizon sand fractions appear to have been affected by weathering, only sand fractions from C horizons are compared.

The sand-fraction ratios in most cases are based on measurements after removal of carbonates (including parent material) and gypsum. This step should result in greater uniformity in the ratios than in untreated soils (table 3). Sand-fraction ratios for four profiles (B-5, B-7, B-3, and B-4) are more variable among C horizons with carbonates (80 percent for C/VC and 76 percent for M/C+VC) than among C horizons without carbonates (38 percent and 53 percent respectively, table 3). Remarkably, within-profile variability is the same for both sets of profiles.

Of the 39 C horizons in 13 profiles examined, none have C/VC or M/C+VC ratios that exceed 2s from the mean ratio; eight horizons exceed 1s from the mean ratio for either C/VC or M/C+VC, but not both. These horizons are considered to lie within the range of normal variability for this population. Considering the textural variation that is possible on alluvial fans, the sand fractions are unusually uniform both with depth and with distance from the mountain front (variability about 40-80 percent). For comparison, Merced River soils (Harden, 1982a) average 44 percent for within-profile variability of both sand-fraction ratios, whereas between-profile variability is 114 percent for M/C+VC and 90 percent for C/VC ratios.

In summary, the fan alluvia are chemically and lithologically similar, but do vary somewhat in grain

Table 2. Field description of Kane fan deposits relative to distance from fan head

[Proximal is nearest remnant to fan head, distal is farthest remnant from fan head. Refer to figure 2 and to U.S.G.S. 7-1/2-minute Kane and Cottonwood Canyon, Wyo., quadrangles for approximate locations. Percent lithologies and mean clast diameters were visually estimated; maximum clast diameters were measured. Mean diameters are not given for poorly exposed sites. No proximal exposures are available for fans 2, 3, and 4]

Fa	an Description	Estimated	Mean clast	Maximum clast	Gypsum	
num	nber of exposure	thickness (m)	diameter (cm)	diameter (cm)	stage	
	Proximal deposits					
l	Arroyo, NW/4SW/4SW/4 sec. 6, T. 56 N., R. 93 W; poor exposure Crude bedding, poorly sorted. 95 percent carbonates (carb.), 5 percent other sedimentary (sed.) rocks.	. 2	10	70	0	
•	SE/4NE/4NE/4 sec. 2, T. 56 N., R. 94 W.; poor exposure. Grave overlies shale. No bedding visible; poorly sorted, gravelly 90 percent carb., 10 percent other sed. rocks. Gypsum decreases with depth.			60	II~II)	
•	SW/4SE/4SW/4 sec. 6, T. 56 N., R. 93 W.; poor exposure. No bedding visible; poorly sorted, gravelly. 95 percent carb., 5 percent other sed. rocks. No visible gypsum at base of gravel.	15		100	0-IV	
	Poor exposure, at soil site B-19. Gravel overlies shale. No bedding visible; poorly sorted, gravelly. 95 percent carb., 5 percent other sed. rocks. No visible gypsum at base of gravel.	10		1 50	0-IV	
	Distal deposits	<del></del>				
	0.2 km north of Kane highway bridge. Crudely bedded; some 10-cm-thick lenses of sand and silt intercalated with poorly sorted, imbricated gravel. 85 percent carb., 10 percent other sed. rocks, 5 percent andesites (and.).	>3	15	45	0	
	Low river bluff 0.1 km north of bridge; gravel overlies shale. Crudely bedded; discontinuous 5-cm-thick lenses of sand and silt intercalated with poorly sorted imbricated gravel lense 85 percent carb., 10 percent other sed. rocks, 5 percent and	S.	10	45	0	
3	High river bluff 0.8 km north of bridge; fan gravel overlies lof fluvial gravel above shale. Crudely bedded; lenses of finer material up to 50 cm thick intercalated with poorly sorted imbricated gravel. 85 percent carb, 10 percent other sed. rocks, 5 percent and. Gypsum decreases with depth.		5	25	0-1	
	Poor exposure, NW/4NW/4 sec. 3, T. 56 N., R. 94 W.; gravel overlies shale. No bedding visible; poorly sorted, gravelly 90 percent carb., 10 percent and. Gypsum decreases with dep			80	0-11	
	Poor exposure, NE/4NW/4 sec. 3, T. 56 N., R.94 W.; gravel overlies shale. No bedding visible; poorly sorted, gravelly 90 percent carb., 10 percent and. Gypsum decreases with dep				0-111	
	Poor exposure, NE/4NW/4 sec. 3, T. 56 N., R. 94 W.; fan gravel overlies thin fluvial gravel over shale. No bedding visible; poorly sorted, gravelly. 90 percent carb., 10 percand. Gypsum decreases with depth.				0 <b>-1</b> V	
,	Corner secs. 2, 3, 10, 11, T. 56 N., R. 94 W.; fan gravel overlies 30 m of ash. Crudely bedded; 25-cm-thick lenses of fine material intercalated with poorly sorted imbricated gravel. 90 percent carb., 10 percent and. Gypsum decreases with depth.	15	15	50	0-IV	

size. Textural variability is greater than chemical variability due to the nature of fluvial processes, assuming that different size fractions have similar

chemical compositions. Hence, differences among the soils developed on the Kane fan surfaces reflect differences in age and paleoclimate.

Table 3. Parent-material variability within and among soils of the Kane fans

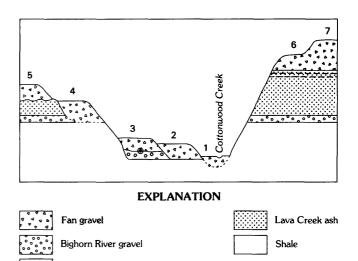
 $[\underline{N}, \text{ number of cases; } \underline{x}, \text{ mean; } \underline{s}, \text{ standard deviation; } \underline{s/x}, \text{ relative variability in}$ percent. M, medium sand; C, coarse sand; VC, very coarse sand. Sand ratios are given both for samples with carbonate and without carbonate (gypsum was removed from all samples)

Variability	N	<u>x</u>	s	$\underline{s}/\underline{x}(percent)$
	Samples	without car	bonates	
Among profiles:	*****			
C/VC, C horizons M/C+VC, C horizons	39 38	1.2 .8	0.5 .5	38 53
Within profiles:				
mean of $\underline{s/x}$ for C/VC mean of $\underline{s/x}$ for M/C+VC	13 13			47±25 (1 <u>s</u> ) 51±21 (1 <u>s</u> )
	Sample	s with carbo	onates	
Among profiles:				
TiO <sub>2</sub> /ZrO <sub>2</sub> , all horizons TiO <sub>2</sub> /ZrO <sub>2</sub> , C horizons	24 12	8.9 10.1	2.0 1.6	23 16
C/VC, C horizons M/C+VC, C horizons	12 12	1.9 .77	1.5 .59	80 76
Vithin profiles:				
mean of $\frac{s/x}{s/x}$ for $TiO_2/ZrO_2$ mean of $\frac{s/x}{s/x}$ for $C/VC$ mean of $\frac{s/x}{s/x}$ for $M/C+VC$	4 4 4	 	 	19± 6 (1 <u>s</u> ) 48±14 (1 <u>s</u> ) 49±24 (1 <u>s</u> )

#### Stratigraphic Units

I recognize seven prominent surfaces that are developed on the Kane fans (figs. 2 and 4). The surfaces are numbered 1 to 7 from youngest to oldest. For brevity, these surfaces will henceforth be referred to as fan 1, fan 2, and so forth. Other minor surfaces and deposits occur in the area but are omitted from this study. Correlation of discontinuous deposits was based on both height above Cottonwood Creek and stratigraphic relations. Each fan surface overlies gravel beds resting either on a bedrock surface or on pre-existing fan or other fluvial deposits. Because the Bighorn River is the local base level for Cottonwood Creek, the fan surfaces probably are graded to former levels of the Bighorn-Shoshone river system.

A schematic cross section (fig. 4) sketches the relations of the fan deposits. The fan deposits are 5-15 m thick near the Bighorn River; exposures on fans 6 and 7 nearest the mountain front indicate little thinning toward the mountains (table 2). The deposits underlying fans 2, 4, and probably 1 rest on cut surfaces of Cretaceous shale and sandstone. deposits underlying fans 3, 6, and 7 grade into the underlying fluvial deposits of the Bighorn-Shoshone river system. Where exposed, andesitic river gravel grades upward, over a 1- to 3-m interval, into calcareous fan deposits. The oldest fan deposits



Schematic cross section showing relationships of Quaternary deposits. Thickness of deposits not to scale; maximum thickness of Quaternary deposits under fan 7 is 55 m, including 15 m fan gravel, 35 m ash, and 5 m river gravel. Cottonwood Creek is presently incised about 13 m below the base of the ash.

Silty sand

Musk ox fossil

conformably overlie fluvially deposited Lava Creek volcanic ash that was erupted from the Yellowstone National Park area (Izett and Wilcox, 1982). The crossbedded ash and pumice deposits, as much as 35 m thick, overlie and are interbedded with andesitic riverchannel gravel (Birdseye, 1983). Remnants of the channel gravel and overlying ash extend north from fan 7 to fan 5 but pinch out toward the mountain front (fig. 2), suggesting that the Bighorn River formerly flowed about 2-3 km east of its present course. Fan 5 gravel unconformably overlies the fluvial ash and gravel on the north side of the study area, hence this fan gravel postdates the deposition of the material composing fans 6 and 7.

The Kane fan gravels consist of crudely bedded lenses of sand and silt intercalated with imbricated, poorly sorted gravel lenses (table 2). exposures of fans 1, 2, 3, and 7, near the Bighorn River, reveal that visible gypsum decreases from the surface downward and does not occur at the base of the gravels. Gypsum does occur at the base of the thin (3-5 m) gravel deposits of fans 4 and 5 but is less

abundant in the lower parts. Maximum clast size is greater toward the head of each fan (table 2), but little change occurs in gravel thickness, bedding, or sorting.

#### Geochronology

The difference in age of the fan surfaces, and to some extent paleoclimate, have determined the differences observed among the Kane fan soils. In order to study time-dependent changes in soils on the Kane fans, estimates of the time of stabilization of each surface must be obtained. Tephrochronology, paleontology, and incision rates have been used to estimate the age of the deposits and their surfaces (table 4). The ages of fans 1, 3, and 7 are discussed first because these ages are most closely constrained. From these ages, incision rates are calculated in order to estimate the ages of fans 2, 4, 5, and 6.

Fan 1 is the most recently abandoned surface (figs. 2 and 4) and still exhibits bar-and-swale

Table 4. Stratigraphic units and ages of the Kane fan soil chronosequence

[Height column gives mean and standard deviation ( $\underline{N}$  = 4-7) of height of fan surface above Cottonwood Creek used in figure 5]

Geologic unit	Age con- trol (ka)	Hest estimate and range for surface age (ka)	Height (m)	Comments
Fan l	0 110	5 (0-10)	3.0±0.6	Post-last glaciation.
Fan 2	<sup>2</sup> 20 <sup>2</sup> 100	<sup>3</sup> 65 (20–100)	6.6±0.6	Younger than fan 3, older than fan 1.
Fan 3	<sup>4</sup> 70 <sup>5</sup> 100 <sup>6</sup> 175	100 (70-150)	8.4±2.0	Correlated with Yellowstone- type volcanic ash; musk-ox vertebra in basal fan gravel.
Fan 4	<sup>2</sup> 2 30 <sup>2</sup> 400	<sup>3</sup> 315 (230-400)	20.8±1.3	Older than fan 3, younger than fan 5.
Fan 5	<sup>2</sup> 320 <sup>2</sup> 480	<sup>3</sup> 410 (320–480)	25.9±1.3	Older than fan 4, younger than fan 6.
Fan 6	<sup>2</sup> 415 <sup>2</sup> 570 <sup>7</sup> 610	<sup>3</sup> 505 (415–570)	31.4±1.3	Overlies Lava Creek ash; younger than fan 7.
Fan 7	<sup>8</sup> 542 <sup>7</sup> 610	585 (540-600)	35.8±1.6	Overlies Lava Creek ash; older than fan 6.

<sup>&</sup>lt;sup>1</sup>Date from obsidian hydration on Pinedale-age deglacial deposits, West

Yellowstone, Montana (Pierce and others, 1976).

<sup>2</sup>Estimated incision rate of Cottonwood Creek, calibrated from fan surfaces 1, 3,

and  $_{7}^{7}$  (fig. 5 and text discussion).  $^{3}\mathrm{Best-estimate}$  age derived from incision-rate calculations as described in text;

<sup>5-</sup>ka accuracy of the ages is not implied.

4 Youngest possible ash eruption from Yellowstone area (K.L. Pierce, oral commun., 1984); ash contained in colluvial deposits overlying alluvium correlated to that which underlies fan 3.

Superhydration of volcanic ash (described in 4 above) on Shell Creek, Wyo.

<sup>(</sup>Palmquist, 1983).

Earliest possible arrival of musk ox in North America (C.A. Repenning, written commun., 1982). Lava Creek ash erupted from Yellowstone Park area (Izett, 1981).

<sup>&</sup>lt;sup>8</sup>End of interglacial stage following eruption of Lava Creek ash, as determined from marine oxygen-isotope curves (Shackleton and Opdyke, 1976).

morphology. The surface is 3 m above the modern Bighorn River channel. Moss (1974) believed that the Cody terrace of Wisconsin age merges with the Shoshone River floodplain 25 km above its junction with the Bighorn River, but small remnants of this terrace occur 10 km above the junction and are 6 m higher than the river (Mackin, 1937; Ritter and Kauffman, 1983). Because fan 1 is graded to a level only 3 m above the Bighorn River, it probably postdates the last glaciation; hence this surface is of Holocene age, or  $5 \pm 5$  ka (table 4).

The lower age limit of the fan 3 surface is constrained by a musk-ox vertebra (found by Lucy Piety) at the contact between the fan 3 deposits and the underlying river gravel (figs. 2 and 4). C. A. Repenning (U.S. Geological Survey, written commun., 1982) noted that modern musk oxen (genus Ovibos) did not appear in North America before the peak of the Illinoian glaciation (between 130 and 175 ka). Most modern musk-ox fossils (genus Ovibos) in the conterminous United States are considered to be of Wisconsin age, but in two localities, the fossils may be of Illinoian age (Walker, 1982). Thus, the vertebra at Kane constrains the lower age limit of fan 3 to 175 ka.

The lower age limit of the fan 3 surface is also post-Bull Lake glaciation, based on mapping, and on correlation by height above stream level of the underlying river gravel with that along the Bighorn and Shoshone Rivers. The underlying river gravel extends up the Shoshone River to Powell, where it constitutes the sub-Powell terrace of Ritter and Kauffman (1983) (also called the "Intermediate Terrace" by Mackin, 1937). The slightly higher Powell terrace has been traced upstream from Powell to possible moraines west of Cody (Moss, 1974) that are considered to be of Bull Lake age, dated at no younger than 125 ka in West Yellowstone (Pierce and others, 1976).

Fan 3 must be younger than the river gravel on which it lies. Palmquist (1983) estimated the age of the Bighorn River terrace that underlies fan 3 at about 90 ka, using the average incision rate of the Bighorn River for the past 600,000 years.

The upper age limit of fan 3 is further constrained by an ash deposit in sheetwash and colluvium overlying fluvial deposits on Shell Creek (fig. 1). These fluvial deposits were correlated to those underlying the fan 3 gravel by Palmquist (1978, 1983), based on similar heights above the Bighorn River. He assigned a tentative date of 100 ka, based on superhydration of the ash (dated by V. Steen-MeIntyre). The ash is not younger than 70 ka, which is the age of the youngest ash eruption from the Yellowstone Park area, and it is probably not older than 150 ka, the inferred age of one of the more widespread young ash eruptions from Yellowstone (K. L. Pierce, oral commun., 1984).

These age constraints indicate that the fan 3 deposits are about 100 ka old, and that the river gravel underlying fan 3 probably is of Bull Lake age. Hence, deposition of the fan gravel may have required about 25,000 years, which is the time interval between the probable youngest age of Bull Lake outwash (Pierce and others, 1976) and the age of the ash on Shell Creek in colluvium that overlies Bull Lake(?) outwash. I estimate the time of stabilization of the fan 3 surface to be  $100 \ ^{+50}_{-30}$  ka (table 4).

Deposits of fan 7 overlie the 610-ka Lava Creek ash (Izett, 1981; Izett and Wilcox, 1982). As neither unconformities nor buried soils were found at this ashfan gravel contact or within the fan material, the surface of fan 7 is probably not much younger than 600 ka. If aggradation is associated with glacial episodes and incision with deglaciation and interglacials, as Moss (1974) believed, then the fan gravel overlying the ash at Kane must have accumulated rapidly just prior to downcutting by the Bighorn-Shoshone river system. If accumulation of the fan 7 gravel took as long as that of fan 3 gravel, the surface of fan 7 may have been abandoned about 585 ka. Alternatively, if Palmquist's (1983) views are accepted, that the Bighorn River is stable or aggrades during interglacials and incises during glacials, then the fan gravel may have been deposited during an interglacial period. From the marine oxygen-isotope record (Shackleton and Opdyke, 1976), the interglacial that followed 610 ka ended about 540 ka. The fan 7 surface dates from about 585 +15 ka (table 4).

Age estimates for fans 1. 3. and 7 can be used to estimate ages of stabilization for the remaining fan surfaces (fig. 5 and table 4). The means and standard deviations of the heights above Cottonwood Creek of the three dated fans are plotted against age. The dashed line connecting the three points of known age yields an average stream-incision rate of 5.6 cm/1,000 yr for the past 585,000 years. When the means and standard deviations of heights above the creek of the remaining fan surfaces are plotted on the resultant line, approximate ages and age errors can be derived for the undated surfaces (method from Palmquist, 1979, 1983). The ages are as follows: fan 2, 65±10 ka; fan 4, 315±25 ka; fan 5, 410±25 ka; and fan 6, 505±25 ka. These are minimum age errors because they only

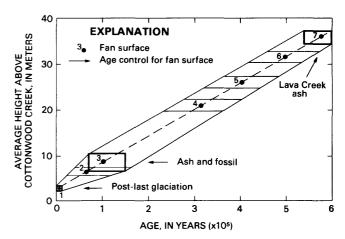


Figure 5. Graph for estimating fan surface ages from average heights above Cottonwood Creek and incision rates (dashed line) derived from ages for fans 1, 3, and 7. Boxes around points for fans 1. 3, and 7 show error limits of age (horizontal) and height (±1 standard deviation, vertical). Lines connecting outermost points of the boxes show the error envelope. Parallelograms show the age errors for other fans, derived by projecting their height errors (±1 standard deviation) horizontally to meet the lines of the error envelope.

include uncertainties in the fan heights above stream level.

maximum age errors also include uncertainties in the heights and ages of the dated The error envelope (fig. 5) permits estimation of the maximum errors on the ages of the undated surfaces. From figure 5, for example, the age of fan 2 is estimated to be 65 ka and to range from 20 ka to 175 ka. However, fan 2 is younger than fan 3, so the lower age limit is propably about 100 ka. The upper age limit of 20 ka for fan 2 is reasonable, because this fan surface is roughly the same height above the Bighorn River as the Cody (Pinedale-aged) terrace is above the Shoshone River. Age errors for fans 4, 5, and 6 can only be evaluated from their relative age order. For example, if the lower age limit of 400 ka for fan 4 is the true age, then the age of fan 5 is older.

Estimates of age based on incision rates involve two major assumptions: first is that the long-term incision rate has remained constant. In the short term, this assumption is false, because terraces and fans are formed during stable or aggrading fluvial conditions and isolated by later incision. The second assumption is that the lateral position of the Bighorn River has either remained stable or has moved in the same direction at a steady rate. If the river has moved irregularly, the fan surface slopes and elevations may vary unpredictably. Fluvial gravels underlying fans 7 and 3 indicate that the Bighorn River has shifted to the west, but whether the shift was gradual or episodic is not known.

#### SOIL DEVELOPMENT IN THE CHRONOSEQUENCE

The physical appearance or morphology of soils commonly reflects time-dependent soil properties. The Kane fan soils have light-colored, ochric A horizons with vesicular pores and desert pavements; in age sequence the soils progress from having only traces of gypsum, to having gypsic horizons, to having petrogypsic horizons (fig. 6). Soil morphology is commonly described with reference to the U.S. soil classification system (Soil Survey Staff, 1975). Unfortunately, this system does not discriminate the Kane fan soils with respect to their age. These soils have argillic (clay-rich) horizons, and possession of such horizons overrides differences in gypsum content in the process of classification. Translocated clay in Kane fan soils cannot be seen with a hand lens, but such clay surrounds calcite, dolomite, and quartz grains in petrographic thin sections of B horizons. Observations (Nettleton and others, 1969; Holliday, 1982) and experiments (Goss and others, 1973) have shown that clay can be translocated in calcareous soils if pore spaces or cracks are large enough. Thus, all the Kane fan soils are classified as Typic Haplargids (aridic soils with argillic horizons), regardless of their CaCO<sub>3</sub> or gypsum content.

#### Profile Indices

Some soil properties that change with soil age include texture (in this study, determined in the field

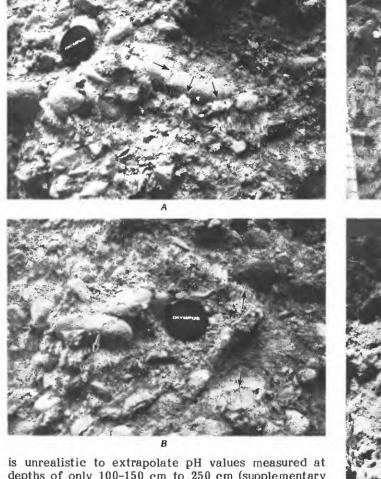
by hand), rubification (reddening and brightening), and pH. These properties can be described, but in this study they were converted to numerical quantities using Harden's (1982b) soil development index (see "Field Methods and Calculations" section). Some other soil properties did not show consistent age trends. Structure tends to be poorly developed in gravelly desert soils, and in the Kane fan soils is disrupted by gypsum. Organic matter content does not increase consistently with time in desert climates, where organic material is rapidly oxidized. The gradual accumulation of CaCO<sub>3</sub> causes soils to whiten with age. Gypsum, rather than CaCO<sub>3</sub>, increases with age on the Kane fans, and indices for color-paling and color-lightening poorly reflect gypsum accumulation.

This study extrapolates or limits all soil depths to 250 cm when calculating property values using the Harden (1982b) index, in order to minimize the variability caused by different sampling depths in the different excavations. However, total-texture (texture plus wet consistence), pH, and rubification vary with depth (fig. 7). Because these three properties reflect weathering and additions to the soil, they should be most pronounced near the surface, where weathering is most intense. Thus, limiting the calculations to 250 cm should not seriously affect the value of a given property summed for a soil profile. Index values for total-texture are indeed largest in the B horizons, but those for pH increase and rubification are often greatest at depth. Hence, index values for these properties summed over the profile may be minimum values for soil development, especially in the older soils.

Total-texture consistently increases with age for soils on fans 1-4, and is significant at the 0.01 level in distinguishing between these soils (fig. 8). The texture-age correlation is also significant (at the 0.025 level) for all soils, but the r² value (coefficient of determination) for this correlation is quite low. Laboratory textural data for these soils cannot be used because in most cases, carbonate parent material was removed along with the gypsum before particle size was determined. Hence texture was determined in the field by hand. Expectably, such a technique reflects the high gypsum content of older soils, for gypsum crystals impart a sandy feel to the soil.

Rubification (both moist and dry colors used) increases consistently with age (fig. 9); the  $\underline{r}^2$  value is significant at the 0.01 level. If extrapolated values are not used, the correlation is improved from  $\underline{r}^2$ =0.69 to  $\underline{r}^2$ =0.86. This trend reflects reddening at depth in the profiles, rather than mainly in the B horizons (fig. 7). The most reddened horizons occur at or just below gypsic horizons; this suggests a connection between rubification and gypsum content. This association also appears in gypsic soils developed on similar parent materials in southern Israel (Dan and others, 1981, p. 268-296; Dan and others, 1982).

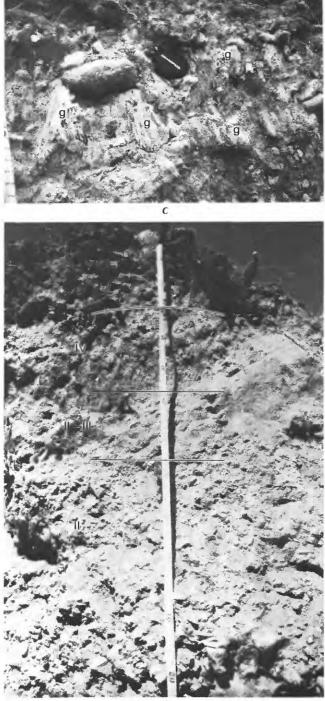
The patterns of pH change are complex (figs. 7 and 10). Values of pH become more alkaline with depth. When pH index values are extrapolated to depths of 250 cm, anomalously high values of pH change for the young soils are derived (fig. 10). The young soils, especially B-21, B-5, and B-15, were sampled shallowly because they showed no visible evidence of pedogenic alteration at shallow depths. It



depths of only 100-150 cm to 250 cm (supplementary table 1).

The minimum index values for pH (calculated to the sampling depth for each soil) increase with time for fans 1-4 but decrease in the older fans (fig. 10). This pattern is consistent with probable changes in soil chemistry as interpreted from soil morphology and soluble salts. The amount of secondary CaCO3 increases with time in the young soils (supplementary table 1), which causes an increase in pH. accumulation of secondary gypsum in older soils (fig. 6) causes a decrease in pH, because the pH of gypsum is 6, as opposed to 8 for CaCO3 (Stevens and Carron, 1948).

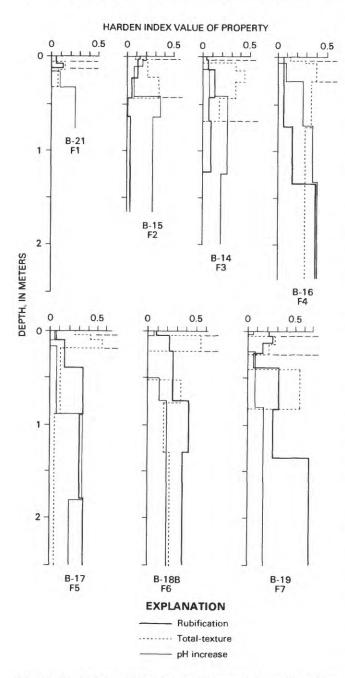
The profile index is calculated in two ways. The first (fig. 11A) combined texture, rubification, and pH increase; the second (fig. 11B) used texture and rubification, in order to eliminate the uncertainties in the calculation of index values for pH. Using three properties, correlation of the profile index with age is not significant when the extrapolated maximum values of the index for young soils are used. If the minimum values are used, the correlation is significant at the 0.01 level. The  $\mathbf{r}^2$  value of the regression increases from 0.57 to 0.86 by deleting the soils on fan 7 that appear eroded. When pH is excluded, the correlation of the profile index with time is significant at the 0.01 level, and the  $\underline{r}^2$  value again increases from 0.69 to 0.86 if the soils on fan 7 are deleted. Comparing these methods, correlation using all the soils is better if pH



Typical appearance in soil profiles of gypsum stages (see "Field Methods and Calculations" section). Gypsum pebble coats and small masses shown by arrows, larger masses indicated by "g." A, Stage I, from the 2C4cs horizon in soil B-17; note gypsumcracked clast in center. B, Stage II, from the 2C3cs horizon in soil B-17. C, Stage IV, from the 2C1csca horizon in soil B-17; note clasts floating in gypsum matrix. D, Soil B-18B, with stage IV gypsum from 20 to 50 cm, stage II-III gypsum from 50 to 75 cm, and stage II gypsum below 75 cm.

values are excluded, but there is no difference between the results when soils on fan 7 are excluded.

Linear rates of soil development (figs. 8-11) are contrary to results obtained by workers in warmer, more moist climates (Colman and Pierce, 1981; Harden, 1982a, b; Birkeland, 1984, p. 204 and 225), where rates were found to be logarithmic or exponential in character. Bockheim (1980b) obtained logarithmic trends with time for soils formed in climates ranging from tropical rainy to cold desert; his



**Figure 7.** Values of pH, total-texture, and rubification in selected soils (B-numbers) on each fan surface. Note increase of rubification and pH at depth in soils. F1-F7, fan numbers. Horizontal dashes separate A, B, and C horizons.

study, however, used maximum values in each profile, rather than summations over the total profile depth. Muhs (1982) expressed soil properties in an arid xeric climate as profile summations and found linear trends with time; these soils probably received considerable

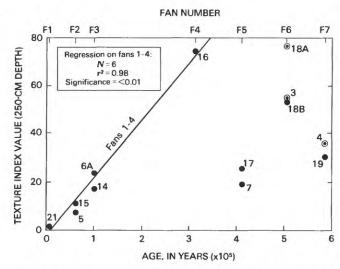


Figure 8. Regression of total-texture (texture + wet consistence) with soil age. Dots are index values for each soil; circled dots are values derived by large extrapolations to 250-cm depth. Numbers correspond to profile numbers without prefix "B-". The r<sup>2</sup> value is for linear regression based on index values for fans 1-4. Index for older soils is nonsystematic in part because of effect of gypsum on field determination of texture and consistence (see text).

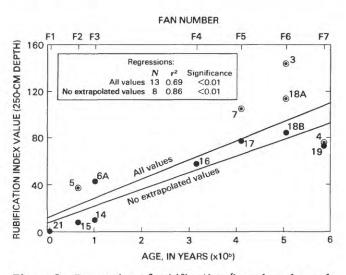


Figure 9. Regression of rubification (based on dry and moist colors) with soil age. Dots are index values for each soil; circled dots represent values derived by large extrapolations to  $250\,\text{-cm}$  depth. Numbers correspond to profile numbers without prefix "B-". The  $\underline{r}^2$  values are for linear regression based on all index values, and on values excluding those extrapolated. Values for older soils are minimum because calculations are limited to  $250\,\text{-cm}$  depth.

dust influx. Therefore, linear trends with soil age of index values based on field criteria for the Kane fan soils may best be explained, not by weathering phenomena, but by cumulative additions atmospheric sources to the soil profiles, and by secondary effects of these additions.

#### Calcium Carbonate and Gypsum

Early in soil development on the Kane fans, CaCO3 is dissolved from the calcareous parent material and from eolian additions, and is reprecipitated lower in the profile mainly as coats and pendants under clasts (supplementary table 1). The depth to and the amount of secondary CaCO3 increase with age in soils on fans 1-3. However, CaCO3 does not increase beyond stage II and, in fact, appears in lesser proportions in the oldest soils, despite the availability of CaCO3 in the parent material and aerosolic dust.

Gypsum first appears in trace amounts in soils on fan 2 and steadily increases in concentration and in thickness with age (fig. 6 and supplementary tables 1 and 3). Petrogypsic horizons first appear in fan 5 soils. Expectably, gypsum accumulates below the first appearance of secondary CaCO3 because gypsum is more soluble. I was not able to find the maximum depth of gypsum accumulation in backhoe pits up to 3 m deep, but observations (table 2) indicate that gypsum amounts decrease with depth and that gypsum does not occur in most of the basal fan gravels.

The composition and amount of dust collected in traps (table 5) suggest that atmospheric additions of silicates, CaCO<sub>3</sub>, and gypsum determine

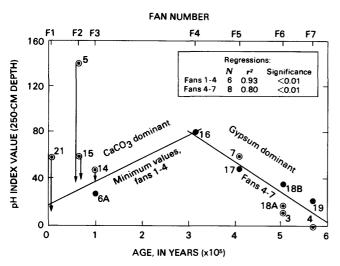


Figure 10. Regression of pH with soil age. Dots are index values for each soil; circled dots are values derived by large extrapolations to 250-cm depth; arrows extend down to minimum index values for young soils. Numbers correspond to profile numbers without prefix "B-". The  ${\bf r}^2$  value for fans 1-4 is for regression using minimum index values for young soils. Values for older soils are minimum because calculations are limited to 250-cm depth.

development of soils on the Kane fans. The annual infall of mineral particles (non-salt, non-organic fraction) is about 5 X 10<sup>-4</sup> g/cm<sup>2</sup> (based on one year of data)<sub>2</sub> CaCO<sub>3</sub> is added at the annual rate of 2 X 10<sup>-4</sup> g/cm<sup>2</sup>, whereas the gypsum rate is about 0.4 X 10<sup>-4</sup> g/cm<sup>2</sup>.

Eolian silt and clay may compose a large part of the silt and clay in the upper parts of the soils. Continual input of silt and clay through time can account for the linear textural increase seen in the younger soils (fig. 8). Gravelly material makes an effective long-term sediment trap, especially where a desert pavement helps protect the surface from deflation. Eolian dust influx also helps to account for

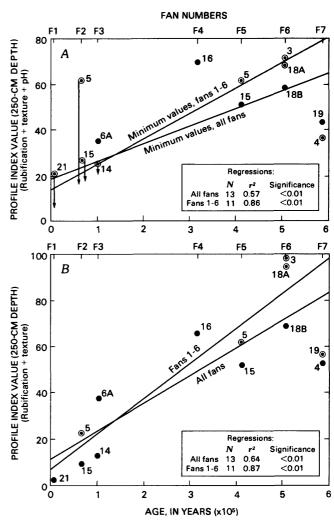


Figure 11. Regression of profile index with soil age. (A, Combined pH, total-texture, and rubification. B, Total-texture and rubification). Dots are index values for each soil; circled dots represent values derived by large extrapolations to 250-cm depth; arrows extend down to minimum index values for young soils. Numbers correspond to profile numbers without prefix "B-". The  $\underline{r}^2$  values in  $\underline{A}$  are for linear regression using minimum values for young soils. Values for older soils are minimum because calculations are limited to 250cm depth.

Table 5. Data from dust traps

[Analytical methods are described in text. The first major oxides column gives percent as measured on the nonsalt silt-plus-clay fraction. The second column gives recalculated percent oxides assuming that all CaCO<sub>3</sub> and gypsum in the traps were contained in the silt-plus-clay fraction. Oxides were determined on combined samples from D-1 and D-2]

	Trap D-1	Trap D-2
	Air-dry weight 0.6818 g	0.9360 g
Percent	of total sample:	
	Organics 19.3	15.6
	CaCO <sub>3</sub> 16.6	22.9
	Gypsum 3.3	5 • 6
	Mineral matter 60.8	55.9
Percent	of mineral matter fraction:	
	Sand 4.0	12.0
	Silt 59.6	54.2
	Clay 36.3	33.8
Annual	flux rates $(10^{-4} \text{ g/cm}^2)$ :	
	Mineral matter 4.8	6.0
	CaCO <sub>3</sub> 1.3	2.5
	Gypsum26	.60

Percent major oxides of silt-plus-clay fraction, combined samples:

		With CaC
	Nonsalt	and gyps
SiO <sub>2</sub>	67.	59.
A1,0,	11.7	10.3
Fe 203	4.4	3.9
Mg 0		1.8
Ca 0	2.0	14.0
Na 20	3.2	2 • 8
к, б		2.6
Tio,		.46
Mn 0		.033
ZrO2	026	.023

the presence of illuvial clay in B horizons of soils that show few signs of weathering (Holliday, 1982).

The CaCO3 dust rate is similar to the Holocene rate calculated by McFadden and Tinsley (1985) for soils in southern California and is slightly higher than modern dustfall rates in the area of Las Cruces, N. Machette (1985) Mex. (Gile and Grossman, 1979). calculated that annual Pleistocene rates of carbonate accumulation were 1.4, 2.2-2.6, and 5.1 X 10<sup>-4</sup> g/cm<sup>2</sup> for the respective areas of Beaver, Utah, Albuquerque-Las Cruces, N. Mex., and Roswell-Carlsbad, N. Mex. Formation of calcic soils in the United States is commonly attributed to atmospheric additions of CaCO3, including calcareous dust and calcium ions contained in precipitation (for example, Gardner, 1972; Bachman and Machette, 1977; Gile and Grossman, 1979).

If  $CaCO_3$  is being added to the Kane fan soils at the rate indicated, the amount of secondary  $CaCO_3$  in the older soils should be much greater than the amount actually observed. For example, soils at Las Cruces as young as 25-75 ka have stage III  $CaCO_3$  horizons (Gile and others, 1981). At Silesia, Mont., in a climate

moister than the Kane area with a smaller carbonate dust flux, 400-ka soils in granitic alluvium have stage III CaCO<sub>3</sub> (Reheis, 1984). Some mechanism that diluted or prevented eolian CaCO<sub>3</sub> from being added to the older Kane fan soils appears to be required. One possible answer is that CaCO<sub>3</sub>, much less soluble than gypsum, could have been removed from the fan surfaces by wind before it could be dissolved and moved into the soils.

Potential sources of gypsum in the study area include: (1) local evaporite fragments in the fan detritus, (2) ground water from nearby springs, (3) the past and present base-flow component of runoff within the fan gravels, and (4) upwind dust sources.

Incorporation of gypsum into the fan debris from local bedrock sources is not significant. Analyses of relatively unaltered fan debris from the base of profiles developed on fans 1-3 and from the base of fan 3 deposits (supplementary table 4), plus field observations at depth (table 2), show only trace amounts of gypsum (about 0.07 percent). The bulk density of unweathered fan gravels is about 2.0 g/cm $^3$ (supplementary table 2). Hence, only 0.0014 g of gypsum can be obtained from a cubic centimeter of parent material. The profile weight of gypsum in a 250-cm thickness of soil on fan 7 is about 43 g/cm<sup>2</sup> (fig. 12). In order to derive all the gypsum in the uppermost 250 cm of fan 7 soils from weathering and solution of the carbonate parent material, about 305 m of deposits above the present land surface would be required. This is extremely unlikely because the fan surfaces are quite well preserved.

A ground water source for the secondary gypsum cannot be dismissed, but several facts argue against (1) Analyses (supplementary table 3) and (table show that gypsum is observations concentrated about 0.25-1.0 m below fan surfaces. It decreases with depth and does not visibly occur in basal gravel in most cases, even where the gravel overlies relatively impermeable shale bedrock. Ground water flow would concentrate above such aquacludes, so if the gypsum precipitated out of ground water, it should also concentrate near the base of the gravel. (2) If ground water were the main contributor of gypsum, the observed increase of gypsum with time should not occur. Gypsum would be added only as long as the soil remained within reach of the ground water table. After incision and isolation of the fan surface, gypsum should not continue to accumulate. (3) The vertical distribution of gypsum and carbonate argues against ground water origin. Precipitation of salts from solutions of upward-moving capillary water should result in more soluble salts occurring closer to the surface than less soluble ones, but gypsum-rich horizons in the Kane soils always occur beneath horizons that contain secondary CaCO3. Additionally, with a ground water source, gypsum should occur closer to the surface in younger soils that are near the ground water table; the reverse is true in the study area.

Eolian-derived gypsum (Page, 1972; Watson, 1979) appears to account for most of the observed increase in gypsum content. Profile weights (g/cm $^2$  in a column of soil) of gypsum (calculated to a depth of 250 cm) increase at a linear rate (fig. 12), and the  $\underline{r}^2$  value based on nine soils is significant at the 0.01 level

when ages are estimated from the stream-incision rate (table 4 and fig. 5). The  $\underline{r}^2$  value increases from 0.90 to 0.93 when the soil on  $\overline{fan}$  7 that appears eroded is excluded from the regression. Annual rates of gypsum accumulation calculated from the profile weights and estimated soil ages vary between 1.1 and 120 X  $10^{-6}$  g/cm<sup>2</sup>. Addition of eolian gypsum (table 5) now occurs at annual rates between 26 and 60 X  $10^{-6}$  g/cm<sup>2</sup>. Thus, the observed accumulation rate of aerosolic gypsum is enough to account for most of the pedogenic gypsum.

Local and long-range sources of eolian gypsum are abundant. Presently stabilized clay and gypsum dunes (up to 90 percent gypsum) 4 km south of the Kane fans indicate eolian gypsum transport (Birdseye, 1983), and may be a gypsum source for the Kane fan soils. Cretaceous rocks underlying the central part of the Bighorn basin commonly contain gypsum. Gypsum deposits are exposed in the Sheep Mountain anticline to the southwest, on the slopes of the Pryor Mountains to the northwest (Pierce, 1978), and in the Cody area on the west side of the Bighorn Basin (Bullock and Wilson, 1969). At present, prevailing winds in the study area are westerly.

#### Effects of Gypsum on Other Soil Properties

#### Depression of CaCO3 solubility

Mobility of CaCO<sub>3</sub> in gypsic soils is restricted because of the common ion effect (Krauskopf, 1967, p.17): the solubility of any salt (CaCO<sub>3</sub>) in equilibrium with a saturated solution is less if an excess of one of its ions (Ca<sup>+2</sup>) is present. Calculations indicate that the solubility of CaCO3 is depressed by a factor of 64 in the presence of a solution saturated with gypsum at 25°C, assuming that equilibrium conditions are maintained and that the partial pressure of CO2 is constant (see "Field Methods and Calculations" section). The solubility of  ${\rm CaCO_3}$  may be even further reduced by the low partial pressure of CO2 in arid soils (modeled growing-season  $pCO_2$  is less than 0.1 percent in arid climates, 0.6-4 percent in semiarid climates; Brook and others, 1983). The reverse is not true; because CaCO3 is much less soluble than gypsum, gypsum solubility remains nearly constant in a CaCO3saturated solution.

Although trace amounts of gypsum occur in the uppermost soil horizons, gypsum accumulates in horizons lower than the first occurrence of secondary CaCO<sub>3</sub> (supplementary table 1). The depths to accumulations of secondary CaCO<sub>3</sub> and gypsum in the soils can be compared best by calculating these depths in units of yearly water penetration (Arkley, 1963) based on the estimated available water-holding capacity (AWC) of each horizon (fig. 13). The depths then are plotted on a curve that represents the movement of water calculated for 24 years of Loyell

weather data (National Climatic Center, 1976). Theoretically, rainfall will dissolve gypsum and  $CaCO_3$  either from eolian dust or parent material and carry the ions downward. As soon as this solution contacts a significant amount of gypsum, the addition of  $Ca^{+2}$  and  $SO_4^{-2}$  ions to the solution will depress the solubility of  $CaCO_3$  and it will precipitate. Thus, the more gypsum that accumulates, the less movement of  $CaCO_2$  will occur.

 ${\rm CaCO_3}$  will occur. The following method permits estimation of the depth at which gypsum saturation occurs and CaCO3 precipitates for the Kane fan soils (table 6). The amount of water per year that is available to dissolve the gypsum in a 1-cm<sup>2</sup> column through a soil horizon is equal to the yearly amount of water passing through the soil at the base of that horizon (fig. 13), plus the water held within the horizon (AWC, table 6). Finely powdered gypsum is soluble at the rate of 19 X 10 g/mL H<sub>2</sub>O (distilled) at 10°C (Hardie, 1967). The potential amount of gypsum that can be dissolved yearly in a horizon is equal to the gypsum solubility multiplied by the total amount of water held in and passing through that horizon. The amount of gypsum present in a 1-cm<sup>2</sup> horizon column is known (table 6). The percent saturation with gypsum of water percolating down from the base of the soil horizon is equal to the amount of gypsum actually present divided by the amount that could be dissolved, assuming that if sufficient unsaturated water is present, all the gypsum in the horizon will dissolve. The calculation is repeated for each successively lower horizon, except that the capacity of water to dissolve gypsum in each lower horizon is decreased by the percent saturation of the water when it enters that horizon.

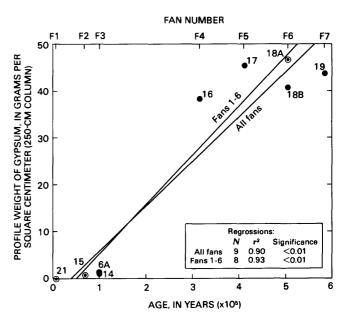


Figure 12. Regression of profile weights of pedogenic gypsum with soil age. Dots are values for each soil; circled dots represent values derived by large extrapolations to 250-cm depth. Numbers correspond to profile numbers without prefix "B-". The r<sup>2</sup> values are for linear regression based on soils for which bulk density is known.

<sup>&</sup>lt;sup>3</sup>The variation in rates of gypsum accumulation calculated from the soils may be due to episodic solution of gypsum, or to variation in the gypsum dust flux with time.

The gypsum-saturation depth of the normal yearly amount of water occurs within 10 cm of the surface in all the Kane fan soils except B-21 on fan 1 (table 6). In most of the soils, the water becomes saturated with gypsum at or just above the uppermost level of accumulation of CaCO<sub>3</sub> (B-14 and B-19, table 6). This observation supports the above hypothesis concerning the inhibition of CaCO<sub>3</sub> dissolution by gypsum. In profile B-19, the water percolating through the soil is not saturated with gypsum until it enters the uppermost gypsic horizon, implying ongoing dissolution of gypsum. Field observations of this soil indicate that dissolution is occurring, for the gypsum distribution in

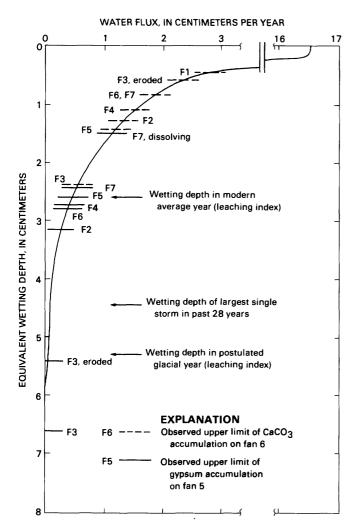


Figure 13. Average yearly water movement (24-year record at Lovell, Wyo.) and relationship to gypsic and calcic horizons using Arkley's (1963) method. Upper limits of gypsum (solid line) and CaCO<sub>3</sub> (dashed line) horizons for soils on fans F1-F7; units of depth were calculated by estimating AWC for each horizon from textural data. F1 has no gypsum accumulation in the depth sampled. "F3. eroded" denotes top of carbonate horizon of a stripped soil (there are two F3 soils); "F7. dissolving" is top of gypsic horizon that shows signs of ongoing dissolution. Soil profiles used include: F1 = B-21, F2 = B-14 and B-15, F3 = B-6A, F4 = B-16, F5 = B-17, F6 = B-18A, F-7 = B-19.

the B horizon is patchy and irregular. This is shown by two different lines representing the depths to gypsum accumulation for the soil on fan 7 in figure 13.

Several assumptions are involved in calculations of gypsum-saturation depths: (1) All of the gypsum in each horizon is dissolved if sufficient unsaturated water is present. This assumption is probably justified for the A and B horizons of the Kane fan soils, because in these horizons, gypsum occurs in such small, disseminated amounts that it is not visible. Gypsum in these horizons is probably finely crystalline and can dissolve rapidly. The assumption is not justified for the horizons with coarsely crystalline gypsum (all those horizons designated "cs" in table 6). (2) The solubility of gypsum is not depressed by a common ion present in a more soluble salt. No other salts were observed in X-ray diffraction analyses of the Kane fan soils. (3) The amount of gypsum calculated in each horizon is reasonably accurate. Horizon weights of gypsum are calculated from the percent gypsum, bulk density, and horizon thickness (supplementary table 7). The total horizon weight includes both pedogenic gypsum and parent-material gypsum, assuming the latter has not been leached Due to uncertainties, only the amount of pedogenic gypsum was used to calculate saturation depth of gypsum. Hence, saturation depths in table 6 are probably minimum estimates.

An obvious implication of this work is that, through time, the amount of gypsum in the A and B horizons should decrease and disappear as it is dissolved. That this does not occur is proved by the occurrence of gypsum in small amounts in the upper parts of all soils. Several factors may explain this phenomenon: (1) Small amounts of gypsum are added by dust influx (table 5). This factor is considered unlikely, for present rainfall is more than enough to dissolve the very small amounts of gypsum added (26-60 X  $10^{-6}$  g/cm<sup>2</sup> per year). (2) Small amounts of gypsum, if present, may be released from the limestone and dolomite clasts as they weather on and near the ground surface. Most carbonate rocks contain trace amounts of gypsum and other salts (less than 0.1 percent; Lamar and Shrode, 1958). These two factors are incompatible: if percolating water dissolves aerosolic gypsum near the ground surface, then the common ion effect will hinder this water in dissolving carbonate and releasing gypsum from the near-surface clasts. (3) Evaporation of water at the ground surface probably causes upward movement of soil water in the upper, fine-grained parts of the soil (supplementary table 1). Thus, small amounts of gypsum that are dissolved and moved down by a rainfall event may be redistributed upward as evaporation occurs.

#### Effects on other soil-forming processes

Pedogenic silt and clay shows no increase in soils older than 300 ka (fig. 8); this may be related to the gypsum accumulations in older soils. Gypsum may control the movement of fines in two different ways: First, gypsum is much more soluble than CaCO<sub>3</sub>; thus it flocculates clay particles more effectively than CaCO<sub>3</sub> and can prevent clay from being translocated into the profile. Second, if enough gypsum

Table 6. Estimated depths to saturation of leaching water with gypsum

[Numbers (C1-C6) at top of each column are used to abbreviate equations to calculate the quantities in other columns. Equations are shown in parentheses at the bottoms of column headings. GS, gypsum solubility,  $19x10^{-4}$ g/mL at  $10^{\circ}$ C (Hardie, 1967). C6, value of column C6 for the next higher horizon]

Fan and profile number,	Horizon thick- ness	Est. AWC (cm)	C2 Estimated water flux from horizon (mL/yr) (from fig. 13)	C3 Water in or leaving horizon (mL/yr (C1+C2)	C4 Gypsum capacity of water (g) $(C3 \cdot CS \cdot [\frac{100 - C6}{100}])$	C5 Pedogenic gypsum in horizon (g/cm <sup>2</sup> ) (supp. table 7)	C6 Saturation of water leaving horizon (percent) (C5/C4)
F1 (B-21)							
A	6	0.47	2.9	3.4	0.0065	0	0
2Btca	7	.18	2.3	2.5	.0048	.0002	4
2Cca	20	• 52	1.5	2.0	.0036	.0012	33
2Cn	40	•44	1.2	1.6	.0020	.0008	40
F3 (B-14)							
All	. 4	•47	3.0	3.5	.0067	0	0
A12	4	.47	1.7	2.2	.0042	.0007	17
B1	9	1.46	•6	2.1	.0033	.0049	100
B2ltca	11	1.78	•2	2.0	0,	0	100
B22tca	16	2.45	0	•2	0	0	100
2B3tca	26	• 57	0	0	0	.0037	100
2Cca	53	1.17	0	Ö	0	.0935	100
2Cnca	<b>7</b> 0	1.40	0	0	0	.218	100
F7 (B-19)							
A	6	.86	1.8	2.7	.0051	.0019	37
B21tca	8	. 58	1.3	1.9	.0023	.0007	31
B22tcsca	11	.99	•6	1.6	.0021	.509	100
2Clcsca	15	1.56	•2	1.8	0	7.13	100
2C2cs	45	3.38	0	•2	0	13.9	100
203cs	52	4.06	0	0	0	13.0	100
3C4cs	125	3.25	0	0	0	9.47	100

accumulates, the soil horizon may be plugged, and clay particles physically cannot pass through it. Petrogypsic horizons in soils on fans 5, 6, and 7 might act as barriers to clay movement.

Gypsum appears to cause rapid physical disintegration of gravels (fig. 6). Limestone and dolomite cobbles within gypsic and petrogypsic horizons are commonly split, and gypsum crystals fill the cracks. Gypsum crystallization has apparently split the clasts. Gypsum crystals also have disaggregated andesite clasts along grain boundaries. Such weathering has been reported in other arid regions (Coleman and others, 1966; Dan and others, 1982) and has been duplicated in the laboratory (Goudie and others, 1979). It is commonly attributed to pressures caused by hydration and crystallization.

Rubification may be associated with gypsum in the Kane fan soils. Rubification does not appear to be related to weathering processes in the B horizons of these soils, because it continues to increase with depth in the profile (fig. 7). However, Dan and others (1982) suggested that high salt contents of older desert soils in southern Israel may enhance deep weathering, because the hygroscopic nature of the salts encourages moisture retention in the soils. In addition, gypsum may encourage rubification by disintegrating clasts and mineral grains, which increases the surface area of debris exposed to oxidation. Petrogypsic horizons do not appear most rubified simply because gypsum dilutes the red coloration.

#### Possible Climatic Implications

Depth to accumulations of secondary CaCO<sub>3</sub> and gypsum should bear some relation to climate (Jenny

and Leonard, 1939; Arkley, 1963). These relations are examined by plotting the depths to secondary  $\text{CaCO}_3$  and gypsum in units of yearly water penetration based on the AWC of each horizon (fig. 13).

The wetting depths to accumulations of secondary  $CaCO_3$  in the Kane fan soils do not show a relation with soil age (fig. 13). As discussed above, the wetting depth to secondary  $CaCO_3$  is probably controlled in these soils by the amount of secondary gypsum contained in the uppermost horizons. The measured depths to secondary  $CaCO_3$  in the profiles (supplementary table 1) are only a few centimeters. Thus, the variable wetting depths to  $CaCO_3$  in soils of various ages (fig. 13) may be related to slight erosion or deflation of fan surfaces.

Different depths to gypsum horizons in younger (F3) and older soils (F4-7, fig. 13) suggest a drier climate before 100 ka. Because temperature has little effect on gypsum solubility (Hardie, 1967), effects on the soils would probably be related to changes in effective moisture caused by changes in temperature or in the amount and seasonal distribution of precipitation. Gypsum horizons of fan 3 soils occur near the leaching depth calculated from the postulated climatic parameters (table 1) during glacial periods, but positions for soils on fans 2 and 4-7 cluster closely around the modern leaching index of 2.6 cm.

The soils on fans 2 and 3 have been subjected to the climates of the last glaciation and of the Holocene. Accumulation of silt and clay in the A and B horizons with time raises the soil AWC. If the accumulation of gypsum at a given wetting depth occurred prior to the accumulation of silt and clay above the gypsum, then the wetting depth to the gypsum measured at present will be greater than that

at which it accumulated. Therefore, greater depth to gypsum in fan 3 as opposed to fan 2 soils may be agerelated rather than climate-related. Alternatively, the assumed age of 65,000 yr for fan 2 may be incorrect. If fan 2 is younger, perhaps dating to the end of the last glaciation, then it has developed only under the recent interglacial climate. In that case, the depth to gypsum in fan 2 could reflect the modern, shallow leaching depth, whereas gypsum depth in fan 3 soils mainly reflects the greater available moisture of the last glaciation.

Soils on fans 4 through 7 have developed through many glacial-interglacial cycles. The presence of secondary gypsum at depths well below the maximum concentration of gypsum demonstrates that greater available moisture during glacial episodes probably affected the soils. But the maximum gypsum horizons correspond to the modern interglacial wetting depth; this suggests that the positions of these horizons is controlled by the interglacial episodes.

The simplest hypothesis that explains gypsum depth in soils older than 100 ka invokes the hygroscopic qualities of gypsum. It is well known that salts have the ability to sorb larger amounts of moisture than mineral grains (Winkler, 1975; Dan and others, 1982). Experiments with brine-saturated bricks have shown that the volume percent of hygroscopic moisture content of a brick increases by a factor of 20 when the amount of salt (NaCl) in the brine increases from 0 to 4 percent (Winkler, 1975). Because the calculated AWCs are based on field texture, they probably underestimate the real AWCs of gypsum-rich soils. If gypsum has even 10 percent the water-holding capacity of NaCl, then a gypsic horizon with 5 percent gypsum can double the AWC of that horizon. Gypsum quantities in Kane fan soils approach 50 percent in some horizons, including gravel content. Estimated from field textures, the AWC of the uppermost gypsic horizons in fans 5, 6, and 7 is 2 cm in a 20-cm-thick, 1cm<sup>2</sup> section. The gypsum in these horizons is probably sufficient to hold all of the estimated additional moisture of a glacial climate. The water-holding ability of the gypsum is further enhanced by the high porosity of the gypsic horizons: the bulk density of gypsic peds in some of these horizons is less than 1,0  $g/cm^3$ , whereas the density of gypsum is 2.3  $g/cm^3$ . When gypsum accumulates in a near-surface position over a lengthy interglacial, it can increase the soil AWC enough to hold the increased moisture of a glacial period within the gypsic and overlying A and B horizons. Thus, the gypsic horizon may prevent its own leaching.

Two other hypotheses could help account for shallow depth to gypsum in older soils. First, the rate of gypsum influx during glacial periods may have increased to levels that overwhelmed the ability of the downward flux of water to transport gypsum. Because loess deposition accompanies and closely follows glaciations, this hypothesis has some merit. But even with modern rainfall, the influx must increase by two orders of magnitude to allow gypsum to remain at shallow depths in the old soils. This hypothesis alone seems inadequate. Second, moisture during glacial periods may have actually decreased from modern levels, even though evapotranspiration was greatly reduced, because of the predominant northerly winds

and the increased proportion of snow to rain. If no winter moisture was added to the soils due to the presence of frozen ground and wind deflation, summer precipitation must have decreased to half the present level to place the calculated glacial leaching index at the level of the gypsum horizons. Such a decrease is much larger than that previously postulated (Gates, 1976) but is within the realm of possibility.

In summary, the most likely hypothesis that accounts for the retention of gypsum near the surface of soils that have undergone glacial climates is the hygroscopic property of gypsum. This hypothesis is most simple because it does not require any changes external to the soils. It is also possible that increased dust flux and decreased moisture during glaciations may have played a part in the development of the Kane fan soils. In any case, because gypsum increases linearly with time, precipitation in the past 600,000 years has never been sufficient to flush gypsum completely from the soils.

#### SUMMARY

Field properties and gypsum contents of Kane fan soils change at statistically significant linear rates. These linear rates imply an external source for pedogenic materials. Total texture and pH increase with time in soils on fans 1-4, whereas pH appears to decrease in soils on fans 5-7. Rubification increases with time in all soils. The profile index increases with time, but the correlation is stronger if the index is calculated using only texture and rubification.

Evidence from dust traps indicates that the modern influx of aerosolic gypsum is sufficient to produce a large proportion of the observed amounts of pedogenic gypsum. Despite the presence of alkali springs and gypsum in bedrock units in the Cottonwood Creek drainage, gypsum occurs in very small quantities (less than 0.1 percent) in fan gravels at depth, and it does not concentrate in basal gravels over impermeable shale where ground water should concentrate.

Gradual accumulation of pedogenic gypsum strongly affects development of other soil properties. Gypsum causes rapid clast weathering and may cause reddening of the C horizons; it also reduces the solubility of CaCO<sub>3</sub> so much that redistribution of CaCO<sub>3</sub> has stopped in soils older than 300 ka. The small amounts of gypsum in the A and B horizons of most of the soils is apparently sufficient to prevent CaCO<sub>3</sub> from being carried in solution to depths greater than about 10 cm by the average yearly precipitation. Gypsum may also inhibit downward translocation of clay by flocculating clay particles, which would account for the lack of continuing increase in fines near the surface of older soils.

Large accumulations of gypsum can so increase the AWC of the soil that increased moisture during glacial episodes can be held at shallow depths. The positions of gypsic horizons should be related to available soil moisture, which was probably increased during past glacial periods. Gypsic horizons in soils older than 100 ka correspond to the modern average yearly wetting depth, which suggests that over a long period of time, depth to gypsum is controlled by soil

moisture available during interglacials. Gypsum that accumulates close to the surface in an interglacial climate can, by increasing soil AWC, hold the increased moisture available during a glacial period. Some younger soils have accumulations of gypsum deeper than older soils; this may be because younger soils reflect the climate of the last glaciation, when leaching depths were greater, or because the small amount of gypsum added to young soils can be easily moved to greater depths by large wetting events.

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#### FIELD METHODS AND CALCULATIONS

#### Sampling and Description

Soils developed on the Kane fans were sampled in hand-dug holes or backhoe pits at sites that appeared least affected by erosion or deposition subsequent to abandonment of each fan surface (fig. 2 and sites were supplementary table 1). These supplemented by exposures in an ash pit (B-7) and a diversion ditch (B-6A). I sampled at least two soils on each fan, except fans 1 and 4, to assess soil variability. Four soils appear to have undergone slight truncation of the upper parts of the B horizons (supplementary tables 1 and 2). Humans probably truncated B-5, located near a parking area, and B-6A, on the sloping side of a diversion ditch. The surface of fan 7 is more bouldery and dissected than the others; thus, soils B-4 and B-19 may have been eroded. Alkali springs are just east of remnants of fans 4 and 5. In order to minimize possible groundwater effects, soils B-16, B-17, and B-7 were sampled as far away from the springs as possible. The surfaces of these fans at their westward ends are presently well above any potential groundwater influence, but earlier alteration could have occurred when the fans were less dissected.

Horizon nomenclature and soil property descriptions follow Soil Survey Staff (1975) usage, with the exception of texture modifiers describing gravel content. Field textural descriptions (determined by hand) in supplementary table 2 were not corrected using laboratory-determined carbonate-free particlesize data, because removal of carbonates prior to laboratory work resulted in removal of the limestone and dolomite parent material. Hand-determined field textures are influenced not only by carbonates, but also by the feel of gypsum crystals in gypsum-rich horizons. Moist and dry colors were obtained with a Munsell soil color chart. pH was determined in the laboratory with a pH meter. CaCO<sub>3</sub> stages follow those proposed by Bachman and Machette (1977). I adapted gypsum stages (fig. 6) from CaCO3 stage descriptions as follows:

1-thin, discontinuous coatings on lower surfaces of stones

Il-abundant gypsum pendants under stones;

gypsum crystals scattered through matrix or forming small, soft powdery nodules

III--gypsum continuous through matrix and forming large (up to 10 cm long) pendants under stones

IV--continuous gypsum-plugged fabric; stones and smaller debris float in gypsum matrix

#### Index of Soil Development

Soil field data, which are measured mainly on ordinal scales, were converted to ratio data using Harden's (1982a) index of soil development, in which a point system is used to quantify various field properties of soils. Points for each horizon are then normalized to give all properties equal weight and multiplied by the horizon thickness; the products are summed for each soil. Thus, individual properties can be examined with depth in one soil or compared between soils, or several properties can be combined to obtain a total profile-development index value for each soil. (See flow diagram in Harden and Taylor, 1983).

I have modified Harden's index for the Kane fans. The change of pH in the index is assessed in terms of pH lowering, and this works well in relatively moist environments. In the case of soils developed on the Kane fans, however, I believe that pH increases with soil age, because of the aridity and sparse vegetation of the area and because of the availability of CaCO<sub>3</sub>. This phenomenon occurs in semiarid soils along Rock Creek, Mont. (Reheis, 1984). The change in pH for the Kane fans, therefore, was shown as a pH increase from the parent material state. A value of 1.5 was used as the normalization value for pH increase, because it was the maximum pH increase observed in soils of the northern Bighorn Basin (Reheis, 1984).

It is very difficult to assess parent-material pH for the Kane fan soils. The pH of basal C horizons of the young soils on fans 1-3 is about 8.4. If this value accurately reflects parent-material pH, then nearly all the soil horizons show decreases in pH (supplementary table 3). Such a trend is inconsistent with observations of calcareous soils made elsewhere (Reheis, 1984). In theory, the pH in calcareous soils is unlikely to decrease unless leaching is sufficient to remove a significant proportion of the exchangeable Ca<sup>+2</sup> ions (Birkeland, 1984, p. 22-23). Thus, I believe that the pH of 8.4 representing the lowest depths sampled in young Kane fan soils is a pedogenic product; consequently the original pH is likely less than 8.4. The mean pH of leachate from limestone is about 7.9, whereas the mean pH of leachate from dolomite is about 8.5 (Lamar and Shrode, 1958). As limestone is slightly more soluble than dolomite, weathering of limestone probably dominates soil pH on the Kane fans. These considerations led me to select arbitrarily a value of 8.0 as the parent-material pH in order to calculate pH increase for these soils. Perhaps coincidentally, this pH value is the same as the mean pH of the calcareous A horizons of Kane fan soils.

#### **Dust Traps**

Dust traps were set out in three locations on the fans (fig. 2) to assess aerosolic additions to the soils.

Plastic-lined aluminum trays, 29.5 x 29.5 x 6.5 cm in size, were filled with marbles and mounted on poles about 2 m above the ground (Gile and Grossman, 1979). After one year, two traps were removed (the third was destroyed by high winds), and the dust was collected by washing the traps and marbles with distilled water.

#### **Water Movement**

Arkley's (1963) method of calculating water movement in soils was used to examine the relationship of local climate to the depth of CaCO<sub>3</sub> or gypsum accumulation. Month-by-month precipitation data for the 24-year record at Lovell and average monthly potential evapotranspiration (using van Hylckama's method, 1959) were used to calculate monthly excesses of precipitation over evaporationthat is, the depth of water available to wet the soil in a given month. The average amount of water per year that passes any given depth in a soil can then be calculated. If the available water-holding capacity (AWC) of the various horizons is known or can be estimated, the amount of water can be plotted against actual soil depth. I estimated AWC of the Kane fan soils from field textural data using Salter and Williams' (1967) method. They state that their method has a mean error of ±17 percent using a 12 textural-class (field) system.

#### Depression of CaCO3 Solubility by Gypsum

Solubilities of gypsum (Sg) and CaCO<sub>3</sub> (Sc) at 25°C are calculated from their solubility products (Kg, Ke) as follows (Krauskopf, 1967):

Kg = 
$$(Sg)^2$$
 =  $(Ca^{2+})(SO_4^{2-})$  =  $2.51 \times 10^{-5}$  mo1/L  
Sg =  $(2.51 \times 10^{-5})^{0.5}$  =  $5.00 \times 10^{-3}$  mo1/L  
Kc =  $(Sc)^2$  =  $(Ca^{2+})(CO_3^{2-})$  =  $6.03 \times 10^{-9}$  mo1/L  
Sc =  $(6.03 \times 10^{-9})^{0.5}$  =  $7.76 \times 10^{-5}$  mo1/L

If a solution is saturated with respect to gypsum, the concentration of  ${\rm Ca}^{2^+}$  in solution is 5.0 x  $10^{-3}$  mole/liter. Let <u>x</u> equal the solubility of  ${\rm CaCO}_3$  in this gypsum-saturated solution; then

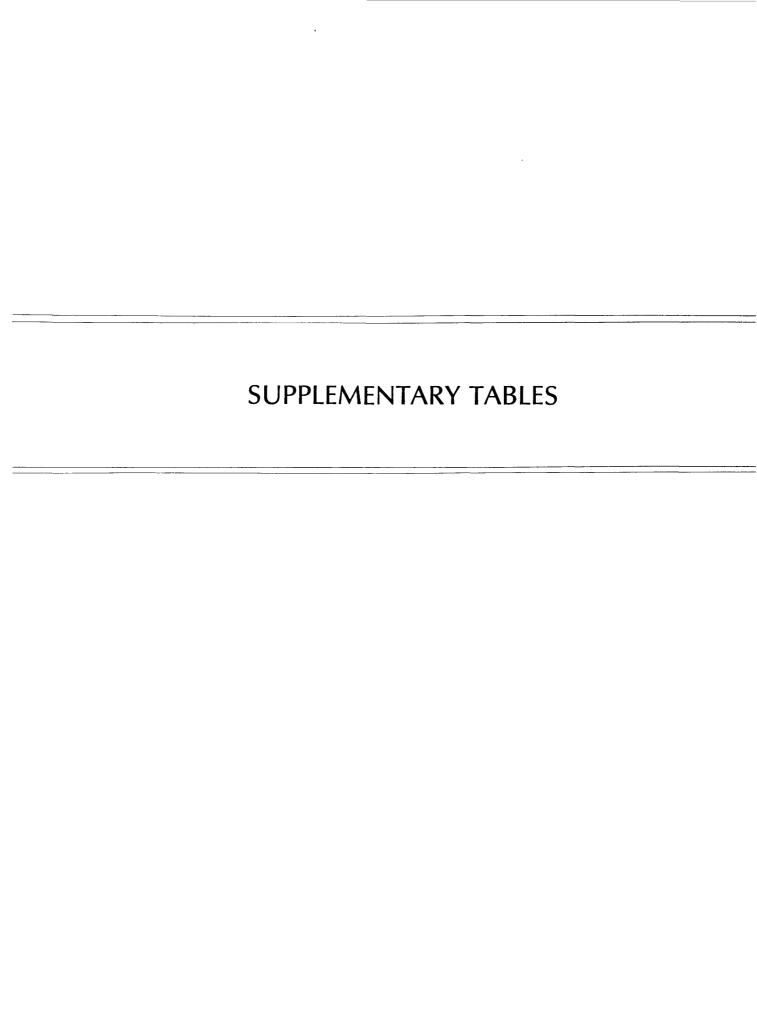
Kc = 
$$(Ca^{2+})(CO_3^{2-}) = (\underline{x}+5.0\times10^{-3})\underline{x}$$
  
=  $x^2+5.0\times10^{-3}x = 6.03\times10^{-9} \text{ mol/L}$ 

 $\underline{x}$  will not be larger than the solubility of CaCO  $_3$  in pure water, so the term  $\underline{x}^2$  is small enough to be ignored. Then

$$5.0 \times 10^{-3} \underline{x} = 6.03 \times 10^{-9} \text{ mol/L}$$
  
 $\underline{x} = 1.21 \times 10^{-6} \text{ mol/L}$ 

These calculations indicate that the solubility of CaCO<sub>3</sub> is depressed by a factor of 64 in the presence of a solution saturated with gypsum at 25°C.

#### C22 Soil Chronosequences in the Western United States



Supplementary Table 1, Part 1. Sample locations and site conditions

Site	Geologic Unit	Elevation (ft)	Modern vegetation <sup>1</sup>	Modern land use	Type of excavation	Parent material texture <sup>2</sup>	Location L (Montana Base Meridian)
B-21	Fan 1	3,735	cactus, sagebrush, grasses	grazing	hand-dug hole	gr/S,Si	SE/4,SE/4,NE/4, Sec. 4, T. 56 N., R. 94 W.
B-5	Fan 2	3,667	cactus, sagebrush, grasses	grazing	hand-dug hole	gr/S,Si	NW/4,SW/4,SE/4, Sec. 4, T. 56 N., R. 94 W.
B-15	Fan 2	3,735	cactus, sagebrush, grasses	grazing	backhoe pit	gr/S,Si	NE/4,NE/4,SE/4, Sec. 4, T. 56 N., R. 94 W.
B-6A	Fan 3	3,715	cactus, sagebrush, grasses	grazing	diversion ditch	gr/S,Si	NE/4,SW4,NE/4, Sec. 4, T. 56 N., R. 94 W.
B-14	Fan 3	3,725	cactus, sagebrush, grasses	grazing	backhoe pit	gr/S,Si	SW/4,SW/4,SW/4, Sec. 3, T. 56 N., R. 94 W.
B-16	Fan 4	3,920	cactus, sagebrush, grasses	grazing	backhoe pit	gr/S,Si	SW/4,NE/4,NE/4, Sec. 3, T. 56 N., R. 94 W.
B-17	Fan 5	3,945	cactus, sagebrush, grasses	grazing	backhoe pit	gr/S,Si	NW/4,NE/4,NE/4, Sec. 3, T. 56 N., R. 94 W.
B-7	Fan 5	3,970	cactus, sagebrush, grasses	grazing	hand-dug hole		NW/4,NW/4,NW/4, Sec. 2, T. 56 N., R. 94 W.
B-3	Fan 6	4,205	cactus, sagebrush, grasses	grazing	hand-dug hole		NE/4,NW/4,SW/4, Sec. 1, T. 56 N. R. 94 W.
B-18A	Fan 6	4,085	cactus, sagebrush, grasses	grazing	backhoe pit	gr,L	NW/4,NE/4,NW/4, Sec. 11, T. 56 N., R. 94 W.
B-18B	Fan 6	4,085	cactus, sagebrush, grasses	grazing	backhoe pit	gr,L	NW/4,NE/4,NW/4, Sec. 11, T. 56 N., R. 94 W.
B-4	Fan 7	4,145	cactus, sagebrush, grasses	grazing	hand-dug hole	gr,L	NE/4,SW/4,NE/4, Sec. 11, T. 56 N., R. 94 W.
B-19	Fan 7	4,180	cactus, sagebrush, grasses	grazing	backhoe pit	gr,L	NE/4,SE/4,NE/4, Sec. 11, T. 56 N., R. 94 W.

 $<sup>^{\</sup>mathrm{l}}\mathrm{Percent}$  bare ground is 50 percent or greater at all sites.

<sup>&</sup>lt;sup>2</sup>Textures: gr/S, gravel and sand; L, loam; Si, silt.

#### Supplementary Table 1, Part 2. Field descriptions

[--, not measured. Analyst: M. C. Reheis, U.S. Geological Survey]

												Assumed Material			
No.	Sample Number	Horizon	Depth (cm)	Lower Boundary	Moist y Color	Dry Color	Texture	Structure	Wet Consis- tence	Clay Films	рН		Wet Con- sistence	Stage CaCO <sub>3</sub>	Stage Gypsum
							Fan 1,	5 ka							
1	B-21	A	0-6	c,w		7.5YR5.5/3.5		lfpl,v	so,po	0	7.7	SL	so,po	0	0
2		28tca 2Cca	6-13 13-33	c,w c,s	7.5YR4.5/4 7.5YR5/3	7.5YR5.5/4 7.5YR6.5/3	gSL- vgSL	lgr m	so,po so,po	0	8.0 8.1	LS LS	so,po so,po	I I+	0 0
4		2Cn	33-73+		7.5YR5/3	7.5YR5.5/3	vgLS	m.	so,po	0	8.4	LS	so,po	0	0
							Fan 2,	65 ka							
5 6	B-5	A B2t no	0-3 3-11	c,W	 7.5YR4.5/6	 7 EVDE E/4	sgL	v 2mabk	so,po	0	 8.6	SL SL	so,po	0 I	0
7		B2tca B3tca	11-20	d a,s	7.5YR5/4	7.5YR5.5/6 7.5YR6.5/5	sgL sgSL+	lfsbk	ss,ps ss,ps	0	8.4	SL	so,po so,po	II	o
8		2Clca	20-31	đ	7.5YR6.5/4	7.5YR7/4	vgSL	m	ss,po	0	8.6	SL	so, po	I	0
9 10		202ca 203ca	31-46 46-71+	d 	7.5YR5/4 7.5YR5/4	7.5YR6/4 7.5YR6/5	vgSL vgLS	m m	ss,po so,po	0	8.9 8.9	SL LS	so,po so,po	I I-	0 0
11	B-15		0-4		7 EVD/ E//	0 25704//		2fa1		0	8.3	SL	00.00	0	0
12	D-13	A B1	4-12	a,s g	7.5YR4.5/4 7.5YR4.5/4	8.25YR6/4 7.5YR4.5/4	gSL sgfSL	2fpl,v lfsbk,lfpl	so,po ss,ps	ő	8.2	SL	so,po so,po	0	Ö
13		B2tca	12-24	c,w	7.5YR4.5/4	7.5YR4.5/4	gfSL	lfsbk	ss,ps	0	8.1	SL	so,po	I	0
14 15		2B3tca 2Cca	25-45 45-65	d d	7.5YR5/4 7.5YR4.5/3	7.5YR6/3 7.5YR5/3	vgSL vgLS	lgr sg,m	ss,ps so,po	0	8.1 8.5	LS LS	so,po so,po	I I– I	0 I-
16		2Cn	65-165+		7.5YR4/3	8.25YR5/2.5	vgLS	m	so,po	Ŏ	8.4	LS	so, po	I-	I-
						:	Fan 3, 1	00 ka							
17	B-6A	A	0-4	a,s	7.5YR4.5/4	7.5YR5.5/3	gL	v	so,po	0	8.1	SL	so, po	0	0
18 19		B2tca 2B3tca	4-29 29-46	a,w a,w	7.5YR4/4 7.5YR5/3	7.5YR5/4 7.5YR6/3	sgL vgSCL	lfabk lgr	ss,ps s,p	0	8.1	SL SL	so,po so,po	I II	0
20		2Cca	46-88	d d	5YR5.5/4	5YR5.5/3.5	vgSL	m	ss,po	Ö	8.2	SL	so,po	I	I-
21		2Clcsca	88-181	ď	5YR5/5	5YR5.5/3.5	vgSL	m	ss,po	0	8.2	SL	so, po	I 0	I I-
22 23		2C2cs base of	181-242 fan 3	2+ d	5YR5/4 10YR6/4	5YR5.5/3 7.5YR7/4	vgLS vgLS	m m	so,po so,po	0	8.1 8.5	LS LS	so,po so,po	0	0
24	B-14	A11	0-4	c,s	7.5YR4/3	7.5YR5.5/3.5		2mgr,1fpl	so,po	0	8.0	SL	so,po	0	0
25 26		A12 B1	4-8 8-17	a,s	7.5YR4.5/4 7.5YR5/4	7.5YR5.5/3 7.5YR5/3	sgfSL	2fpl lfsbk	so,po ss,ps	0	8.0 8.0	SL SL	so,po so,po	0 0	0 0
27		B2ltca	17-28	a,s c,i	7.5YR4.5/4	7.5YR5/4	sgL sgL	lf-mabk, lmpr		ő	8.1	SL	so,po	Ĭ	ő
28		B22tca	28-44	a,s	7.5YR4.5/4	7.5YR5/4	sgL	lf-mabk,1mpr		0	8.1	SL	so,po	I+	0
29 30		2B3tca 2Cca	44-70 70-123	d d	7.5YR5/4 7.5YR4.5/4	7.5YR6/3 7.5YR5/3.5	vgSL- vgLS	sg,lfabk m	ss,po so,po	0	8.4 8.4	LS LS	so,po so,po	II I	0 I-
31		2Cnca	123-193		7.5YR4/3	7.5YR5/3	vgLS	m	so,po	Ö	8.3	LS	so, po	Ī-	ī-
							Fan 4, 3	315 ka							
31	B-16	Al 1	0-4	c,s		7.5YR5.5/3	gSL	v	ss,po	0	8.0	SL	so,po	0	0
33 34		A12 Bt	4-9 9-12	c,s a,s	6.25YR5/4	7.5YR5.5/3	gSL+ gSL+	2gr,v lfsbk,sg	ss,ps ss,p	0	8.0	SL SL	so,po so,po	0 0	0 0
35		Btca	12-28	g,w	7.5YR6/4	7.5YR6/3	gSL+	lfsbk,sg	ss,p	Ö	8.1	SL	so,po	I	0
36		Ccacs	28-75	d	7.5YR7/4	7.5YR7/3	gSCL	m	s,p	0	8.4	SL SL	ss,ps	II I	II
37 38		2Clesea 2C2es	75-137 137-237	d 7+	5YR6/4 3.75YR5/6	5YR6/3 2.5YR5.5/4	vgSCL- vgSCL-		s,p s,p	0	8.5 8.6	SL	ss,ps ss,ps	0	I+
							Fan 5, 4	410 ka							
39	B-7	A	0-6	a,s	7.5YR5/6	7.5YR6/6	gL_	lgr,v	ss,ps	0	7.8	SL	so,po	0_	0
40 41		Clesca 2C2esca	6-30 30-50	d d	7.5YR7/6 5YR6/7	7.5YR8/4 5YR8/6	sgSL gSL+	m m	so,po	0 0	8.0 8.0	LS LS	so,po so,po	II II	II+ III
42		2C3csca	50-83	g	5YR6/6	5YR7/4	gSL-	m	so,po	0	8.3	LS	so,po	I	ΙΙ
43		2C4cs	83-1984		4YR5/6	4YR6/6	gSL-	m	so,po	0	8.5	LS	so,po	0	II
44	B-17	A11	0-4	c,s	7.5YR5/4	7.5YR6/3	gSL+	lfp1,v	ss,ps	0	8.1	SL	so,po	0	0
45		Al 2	4-10	a,w	7.5YR4.5/4	7.5YR5.5/3	sgL	lgr	ss,p	0	8.1	SL	so,po	0	0
46 47		Btca 2Clcsca	10-19 19-40	c,w g	5YR5.5/4 5YR6/4	5YR5.5/3 5YR7/3	gSCL gSL	2msbk m	s,p so,po	0	8.0	SL LS	so,po so,po	II I	0 IV
48		2C2csca	40-89	g	5YR2.5/6	2.5YR6/4	vgSL	m	so,po	0	8.1	LS	so,po	I	II+
49 50		2C3cs 2C4cs	89-179 179-249	d	5YR5.5/6	5YR6/4 2.5YR6/4	vgSL-	m m	so,po	0	8.5 8.3	LS LS	so, po so, po	0 0	II I+
00		20408	1/3-245	,,,	5YR5/6	2.JINO/4	vgSL-	m	so,po	J	0.3	40	30,00	V	Δ'

#### Supplementary Table 1, Part 2. Field descriptions--Continued

		ε							Wet			Assumed Material			
No.	Sample Number	Horizon	Depth (cm)	Lower Boundar	Moist y Color	Dry Color	Texture	Structure	Consis- tence	Clay Films	pН	Texture	Wet Con- sistence	Stage CaCO <sub>3</sub>	Stag Gypsu
						-	Fan 6, 5	05 ka							
51	B-3	A	0-4	c,s			gSL	v	so, po	0		SL	so,po	0	0
52		Btca	4-24	c,s	5YR5/6	7.5YR6/6	sgfSL	lfsbk	ss,p	0	8.0	SL	so,po	II	0
53		Clesca	24-42	g	7.5YR6/6	7.5YR7/4	sgSL	m	ss,ps	0	7.9	SL	so,po	Ι	ΙV
54		2C2cs	42-62	c,w	7.5YR5/6	7.5YR6/5	sgL	m	ss,p	0	8.0	L	ss,ps	0	ΙI
55		2C3cs	62-67+		2.5YR5/8	5YR6/8	sgL	m	s,p	0	8.1	L	ss,ps	0	11
56	B-18A	A11	0-2	a,s	7.5YR4.5/4	7.5YR5.5/3	gL	1fpl,v	ss,ps	0	8.1	L	so,po	0	0
57		A12	2-5	a,s	7.5YR5/4	7.5YR6/3	sgL	lgr,lfpl	ss,ps	0	8.1	L	so,po	0	0
58		Btca	5-10	a,w	5YR5.5/4	6.25YR6/4		lgr,lfsbk	s,p	0	8.0	L	so,po	I+	0
59		2C1csca	10-37	g	5YR6/4	5YR7/3	gSL	m	ss,po	0	8.2	SL	ss,ps	I	IV
60		2C2cs	37-46	ď	5YR5.5/4	5YR6.5/3		m	s,ps	0	8.1	SL	ss,ps	0	II+
51		2C3cs	46-86+		3.25YR5/7	3.25YR5/6	vgSCL	m	s,p	0	8.1	SL	ss,ps	0	II
52	B-18B	A	0-6	c,s			gL	1fp1,v	so,po	0		SL	so,po	0	0
53		Btca	6-22	c,w	6.25YR4/5	6.25YR5/4	sgSCL	2msbk	s,p	0	7.9	SL	so,po	I	0
54		2Clcsca	22-51	c,w	5YR5/5	5YR6.5/4	sgSL	m	so,po	0	8.0	SL	so,po	I	IV
55		2C2cs	51-75	c,w	5YR4/6	5YR5.5/3		m	s,p	0	8.2	SL	ss,ps	0	II+
66		2C3cs	75-130	a,s	5YR4/8	5YR5/4		m	ss,po	0	8.3	SL	so,po	0	ΙI
67		2C4cs	130-255	·+	3.25YR4/6	3.25YR5/4	vgSCL	m	s,ps	0	8.3	SL	ss,ps	0	1
							Fan 7, 5	85 ka							
68	B-4	A	0-6	a,s	7.5YR5/4	7.5YR6/4	sgSiL	v	so,po	0	8.0	SL	so, po	0	0
59		B21tcacs		a,w	7.5YR5/5	7.5YR5.5/6	gSiCL	lmsbk	ss,p	0	7.7	SL	so,po	II	I
70		B22tcsca	12-30	g	7.5YR5/6	5YR8/3	gSCL	1msbk,m	s,p	0	7.8	SL	so,po	I	III
71		2C1csca	30-55	ď	7.5YR5/6	7.5YR6/6	gSL	m	so,po	0	7.9	SL	so,po	I	IV
2		2C2cs	55-75+		7.5YR5/6	7.5YR5/6	gL	m	ss,ps	0	8.0	SL	ss,ps	0	III
3	B-19	A	0-6	a,s	7.5YR4.5/6	7.5YR5.5/4	gL'	lfpl,v	so,po	0	8.1	L	so,po	0	Q
4		B21tca	6-14	a,i	7.5YR4.5/6	7.5YR5/5	gL+	lfsbk	ss,ps	0	8.0	L	so,po	I	0_
75		B22tcsca		c,w	7.5YR5/5	7.5YR6/4	g L-	lgr	ss,ps	0	8.0	L	so,po	I	11
76		2Clcsca		d	7.5YR5/4	7.5YR6/3	gSL	m	so,po	0	8.1	SL	so, po	0	IV
77		2C2cs	40-85	d	5YR5/6	5YR5.5/4		m	s,p	0	8.1	SL	so,po	0	III
78		2C3cs	85-137	d	5YR4/5	5YR5/4	gSL-	m	so, po	0	8.2	SL	so,po	0	ΙΙ
79		3C4cs	137-262	+	5YR4/6	5YR5/5	vgSL	m	so,po	0	8.2	SL	so,po	0	Ι

#### EXPLANATION 1

#### SOIL STRUCTURE

m, massive vf, very fine (v thin) gr, granular sg, single grained f, fine (thin) pl, platey 1, weak m, medium pr, prismatic 2, moderate c, coarse (thick) cpr, columnar 3, strong vc, very coarse abk, angular blocky	Grade	Size	Type
	sg, single grained l, weak 2, moderate	f, fine (thin) m, medium c, coarse (thick)	pl, platey pr, prismatic cpr, columnar

If two structures, listed as primary and secondary  $% \left\{ 1\right\} =\left\{ 1\right\}$ 

#### SOIL TEXTURE

vsg, very slightly	co, coarse	S, sand	SCL, sandy clay loam
gravel <b>l</b> y (<5%)	f, fine	LS, loamy sand	CL, clay loam
sg, slightly gravelly	vf, very fine	SL, sandy loam	SiCL, silty clay loam
(5-20%)		L, loam	SC, sandy clay
g, gravelly (20-50%)		SiL, silt loam	C, clay
vg, very gravelly (>50%)		Si, silt	SiC, silty clay

#### SOIL CONSISTENCE

#### Wet

so, nonsticky	po, nonplastic
ss, slightly sticky	ps, slightly plastic
s, sticky	p, plastic
vs, very sticky	vp, very plastic

#### HORIZON BOUNDARIES

Distinctness	Topography						
va, very abrupt	s, smooth						
a, abrupt	w, wavy						
c, clear	i, irregular						
g, gradual	b, broken						
d, diffuse							

#### CLAY FILMS

Frequency	Thickness	Morphology							
v <sub>1</sub> , very few 1, few 2, common 3, many 4, continuous	n, thin mk, moderately thick k, thick	pf, ped face coatings br, bridging grains po, pore linings (w, occurs as waves or lamellae) co, coats on clasts							

#### STAGES OF CaCO3 AND GYPSUM2

Roman numerals indicate increasing content of  ${\tt CaCO_3}$  or  ${\tt gypsum}$ 

 $<sup>^12</sup>$ For more detailed information, see Soil Survey Staff (1975) and Birkeland (1984). For more detailed information, see section on "Field Methods and Calculations", and Bachman and Machette (1977).

#### Supplementary Table 2. Physical properties

[--, not measured. Analyst: M. C. Reheis, U.S. Geological Survey]

Mothode

Bulk density was determined by two different methods (Blake, 1965). Two to four paraffin-coated peds taken from relatively fine-textured horizons provided mean bulk densities for A, B, and some C horizons. Bulk density was measured on many gravelly C horizons in the field by weighing an excavated amount of soil and measuring the volume of the hole with water-filled plastic bags. For C horizons not measured by either of these methods, bulk density was estimated by comparison to horizons with similar gravel contents.

Gravel percent by volume was visually estimated for all horizons. Gravels were separated and weighed from bulk samples of finer textured horizons to obtain gravel weight percent as a check on the visual estimate.

Particle size of the <2-mm fraction was obtained by methods described in Day (1965). All samples were pretreated to remove organic matter by gentle heating in a weak solution of hydrogen peroxide. Gypsum and carbonates were removed from samples by gentle heating in a weak solution of sodium acetate. This method was time-consuming and required frequent washing and replacement of the solution. The sand fraction was separated by wet sieving; the dried sand was weighed and dry-sieved to obtain the various sand fractions. Silt and clay fractions were obtained by the pipette method, except for dust-trap samples, for which there was insufficient material after removal of gypsum and CaCO<sub>3</sub>. Sand was removed from both dust-trap samples by wet sieving, and the remaining sediment was combined to determine silt and clay fractions using the 5000b Sedigraph Analyzer (analyses by R. Kihl, Institute of Arctic and Alpine Research, Boulder, Colo.).

			Basal		percentage of <2-mm fraction							Bulk	percentage of <2-mm fraction						
No.	Sample No.	Horizon			n Total sand	vco sand	co sand	med sand	fi sand	vfi sand		<2μ clay		density (g/cm <sup>3</sup> )					y ficlay μ(<.5μ)
									Fa	n 1, 5	ka								
1 2 3	B-21 B-21 B-21	A 2Btca 2Cca	6 13 33	40 80 80	44.31 31.74 37.66	0.12	1.12		16.42 11.54 11.87		41.14		10.39 22.46 13.35	2.0	31.34 30.00 35.32	7.18 7.21 9.58	3.23 3.94 4.10	 	  
4	B-21	2Cn	+73	90	37.77		7.77		11.54	7.94		6.65	6.65	2.1			1.37		
									Fai	n 2, 65	ka								
5 6	B-5 B-5	A B2tca	3 11	 15	 29.19	0.39	0.28	 2.26	 5.79	 20.47	 47.40				 37.61	 6.16	 3.63	 6.58	16.83
7 8	B-5 B-5	B3tca 2Clca	20 31	15	38.63 50.52	0.16	0.58	7.11 6.59	13.07	17.71 18.20	42.79				33.62	6.30 13.98	2.87	2.58	16.00
9	B-5	2C2ca	46	80	55.23	3.44	4.05	8.25	22.23	17.26	31.78				19.13	8.12	4.53	6.81	6.18
10	B5	2C3ca	+71	80	67.40	3.53	5.87		30.28	13.86	21.46				14.80	6.07	0.59	2.82	8.32
11 12	B-15 B-15	A B1	4 12	30 10	33.98 33.65	0.25 0.19		0.99	12.31 12.10	19.91 19.96		15.36 16.06	9.46 11.49	1.6	35.96 35.68	8.83 14.59	5.87 0.00		
13 14	B-15 B-15	B2tca 2B3tca	24 45	10 80	26.94 36.89	0.33	0.57 3.27	1.57 3.42	10.28 15.52	14.19 11.25	52.38 47.53	20.68 15.58	18.45 12.69		39.10 34.06	8.24 9.88	5.03 3.58		
15 16	B-15 B-15	2Cca 2Cn	65 +165	80 80	36.02 45.68	2.49 3.64	5.22 5.45	5.33 5.48	12.22 18.65			16.06	11.07 11.03	2.1	26.90 24.26	13.79 2.67	7.23 11.80		
									<del></del>	2 100									
									ran	3, 100	ка								<del></del>
17 18	B-6A B-6A	A B2tca	4 29	20 10	37.82 38.37	0.50 0.17	1.04		14.51 15.77	19.54 19.25		15.97 18.22	11.22 14.59		31.28 33.94	8.61 6.22	6.32 3.25		
19	B-6A	2B3tca	46	70	39.83	1.11	2.00	4.42	18.60	13.69	41.09	19.08	15.87	1.5	31.32	7.24	2.53		
20 21	B-6A B-6A	2Cca 2Clcsca	88 181	80 80	31.27	2.45 8.56	3.36 6.35	3.35 4.49	12.71 10.03	9.40 5.51	42.51 32.61	26.24	21.92		31.26 17.70	7.51 8.61	3.74 6.30		
22	B-6A	2C2cs	+242	80	37.56	4.11	4.27	4.25	14.17	10.76	38.19	24.25	21.35	2.1	26.99	7.11	4.09		
23	B-6A	base of	tan 3	80	51.82	3.66	2.93	4.32	23.28	17.63	35.24	12.94	10.83	2.0	27.31	5 <b>.9</b> 0	2.03		
24 25	B-14 B-14	A11 A12	4 8	10 10	50.09 38.12	0.49 0.37			18.20 13.07		41.06 45.70	8.86 16.18	7.62 11.95		32.23 33.37	5.99 7.39	2.84 4.94		
26	B-14	B1	17	10	34.66	0.18	0.39	1.20	11.13	21.76	46.23	19.11	17.34	1.5	33.98	8.28	3.97		
27 28	B-14 B-14	B21tca B22tca	28 44	10 15	30.90 30.87	0.13		1.67 1.30	9.87 9.66	18.57 19.21		17.27 18.51	15.45 15.72		38.82 38.88	8.36 7.51	4.64		
29	B-14	2B3tca	70	80	56.88	1.48	6.53		22.32	12.30		14.22	13.10		18.42	7.00	3.47		
30 31	B-14 B-14	2Cca 2Cnca	123	80	53.65	4.26		9.55	21.43	10.69		15.22	15.22		19.03	7.95 9.60	4.15 5.84		
	D-14	2011Ca	+193	80	44.69	4.82	8.24		15.60	8.80	34.76	20.36	1/.00	2.1		9.00	J.04		
									Fan	4, 315	ka 								
32	B-16	All	4	30	20.12	0.89		0.58	7.26			22.84	17.29		37.84	12.13	7.07		
33 34	B-16 B-16	Al2 Bt	9 12		19.00 19.39	0.21		0.45	6.82 7.26	11.18 10.31		22.47 33.77	16.47 28.93		39.35 33.44	11.52 8.46	7.65 4.94		
35	B-16	Btca	28	25	19.39	0.46	0.63	0.73	7.26	10.31	46.84	33.77	28.93	1.3	33.44	8.46	4.94		
36 37	B-16 B-16	Ccacs 2Clcsca	75 137	25 50	21.72	1.56 2.21	2.09 1.74	2.28 1.45	8.87 7.46	6.91 8.19	38.76 38.50	39.52 40.46	37.03 38.08		22.89 25.67	15.87 8.79	0.00 4.04		
38	B-16	2010sca 202cs	+237		20.23			2.06	8.61			46.52			19.47	8.04	5.73		

# Supplementary Table 2. Physical properties—Continued

			Basal					percent	age of	<2-mm	fractio	n		Bu1k		centage	of <2	mm fra	ction
No.	Sample No.				n Total sand		co sand	med sand	fi sand	vfi sand		<2μ clay	<lu>clay</lu>	density (g/cm <sup>3</sup> )	co silt	med silt	vfi silt		y ficlay ιχ<.5μ)
									Fan	5, 410	) ka								
39	B7	A	6		29.67					18.96						11.24			18.35
40	B-7	Clesca	30		28.28	1.41	2.28	2.46	10.50	11.63	33.53					10.06	8.11		25.18
41	B-7	2C2csca	50	30	21.44	1.31	1.16	1.49	6.68	10.80	51.87				25.54	15.22		10.40	16.29
42 43	B-7 B-7	2C3csca 2C4cs	83 +198	40 40	28.38 47.21	2.33	2.10 2.16	2.59	9.86 19.13	11.50	41.41				26.23 19.83	7.59 7.18	7.59 4.92	12.87 5.25	17.34 15.61
43	D /	20405	T190	40	47.21	2.04	2.10	2.02	19.13	21.20	31.93				17.03	7.10	4.72	J•2J	13.01
44	B-17	Al 1	4	10	24.22						55.00	20.78	15.03	1.7	35.03	12.31	7.66		
45	B-17	A12	10	10	25.14	0.44	1.07	1.96	11.55	10.11	47.88	26.99	19.79	1.5	25.26	14.68	7.94		
46	B-17	Btca	19	15	18.35	0.74	0.98	1.54	8.02	7.08	52.96	28.69	23.47	1.3	34.22	12.95	5.78		
47	B-17	2Clcsca	40	20	31.32						33.88	34.81	32.80		20.94	6.86	6.08		
48	B-17	2C2csca	89	65	32.62	4.36	3.52		13.05	8.79	35.01	32.35	31.87		23.08	7.17	4.76		
49	B-17	2C3cs	179	65	36.57	3.20	3.69	3.79	15.49	10.40	35.68	27.75	27.25		25.90	7.16	2.61		
50	B-17	2C4cs	+249	80	32.61	4.64	3.44	2.63	12.22	9.68	42.23	25.16	22.23	2.1	30.01	7.78	4.44		
								•	Fan	6, 505	ka								
51	B-3	A	4																
52 53	B-3	Btca	24			0.88	0.53	0.68	9.17	16.60	41.88					10.94 15.64		8.76 13.31	21.50 16.15
54	B-3 B-3	Clesca 2C2cs	42 62		21.46 26.38	0.35	0.26 0.40	0.42	6.81 9.04	13.62 15.94	49.08 52.72				17.83 24.59	16.29		12.18	8.72
55	B-3	2C3cs	+67		21.67	0.66		1.76	8.67	9.27	44.79				26.42		13.42	10.14	23.40
56	B-18A	A11	2	20	28.45	1.04	0.84	1.92	12.16	12.48	50.47	21.07	15.00	1.7	35.01	10.82	4.64		
57		A12	5		24.60	0.72			10.87	10.86	50.70		17.89		33.83	11.20	5.67		
58		Btca	10		25.61	1.03	1.02		11.55	9.68	39.03	35.36	30.90		25.87	8.90	4.26		
59	B-18A	2Clcsca	37	20	36.78	2.68			19.47	9.49	39.94		17.50	1.5	22.28	9.25	8.41		
60	B-18A	2C2cs	46	20	33.83	2.27	1.55	3.55	15.25	11.21	34.09	32.09	28.90	1.5	26.57	6.41	1.10		
61	B-18A	2C3cs	+86	50	37.38	1.20	1.70	4.14	18.12	12.22	35.54	27.08	25.11	2.0	23.97	7.55	4.02		
62	B-18B	A	6	20															
63	B-18B	Btca	22	20	27.95	0.63	0.78	2.52	11.65	12.36	40.46	31.59	27.93	1.5	26.34	7.71	6.42		
64	B-18B	2Clcsca	51	25	28.61	0.60	0.85	2.75	15.33	9.08	31.65	39.74	35.81		24.27	4.89	2.49		
65	B-18B	2C2cs	75	5	36.52	1.11	1.08		19.52	11.29	33.98	29.50	27.96		21.24	9.23	3.52		
66	B-18B	2C3cs	130		42.30	6.06	4.85	6.61			31.54		22.28		20.15	8.54	2.86		
67	B-18B	2C4cs	+255	70	35.68	3.23	3.01	5.18	15.25	9.01	37.86	26.46	23.56	2.1	24.12	8.75	4.99		
									Fan	7, 585	ka .								
68	B-4	Α	6	15	28.74	1.32	0.77	1.65	11.05	13.95	45.23			:	26.02	12.99	6.22	9.30	16.73
69	B-4	B21tcacs			32.24	2.30	1.40	2.78	13.95	11.81	33.85					12.00	3.72	9.08	24.83
70	B-4	B22tcsca		25	29.61	2.92	2.05		12.60	8.62	28.14				13.39	8.37	6.38	11.81	30.44
71	B-4	2Clcsca	55	40	37.97	3.17	2.42	3.46	11.75	17.17	25.99				14.88	6.97	4.14	7.84	28.20
72	B-4	2C2cs	+75	40	50.40	1.69	1.36	6.25	28.59	12.51	20.83				10.67	6.49	3.67	10.62	18.15
73	B-19	A	6	20	21.12	1.43	0.89	1.89	8.67	8.25	56.02	22.86	17.52	1.6	36.37	12.32	7.33		
74	B-19	B2ltca	14		22.53	0.42		2.26	10.86	8.47	43.74	33.73	30.58	1.4	25.31	13.09	5.34		
75	B-19	B22tcsca	a 25	50	48.12	2.63	1.41	3.32	23.02	17.74	28.54	23.34	22.32	1.6	19.95	6.30	2.30		
76	B-19	2Clcsca	40	20	30.05	0.43	0.66	1.69	19.24	8.04	28.49	41.50	37.35		17.31	8.93	2.25		
	B-19	2C2cs	85	50	27.83	1.20	1.31	2.42	14.20	8.70	36.83	35.34	32.25	1.5	21.89	10.48	4.46		
77 78 79	B-19 B-19	2C3cs 3C4cs	137 +262	40	34.13	1.24	1.15 2.36	2.28	16.65 14.86	12.80 9.91	37.88 40.70	27.99	26.74	1.9	29.40	4.94 10.55	3.54 4.56		

# Supplementary Table 3. Extractive chemical analyses

[--, not measured. Analyst, M. C. Reheis, U.S. Geological Survey]

Organic-carbon content measured using the Walkley-Black titration procedure (Allison, 1965),  ${\sf CaCO}_3$  with the Chittick apparatus (Bachman and Machette, 1977). Several different methods of measuring gypsum content were evaluated. The most accurate method for samples containing <5% gypsum measures the conductivity of gypsum in solution (Bower and Huss, 1948; Jackson, 1958). Greater gypsum contents are more easily and accurately measured by the weight loss upon converting gypsum to anhydrite in a  $100^\circ$ C oven. This method has a standard error of estimate of  $\pm 1.8\%$ , compared to the method which calculates gypsum from  ${\sf SO}_4^{\rm T}$  concentration (Nelson and others, 1978). Early in this study, gypsum in some samples was measured by the weight loss caused by repeated washing of each sample in distilled water. This method is time-consuming and may also result in measurement error if any carbonate is present and is dissolved with the gypsum.

			Basal		percenta	ige	
	Sample		depth	Organi	с		pН
No.	number	Horizon	(cm)	c	CaCO <sub>3</sub>	Gypsum l	1:1H <sub>2</sub> C
			Fan 1,	5 ka			
1	B-21	Α	6	0.55	23.2	0.08	7.7
2	B-21	2Btca	13	0.71	35.8	0.07	8.0
3	B-21	2Cca	33	0.37	59.7	0.09	8.1
4	B-21	2Cn	+73	0.22	53.6	0.08	8.4
	*		Fan 2, 6	5 ka			
5	B-5	Α	3				
6	B-5	B2tca	11	0.40	15.3	0.14	8.6
7	B-5	B3tca	20	0.35	29.8	0.09	8.4
8	B-5	2C1ca	31	0.42	59.0	0.36	8.6
9	B-5	2C2ca	46	0.43	57.5	0.50	8.9
10	B-5	2C3ca	+71	0.14	60.3	0.51	8.9
11	B-15	A	4	0.49	15.4	0.07	8.3
12	B-15	B1	12	0.55	12.8	0.10	8.2
13	B-15	B2tca	24	0.78	17.2	0.07	8.1
14	B-15	2B3tca	45	0.16	44.5	0.09	8.1
15	B-15	2Cca	65	0.29	66.8	0.44	8.5
16	B-15	2Cn	+165	0.15	63.2	0.73	8.4
			Fan 3, 10	00 ka			
17	B-6A	Α	4	0.46	17.7	0.07	8.1
18	B-6A	B2tca	29	0.22	10.9	0.22	8.1
19	B-6A	2B3tca	46	0.32	52.0	0.15	8.3
20	B-6A	2Cca	88	0.26	68.7	0.70	8.2
21	B-6A	2Clcsca	181	0.03	73.8	1.26	8.2
22	B-6A	2C2cs	+242	0.23	65.9	0.15	8.1
23	B-6A	base of fa	n 3	0.00	8.5	0.10	55.8
24	B-14	A11	4	0.44	19.7	0.06	8.0
25	B-14	A12	8	0.52	23.1	0.09	8.0
26	B-14 B-14	Bl B21tca	17	0.61	13.1	0.12	8.0
	B-14 B-14	B21tca B22tca	28 44	0.41	13.8 12.3	0.07	8.1
27 28		paalla	44	0.40	14.3	0.07	8.1
28		2B3tca	70	0.40	46 O	0.11	Q /
	B-14 B-14	2B3tca 2Cca	70 123	0.40	46.0 65.7	0.11	8.4

			Basal		pe	rcenta	ige	
No.	Sample number	Horizon	depth (cm)	(	)rganic		Gypsum l	рН 1:1Н <sub>2</sub> 0
			Fan 4, 3	315 ka				
32	B-16	A11	4		0.95	31.6	0.10	8.0
33	B-16	A1 2	9			25.1	0.07	8.0
34 35	B-16	Bt Btca	12 28			49.8 49.8	0.23	8.1
36	B-16 B-16	Ccacs	75			10 1		8.4
37	B-16	2Clcsca	137		0.01			8.5
38	B-16	2C2 cs	+237		0.04	38.3	<sup>2</sup> 29.6	8.6
			Fan 5, 4	10 ka				
39	B-7	Α	6			37.3	0.49	7.8
40	B-7	Closca	30			24.3	358.8	8.0
41	B-7	2C2csca	50				345.2 342.8	8.0 8.3
42 43	B-7 B-7	2C3csca 2C4cs	83 +198			41.0 57.3	314.1	8.5
44	B-17	All	4			25.2	0.18	8.1
45	B-17	A12	10			30.3	0.11	8.1
46	B-17	Btca	19			42.1	0.29 270.3	8.0 8.1
47 48	B-17 B-17	2Clesea 2C2esea	40 89		0.09	14.1 56.5	228.0 237.2	8.1
49	B-17	2C3cs	179		0.08	39.6	237.2	8.5
50	B-17	2C4cs	+249			49.7	227.2	8.3
			Fan 6, 5	05 ka				
51	B- 3	Α	4					
52	B-3	Btca	24			33.6	0.008	8.0
53	B-3	Clesca	42			13.3	348.6	7.9
54 55	B-3 B-3	2C2cs 2C3cs	62 +67		0.18	11.8 19.3	342.5 333.5	8.0 8.1
56	B-18A	All	2		0.34	33.2	0.21	8.1
57	B-18A	A1 2	5		0.42	36.0	0.16	8.1
58	B-18A	Btca	10		0.47	49.8	2 0.14	8.0
59	B-18A	2C1csca	37		0.13	20.4	260.6 232.6	8.2
60 61	B-18A B-18A	2C2cs 2C3cs	46 +86			43.1 33.0	<sup>2</sup> 32.6 <sup>2</sup> 23.1	8.1 8.1
01	D TON	20300	. 55					
62	B-18B	A	6			34.6	0.37	
63	B-18B	Btca	22			24.6	258.8	7.9
64	B-18B	2Clcsca	51 75					8.0 8.2
65 66	B-18B B-18B	2C2cs 2C3cs	75 130		).16 ).26	9.0 45.3	214.1	8.3
67	B-18B	2C4cs	+255			44.2	214.1 28.5	8.3
			Fan 7, 5	85 ka				
68	B-4	A	6			34.8	0.01	8.0
69	B-4	B21tcacs	12		0.55		2 77	7.7
70	B- 4	B22tcsca	30	(		16.9	3 <sub>51.4</sub>	7.8
71 72	B-4 B-4	2Clesea 2Cles	55 +75			24.3	3 <sub>40.7</sub> 3 <sub>43.2</sub>	7.9 8.0
	B-19	A	6			28.9	0.10	8.1
73		D214	14			20.7	0.11	8.0
74	B-19	B21tca	17					
74 75	B-19 B-19	B22tcsca	25	(	0.43	222	220.2	8.0
74 75 76	B-19 B-19 B-19	B22tcsca 2Clcsca	25 40	(	0.43 0.18	22.2	<sup>2</sup> 20.2 <sup>2</sup> 56.0	8.0 8.1
74 75	B-19 B-19	B22tcsca	25	(	0.43 0.18 0.25	22.2 20.0	220.2	8.0

 $<sup>^1</sup>$  Gypsum measured by electrical conductivity except where numbered.  $^2$  Values measured by weight loss due to heating.  $^3$  Values measured by weight loss from solution in distilled water.

# Supplementary Table 4. Clay mineralogy by X-ray diffraction

[--, not measured. Analyst: M. C. Reheis, U.S. Geological Survey]

Samples of clay fractions were drawn from the settling tubes after

Samples of clay fractions were drawn from the settling tubes after particle-size analyses and were plated on ceramic tiles (Whittig, 1965). X-ray diffraction traces (CuK alpha radiation) were run on these oriented clays after each of the following treatments: air-dried, glycolated, and heated to  $400^{\rm QC}$  and  $550^{\rm QC}$ . At least five clay ninerals were identified in the Kane fan soils, including kaolinite, mica, smectite, and palygorskite. X-ray traces revealed an unknown mineral with a primary peak at 11.4 A in the air-dried state that shifts down to 10.6 A after glycolation. Gentle heating to about  $80^{\rm QC}$  gives a peak at 8.66 A that does not shift upon subsequent glycolation. Heating to  $440^{\rm QC}$  destroys the mineral peaks. This mineral concentrates in the coarse clay fraction, but is not found in the silt fraction. Peaks of the unknown mineral increase with age of the soil subsequent glycolation. Heating to 400°C destroys the mineral peaks. This mineral concentrates in the coarse clav fraction, but is not found in the silt fraction. Peaks of the unknown mineral increase with age of the soil (suggesting a pedogenic origin) and are largest in gypsum-rich horizons. These peaks and their behavior during the various treatments do not coincide with those of any cummon clay mineral. The main peak and some of the minor ones roughly fit the zeolite mineral offretite (Chih-Chun Kao and G. W. Brindley, written commun., 1983), but zeolites should not be affected by glycolation or heating to 400°C (R. L. Hay and R. A. Sheppard, written communs., 1983). One additional candidate is the Ca-rich sepicite-like mineral tubermorite, which is thought to be rare in nature but is commonly found in mortar and concrete (D. McGrath and B. L. Allen, oral communs., 1983). Microprobe analyses of the unknown mineral revealed no concentrations of elements uncommon in silicate clays.

A modification of the Illinois State Survey method was employed to estimate clay mineral percentages, in which peak heights are measured from a curvilinear baseline drawn on the glycolated trace. Heights of the 10 A peak (mica) and the 7 A peak (kaolinite) were multiplied by factors of 2 and 3 respectively, to adjust for intrinsic differences in ability of various clays to produce peaks in soils (Guccione, 1982; Johns and others, 1954). Prominent peaks for smectite (17 A), the unknown mineral (10.6 A), and palygorskite (10.4 A) were not adjusted. The resultant peak heights are summed and used to obtain clay-mineral percentages. This arbitrary treatment is not to be interpreted as indicating actual clay-mineral percentages, but only relative changes in the proportion of the various clay minerals present.

minerals present.

percentage of clay mineral

			Basal	pe	rcenta	ige or	cray miner	- d1
No.	Sample number	Horizon	depth (cm)	Kaoli- nite		Smec- tite	Paly- gorskite	Unknown mineral
			Fan I	, 5 ka				
1	B-21	A	6	36	40	24	0	0
2	B-21	2Btca	13	35	42	23	0	0
3	B-21	2Cca	33	17	25	17	9	32
4	B-21	2Cn	+73	17	23	17	10	33
			Fan 2	, 65 ka				
5	B-5	A	3					
6	B-5	B2tca	11					
7	B-5	B3tca	20					
8	B-5	2Clca	31					
9	B-5	2C2ca	46					
10	B-5	2C3ca	+71					
11	B-15	A	4	31	37	32	0	0
12	B-15	B1	12	27	44	29	O	0
13	B-15	B2tca	24	27	41	32	0	0
14	B-15	2B3tca	45	28	21	30	6	15
15	B-15	2Cca	65	30	32	31	Ð	7
16	B-15	2Cn	+165	29	28	26	6	11
			Fan 3,	. 100 ka				
17	B- 6A	Α.	4	32	33	35	0	0
18	B-6A	B2tca	29	27	34	39	0	ñ
19	B-6A	2B3tca	46	29	27	34	10	0
20	B-6A	2Cca	88	30	31	30	9	0
21	B-6A	2Clcsca	181	37	37	20	6	0
22	B-6A	2C2cs	+242	36	39	25	0	0
23	B-6A	base of f	an 3	27	34	30	0	9
24	B-14	A11	4	22	47	31	0	0
25	B-14	A12	. 8	30	41	29	0	0
26	B-14	B1	17	28	45	27	0	0
27	B-14	B21tca	28	28	34	38	0	0
28	B-14	B2 2t ca	44	23	35	35	7	0
29	B-14	2B3tca	70	22	24 30	31	5	18
30	B-14	2Cca 2Cnca	123	28	30	31	0	11 8
31	B-14	zuica	+193	36	30	26	U	8

			Basal				clay mine	
No.	Sample number	Horizon	depth (cm)	Kaoli- nite		Smec- tite	Paly- gorskite	Unknow
			Fan 4,	315 ka				
32	B-16	A11	4	22	45	33	0	0
33	B-16	A12	9	26	50	24	0	0
34	B-16	Bt	12	21	34	34	11	0
35	B-16	Btca	28 75	21	34 23	34 42	11	0 7
36 37	B-16 B-16	Ccacs 2Clcsca	137	21 35	20	33	7 6	6
38	B-16	2C2cs	+237	34	17	37	6	6
			Fan 5,	410 ka				
39	B-7	A	6					
40	B-7	Clesca	30	~~				
41	B-7	2C2csca	50	~-				
42 43	B-7 B-7	2C3csca 2C4cs	83 +198					
44 45	B-17 B-17	A1 1 A1 2	4 10	30 23	40 39	30 25	0 13	0
45 46	B-17	Btca	19	16	25	21	8	30
47	B-17	2Clcsca	40	9	25	13	9	44
48	B-17	2C2csca	89	28	26	28	9	9
49 50	B-17 B-17	2C3cs 2C4cs	179 +249	21 10	30 37	30 27	10 10	9 16
			Fan 6,	505 ka				
51	В-3	Α	4	~-				
52	B-3	Btca	24	~				
53	B-3	Clesca	42	~~				
54	B-3	2C2cs	62	~-				
55	B-3	2C3cs	+67	~-				
56	B- I 8A	AII	2	26	46	28	0	0
57	B-18A B-18A	Al2 Btca	5 10	20	39 25	27 28	14 8	0
58 59	B-18A	2Clcsca	37	31 7	11	3	7	8 72
60	B-18A	2C2cs	46	29	20	26	13	12
61	B-18A	2C3cs	+86	24	21	38	7	10
62	B-18B	A	6					
63	B-18B	Btca	22	27	32	29	12	0
64	B-18B	2Clesea	51 75	24	22 8	39 44	15 4	0 <b>3</b> 0
65 66	B-18B B-18B	2C2cs 2C3cs	130	14 23	23	33	8	13
67	B-18B	2C4cs	+255	24	22	37	7	10
			Fan 7,	585 ka				
ь8	B-4	Α	6					
69	B-4	B21tcacs	12					
70	B-4	B22tcsca	30	~-				
71 72	B-4 B-4	2Clesea 2C2es	55 +75					
/ 2	D-4	20208	+/3					
7 3	B-19	A	6	30	24	38	8	0
74	B-19	B2ltca	14	19	17	4.5	12	7 4
75 76	B-19	B22tcsca	25	19	16	50	11	. 4
	B-19 B-19	2Clesea 2C2es	40 85	18 17	18 13	41 28	6 5	17 37
	B-19 B-19	2020s 2030s	137	28	16	39	5	12
77 78		3C4cs	+262	24	19	42	6	9

# Supplementary Table 5. Total chemical analyses of the fine fraction by induction-coupled plasma spectroscopy

[--, not measured. Analyst: A. Bartel under J. E. Taggart, U.S. Geological Survey]

### Methods

The less-than-2-mm and the silt-plus-clay fractions of selected soil profiles and dust-trap samples were analyzed for major elements and Zr by the Analytical Laboratories of the U.S. Geological Survey. To avoid chemical contamination in obtaining the silt-plus-clay fraction, soils were dispersed by shaking samples overnight in distilled water and then sonicating them in a water bath. Sand was removed by wet-sieving, and the silt-plus-clay fraction was dried and ground.

Oxides were determined by inductively coupled argon plasma emission spectroscopy (ICP). This method is noted for its high sensitivity to most elements once they are in solution (Taggart and others, 1981)(ICP accuracy is estimated at 3-5% and precision at 1-2%; J. E. Taggart, personal communication, 1982). X-ray fluorescence (XRF) is a more accurate method, but it is not sensitive to small amounts of such elements as Zr (there is less than 0.1% Zr in the silt-plus-clay fractions of soils in the study area). Because determination of Zr was critical to this study, all samples were run using ICP, with revised pretreatments that dissolved Zr. Selected samples were split and analyzed using both ICP and XRF as a check on the accuracy of ICP, and the separate analyses were in very close agreement on all samples (data not shown).

	Sample		Basal			P	ercent	age of s	ilt-plus	-clay f	raction		
No.		Horizon	depth (cm)	SiO2	A1 <sub>2</sub> 0 <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub>	MgO	Ca0	Na <sub>2</sub> 0	к <sub>2</sub> 0	TiO <sub>2</sub>	Mn O	Zr0
						Fan 1,	, 5 ka						
1	B-21	A	6	62	10.34	4.01	3.32	6.42	1.06	2.18	0.63	0.068	0.105
2	B-21	2Btca	13	48	10.98	4.18	3.91	12.8	0.5 <b>9</b>	2.29	0.55	0.043	0.049
3	B-21	2Cca	33	28	4.42	1.65	8.46	25.6	0.34	1.22	0.22	0.037	0.024
4	B-21	2Cn	+73	36	6.03	2.20	7.23	20.1	0.49	1.66	0.33	0.047	0.047
						Fan 2,	65 ka						
5	B-5	Α	3										
6	B-5	B2tca	11										
7	B-5	B3tca	20										
8	B-5	2C1ca	31										
9	B-5	2C2ca	46										
10	B-5	2C3ca	+71										
11	B-15	A	4	59	12.14	4.75	3.83	6.00	1.02	2.53	0.72	0.076	0.092
12	B-15	Bl	12	61	11.61	4.48	3.30	6.64	1.13	2.41	0.68	0.059	0.090
13	B-15	B2tca	24	58	12.04	4.58	3.55	8.06	1.09	2.41	0.65	0.053	0.074
14	B-15	2B3tca	45	39	7.18	2.66	5.56	20.8	0.80	1.57	0.43	0.035	0.050
15	B-15	2Cca	65	44	8.00	2.76	6.93	15.7	1.04	1.93	0.43	0.043	0.033
16	B-15	2Cn	+165	46	9.55	3.28	5.75	13.8	1.12	2.17	0.50	0.048	0.030
					I	Fan 3,	100 ka	l					
17	B-6A	A	4										
18	B-6A	B2tca	29										
19	B-6A	2B3tca	46										
20	B-6A	2Cca	88										
21	B-6A	2Clcsca	181								~-		
22	B-6A	2C2cs	+242										
23	B-6A	base of	fan 3										
24	B-14	A1 1	4	69	9.64	3.82	2.85	4.39	1.24	2.04	0.65	0.074	0.127
25	B-14	A1 2	8	54	9.00	3.38	2.44	3.92	0.90	1.88	0.52	0.057	0.081
26	B-14	B1	17	60	11.02	4.06	3.03	6.70	1.05	2.27	0.58	0.058	0.080
27	B-14	B21tca	28	64	10.72	3.93	2.75	6.43	1.25	2.12	0.60	0.052	0.073
28	B-14	B22tca	44	64	10.84	4.06	2.68	5.61	1.18	2.18	0.63	0.048	0.080
29 30	B-14	2B3tca	70	41	7.37	3.08	4.98	19.9	0.77	1.24	0.43	0.044	0.069
30 31	B-14 B-14	2Cca 2Cnca	123 +193	41 47	7.84 9.96	3.25 3.85	7.28 5.85	15.5 11.6	1.11 0.96	1.43 1.66	0.43 0.53	0.066 0.099	0.057

 $\textbf{Supplementary Table 5.} \ \ \textbf{Total chemical analyses of the fine fraction by induction-coupled plasma}$ spectroscopy--Continued

	C+: 1		Basal			P	ercent	age of s	silt-plus	-clay f	raction	1	
No.	Sample Number	Horizon	depth (cm)	SiO <sub>2</sub>	A1 <sub>2</sub> 0 <sub>3</sub>	Fe <sub>2</sub> 0 <sub>3</sub>	MgO	Ca0	Na <sub>2</sub> O	к <sub>2</sub> 0	TiO <sub>2</sub>	Mn O	Zr0
					]	Fan 4,	315 ka	3					
32	B-16	A1 1	4	58	10.68	4.01	3.62	8.78	1.13	2.29	0.63	0.061	0.079
33	B-16	A12	9	59	10.93	4.02	3.35	8.71	1.13	2.29	0.60	0.052	0.073
34	B-16	Bt	12	40	7.56	2.78	4.73	21.1	0.74	1.33	0.38	0.034	0.037
35 36	B-16 B-16	Btca Ccacs	28 75	40 28	7.56 5.26	2.78 1.92	4.73 9.87	21.1 23.2	0.74 0.65	1.33 0.84	0.38 0.25	0.034	0.037
37	B-16	2Clcsca	137	16	2.84	1.10	7.89	33.6	0.43	0.48	0.13	0.027	0.014
38	B-16	2C2cs	+237	40	9.40	3.42	8.77	14.0	0.88	1.69	0.40	0.043	0.019
					]	Fan 5,	410 ka	a					
39	В-7	A	6										
40	B-7	Clcsca	30										
41	B-7	2C2csca	50										
42	B-7	2C3csca	83										
43	B-7	2C4cs	+198										
44	B-17	A1 1	4	56	10.00	3.55	3.76	10.7	0.96	2.11	0.52	0.058	0.057
45	B-17	A12	10	51	9.55	3.39	3.85	12.3	0.88	1.98	0.47	0.053	0.051
46 47	B-17 B-17	Btca 2Clcsca	19	35	6.99	2.58	5.87	21.3	0.51	1.33	0.37	0.032	0.034
48	B-17	2C1csca 2C2csca	40 89	19 21	3.82 3.67	1.39 1.56	5.11	25.7 25.3	0.22 0.27	0.63 0.73	0.18 0.18	0.022	0.019
49	B-17	2C3cs	179	34	6.37	2.42	8.82	18.6	0.54	1.22	0.32	0.050	0.027
50	B-17	2C4cs	+249	34	6.03	2.26	8.77	19.9	0.57	1.24	0.32	0.048	0.030
					]	Fan 6,	505 ka	a					
51	в-3	A	4			<del></del>							
52	B-3	Btca	24										
53	B-3	Clcsca	42										
54	B-3	2C2cs	62										
55	B-3	2C3cs	+67										
56	B-18A	A11	2	53	9.60	3.61	3.62	11.7	1.01	2.05	0.57	0.059	0.075
57	B-18A	A12	5	48	9.21	3.40	3.90	15.7	0.90	1.93	0.50	0.052	0.053
58	B-18A	Btca	10										
59	B-18A	2Clcsca	37										
60 61	B-18A B-18A	2C2cs 2C3cs	46 +86										
62	B-18B	A	6	50	9.40	3.51	3.76	13.7	0.95	1.99	0.54	0.055	0.064
63	B-18B	Btca	22	55	10.34	3.89	3.90	9.82	0.92	1.93	0.55	0.045	0.062
64	B-18B	2Clcsca	51	25	4.20	1.69	4.33	25.3	0.55	0.72	0.27	0.032	0.038
65	B-18B	2C2cs	75	43	7.83	3.02	4.68	14.1	0.80	1.20	0.45	0.069	0.049
66	B-18B	2C3cs	130	33	5.94	2.96	8.42	18.9	0.84	1.20	0.33	0.125	0.027
67 	В-18В	2C4cs	+255	42	8.11	3.36	6.29	15.7	1.01	1.81	0.45	0.085	0.038
					]	Fan 7,	585 ka	1					
68	B-4	A	6										
69 70	B-4	B21tcacs	12										
70 71	B-4 B-4	B22tcsca 2Clcsca	30 55										
72	B-4	2C2cs	+75										
73	B-19	A	6	56	9.55	3.65	3.58	10.1	0.85	1.93	0.50	0.089	0.055
74	B-19	B21tca	14	49	9.13	4.28	3.28	13.1	0.65	1.58	0.49	0.144	0.043
75	B-19	B22tcsca		41	6.43	2.47	3.32	19.3	0.44	1.12	0.33	0.062	0.047
76	B-19	2Clcsca	40	28	5.03	2.05	7.71	23.1	0.32	0.77	0.25	0.107	0.022
77 78	B-19 B-19	2C2cs 2C3cs	85 137	41 43	7.64	3.10 3.29	7.40 7.18	14.8	0.46 0.55	1.20	0.38 0.40	0.092 0.134	0.036
79	B-19	3C4cs	+262	43 47	7.58 8.73	3.71	7.18 4.15	13.9 13.9	0.65	1.39 1.75	0.40	0.134	0.030
. ,	J 19	30408	. 202	47	0.13	J•/1	4.1)	1,0.7	····	1.13	U•+/	0.143	

# Supplementary Table 6. Total chemical analyses of the less-than-2-mm fraction by inductioncoupled plasma spectroscopy

[--, not measured. Analyst: P. H. Briggs under L. R. Layman, U.S. Geological Survey]

Methods Same methods as in supplementary table 5.

	Sample		Basal depth			P	ercent	age of l	ess-than	-2mm fi	action		
No.	-	Horizon		$\sin \theta_2$	A1203	Fe <sub>2</sub> 0 <sub>3</sub>	Mg0	Ca0	Na <sub>2</sub> 0	к <sub>2</sub> 0	$TiO_2$	Mn O	Zr0 <sub>2</sub>
						Fan l	, 5 ka						
1 2	B-21	A 28+	6	60	7.90	2.98	3.33	10.0	1.09	1.81	0.43	0.057	0.056
3	B-21 B-21	2Btca 2Cca	13 33	48 22	7.56 2.68	2.72 1.13	4.05 8.28	15.8 29.2	0.80 0.38	1.81 0.84	0.37 0.12	0.045	0.037
4	B-21	2Cn	+73	21	2.33	1.24	8.59	30.1	0.39	0.84	0.10	0.034	0.009
						Fan 2,	65 ka						
5	B-5	A	3										
6 7	B−5 B−5	B2tca B3tca	11 20										
8	B-5	2Clca	31										
9	B-5	2C2ca	46										
10	B-5	2C3ca	+71										
11	B-15	A	4	63	8.77	3.32	3.30	8.21	1.08	1.93	0.50	0.061	0.067
12	B-15	B1	12	63	9.17	3.40	2.82	7.06	1.09	2.05	0.50	0.056	0.061
13	B-15	B2tca	24	60	9.24	3.39	3.30	9.79	1.08	1.93	0.47	0.050	0.056
14 15	B-15	2B3tca	45	37 29	4.46	1.86	6.24	24.8	0.55	1.08	0.25	0.043	0.026
16	B-15 B-15	2Cca 2Cn	65 +165	28	3.46 3.23	1.53 1.52	8.26 8.13	26.7 27.8	0.50 0.49	0.96 0.96	0.18 0.17	0.041 0.045	0.014
					I	an 3,	100 ka	1					
17	B-6A	A	4										
18 19	B-6A B-6A	B2tca 2B3tca	29 46										
20	B-6A	2Cca	88										
21	B-6A	2Clcsca	181										
22	B-6A	2C2cs	+242										
23	B-6A	base of	ran 3										
24	B-14	A11	4	62	7.45	2.69	3.18	8.41	1.21	1.69	0.42	0.059	0.058
25	B-14	A1 2	8	63	8.83	3.12	2.95	6.94	1.11	1.93	0.45	0.059	0.055
26	B-14	B1	17	63	9.47	3.35	2.79	6.81	1.13	2.17	0.48	0.052	0.050
27 28	B-14 B-14	B21tca B22tca	28 44	63 64	9.64 9.75	3.43 3.49	2.77 2.67	7.05 6.32	1.27 1.21	2.05 2.05	0.50 0.52	0.049	0.051
29	B-14	2B3tca	70	37	4.61	1.95	5.79	21.0	0.75	0.96	0.23	0.039	0.023
30	B-14	2Cca	123	25	3.23	1.36	8.86	26.6	0.58	0.84	0.15	0.037	0.014
31	B-14	2Cnca	+193	29	4.12	1.86	8.14	24.6	0.62	0.96	0.20	0.077	0.015
					F	an 4,	315 ka						
						2 00	3.98	14.3	0.96	1.81	0.43	0.050	0.046
32	B-16	A1 1	4	53	8.09	2.89	2.0						
33	B-16	A12	9	53	8.45	3.03	3.60	14.3	0.88	1.81	0.43	0.048	0.040
33 34	B-16 B-16	A12 Bt	9 12	53 38	8.45 5.67	3.03 2.13	3.60 4.43	14.3 23.6	0.63	1.08	0.43 0.30	0.028	0.028
33 34 35	B-16 B-16 B-16	A12 Bt Btca	9 12 28	53 38 38	8.45 5.67 5.67	3.03 2.13 2.13	3.60 4.43 4.43	14.3 23.6 23.6	0.63 0.63	1.08 1.08	0.43 0.30 0.30	0.028 0.028	0.028 0.028
33 34	B-16 B-16	A12 Bt	9 12	53 38	8.45 5.67	3.03 2.13	3.60 4.43	14.3 23.6	0.63	1.08	0.43 0.30	0.028	0.028

 $\textbf{Supplementary Table 6.} \quad \textbf{Total chemical analyses of the less-than-2-mm fraction by induction-coupled plasma spectroscopy--Continued} \\$ 

	Ca1 -		Basal			P	ercent	age of l	ess-than	-2mm f	raction		
No.	Sample number	Horizon	depth (cm)	Si0 <sub>2</sub>	A1203	Fe <sub>2</sub> 0 <sub>3</sub>	Mg0	Ca0	Na <sub>2</sub> 0	к <sub>2</sub> 0	TiO <sub>2</sub>	Mn O	Zr0 <sub>2</sub>
					]	Ean 5,	410 ka	1					
39	В-7	A	6										
40	B-7	Clcsca	30										
41	B-7	2C2csca	50										
42	B-7	2C3csca	83										
43	B-7	2C4cs	+198										
44	B-17	A11	4	56	8.96	3.13	3.62	11.1	1.05	2.05	0.47	0.057	0.049
45	B-17	A12	10	51	8.37	2.88	3.83	13.8	0.93	1.93	0.40	0.045	0.035
46	B-17	Btca	19	36	5.95	2.09	5.94	22.5	0.58	1.20	0.28	0.031	0.021
47	B-17	2Clcsca	40	9	1.19	0.46	2.31	29.5	0.19	0.24	0.05	0.006	0.004
48	B-17	2C2csca	89	16	2.04	1.09	9.67	27.4	0.23	0.48	0.12	0.024	0.008
49	B-17	2C3cs	179	23	3.35	1.43	7.96	23.9	0.38	0.72	0.18	0.030	0.014
50	B-17	2C4cs	+249	24	3.25	1.32	7.89	25.6	0.38	0.72	0.17	0.028	0.017
					]	Fan 6,	505 ka	1					
51	B-3	A	4										
52	B-3	Btca	24										
53	B-3	Clcsca	42										
54	B-3	2C2cs	62										
55	B-3	2C3cs	+67										
56	B-18A	A11	2	55	7.54	2.85	3.45	13.9	0.93	1.69	0.40	0.074	0.042
57	B-18A	A12	5	51	7.64	2.92	3.75	15.9	0.82	1.57	0.40	0.071	0.042
58	B-18A	Btca	10										
59	B-18A	2Clcsca	37										
60	B-18A	2C2cs	46										
61	B-18A	2C3cs	+86										
62	B-18B	A	6	53	7.59	2.89	3.60	14.9	0.87	1.63	0.40	0.072	0.042
63	B-18B	Btca	22	56	8.39	3.23	3.90	11.0	0.84	1.57	0.42	0.063	0.039
64	B-18B	2Clcsca	51	24	2.89	1.27	3.25	23.5	0.38	0.60	0.17	0.031	0.019
65	B-18B	2C2cs	75	25	4.20	1.63	7.55	23.1	0.49	0.84	0.18	0.026	0.013
66	B-18B	2C3cs	130	32	3.72	3.08	6.57	22.5	0.53	0.84	0.20	0.193	0.019
67	B-18B	2C4cs	+255	41	4.88	2.58	5.24	20.3	0.65	1.20	0.27	0.094	0.025
					1	Fan 7,	585 ka	<u> </u>					
68	B-4	A	6										
69	B-4	B21tcacs	12										
70	B-4	B22tcsca	30										
71	B-4	2Clcsca	55										
72	B-4	2C2cs	+75										
73	B-19	A	6	58	8.51	3.42	3.22	10.8	0.84	1.81	0.45	0.106	0.044
74	B-19	B21tca	14	51	7.50	3.72	2.82	14.4	0.59	1.45	0.38	0.152	0.029
75	B-19	B22tcsca		47	4.10	1.63	2.32	18.0	0.36	0.84	0.22	0.050	0.028
76	B-19	2Clcsca	40	20	2.63	1.19	4.15	24.2	0.26	0.48	0.13	0.066	0.011
77	B-19	2C2cs	85	33	4.39	2.05	4.38	19.2	0.35	0.84	0.23	0.079	0.018
78	B-19	2C3cs	137	42	5.14	2.56	5.01	16.2	0.49	1.08	0.27	0.114	0.024
, 0											0.32		

## Supplementary Table 7, Part 1. Horizon weights of sand, silt, clay, carbon, CaCO3, gypsum, and clay minerals

[--, not measured. Analyst: Marith C. Reheis, U.S. Geological Survey]

## Methods

Profile weights of soil properties (g/cm²/column of soil) are calculated from the property percentage, bulk density, and texture of each horizon. The pedogenic increase from the amount of the property in the parent material is calculated by estimating the property percentage, bulk density, and texture in the original deposit (method modified from Machette, 1978, and Machette, 1985). Weights are calculated for each horizon, summed to give the total profile weight in a soil, and standardized to 250 cm depth to eliminate bias caused by different denths of sampling (hardbog nits us handday holes). by different depths of sampling (backhoe pits vs. hand-dug holes).

Two basic equations are used to calculate pedogenic increase for clay, silt,  $CaCo_3$ , and gypsum. Equation 1 is used to calculate pedogenic increase in horizon weight of clay and silt (substitute appropriate silt percentages when calculating silt weights):

1. 
$$g \text{ clay} = (\frac{c1ay\%}{100})(BDf)(\frac{100 - salts\%}{100})(H)) - (\frac{i \text{ clay\%}}{100})(i \text{ BDf})(\frac{100 - i \text{ salts\%}}{100})(H))$$

where BDf is bulk density of the <2-mm fraction, H is horizon thickness, salts include organic matter,  $CaCO_3$ , and gypsum, and the letter i designates an estimate of the parent material state. This equation includes a factor to account for salts and organic matter because particle size is measured on an organic-free and salt-free basis. Equation 2 is used to calculate pedogenic weights of gypsum and  $CaCO_3$  (again, substitute  $CaCO_3$  percentages in place of gypsum percentages when calculating  $CaCO_3$  weights):

2. 
$$g \text{ gypsum} = ((\frac{\text{gypsum}\%}{100})(BDf)(H)) - ((\frac{\text{i gypsum}\%}{100})(\text{i BDf})(H))$$

Profile weights for sand, carbon, clay minerals, and major oxides are not corrected for the estimated parent material contents. Values for sand are calculated using the first term of equation 1 above, substituting the appropriate sand percentages for clay percentages. Values for carbon and major oxides (Supplementary Table 7, Part 2) are calculated using the first term of equation 2 above, substituting the appropriate property percentages for gypsum percentage. Horizon weights for clay minerals are calculated by equation 3:

3. g clay mineral = 
$$(\frac{\text{clay mineral}\%}{100})(\frac{\text{clay}\%}{100})(\text{H})(\text{BDf})(\frac{100 - \text{salts}\%}{100})$$

			Page 1				We	eight (g	/cm²/hori	zon col	umn)			
No.	Samp numbe	le er Horizon	Basal depth (cm)	Sand	Silt <sup>l</sup>	Clay	L Carbon	CaCO3 <sup>1</sup>	Gypsum <sup>1</sup>	Kaoli- nite	Mica	Smec- tite		Unknown e mineral
							Fan 1,	, 5 ka						
1 2 3 4	B-21 B-21 B-21 B-21	A 2Btca 2Cca 2Cn	6 13 33 +73	1.02 0.57 0.97 1.47	0 0.16 -0.20 0.26	0.18 0.41 0.13 0	0.017 0.020 0.024 0.018	-1.61 -0.38 0.30	-0.0005 0.0002 0.0012 0.0008	0.11 0.17 0 0.044	0.13 0.20 0 0.064	0.077 0.11 0.059 0.044	0 0 0.031 0.026	0 0 0.11 0.082
							Fan 2,	65 ka						
5 6 7 8 9	B-5 B-5 B-5 B-5 B-5 B-5	A B2tca B3tca 2C1ca 2C2ca 2C3ca	3 11 20 31 46 +71					   		  			   	
11 12 13 14 15	B-15 B-15 B-15 B-15 B-15 B-15	A B1 B2tca 2B3tca 2Cca 2Cn	4 12 24 45 65 +165	0.89 3.10 3.27 1.29 0.99 6.92	0.88 3.42 4.51 0.59 0.12	0.24 1.03 1.82 0.14 0	0.015 0.059 0.115 0.010 0.024 0.063	-1.24 -3.26 -4.39 -1.35 0.99 3.44	0 0.0048 0.0015 0.0004 0.0311 0.277	0.13 0.40 0.68 0.15 0.13 0.69	0.15 0.65 1.03 0.11 0.14 0.71	0.13 0.43 0.81 0.16 0.14 0.61	0 0 0 0.033 0	0 0 0 0.082 0.031 0.26
							Fan 3,	100 ka				· · · · · · · · · · · · · · · · · · ·		
17 18 19 20 21 22 23	B-6A B-6A B-6A B-6A B-6A B-6A	A B2tca 2B3tca 2Cca 2C1csca 2C2cs base of		1.31 10.45 1.46 1.61 3.40 3.27	1.02 8.27 0.47 0.39 0	0.31 3.44 0.25 0.07 0.23 0.17	0.019 0.068 0.025 0.044 0.012 0.059	-1.35 2.30 7.34 2.79	-0.0008 0.0444 0.0047 0.106 0.465 0.0205	0.18 1.34 0.20 0.40 1.17 0.76	0.18 1.68 0.19 0.42 1.17 0.82	0.19 1.93 0.24 0.40 0.63 0.53	0 0 0.070 0.12 0.19 0	0 0 0 0 0 0
24 25 26 27 28 29 30 31	B-14 B-14 B-14 B-14 B-14 B-14 B-14	All Al2 Bl B2ltca B22tca 2B3tca 2Cca 2Cnca	4 8 17 28 44 70 123 193	2.13 1.43 3.32 3.59 4.71 2.87 4.03 5.11	1.09 1.06 2.93 4.24 5.37 0.32 0	0.06 0.40 1.09 1.12 1.66 0.15 0	0.023 0.026 0.068 0.056 0.070 0.037 0.060 0.076	-1.88 -1.79 -5.13 -6.18 -8.33 -0.84 2.38 1.56	-0.0005 0.0007 0.0049 -0.0008 -0.0011 0.0037 0.0935 0.218	0.083 0.18 0.51 0.56 0.65 0.16 0.32 0.85	0.18 0.25 0.82 0.68 0.99 0.17 0.34 0.71	0.12 0.18 0.49 0.76 0.99 0.22 0.36 0.61	0 0 0 0 0.20 0.036 0	0 0 0 0 0 0.13 0.13

Supplementary Table 7, Part 1. Horizon weights of sand, silt, clay, carbon, CaCO3, gypsum, and clay minerals--Continued

			Basal				W	eight (	g/cm <sup>2</sup> /hor	izon col	umn)			
No.	Sampl numbe	e r Horizon	depth (cm)	Sand	Silt	l Clay	<sup>l</sup> Carbon	CaCO <sub>3</sub>	3 <sup>1</sup> Gypsum <sup>1</sup>	Kaoli- nite	Mica	Smec- tite	Paly- gorskite	Unknown emineral
							Fan 4,	315 ka						
32 33 34 35 36 37 38	B-16 B-16 B-16 B-16 B-16 B-16 B-16	All Al2 Bt Btca Ccacs 2Clcsca 2C2cs	4 9 12 28 75 137 +237	0.48 0.48 0.18 0.95 2.09 0.81 2.72		0.18 0.12 0.01 0 1.78 -1.08 2.05	0.033 0.017 0.009 0.049 0.043 0.002 0.017	-1.57 -0.75 -4.03		0.12 0.15 0.065 0.35 0.80 0.54 2.13	0.25 0.28 0.11 0.56 0.87 0.31	0.13 0.14 0.11 0.56 1.60 0.51 2.32	0 0.034 0.18 0.27 0.093 0.38	0 0 0 0 0.27 0.093 0.38
							Fan 5,	410 ka						
39 40 41 42 43	B-7 B-7 B-7 B-7 B-7	A C1csca 2C2csca 2C3csca 2C4cs	6 30 50 83 +198			  		  		  	   		   	  
44 45 46 47 48 49 50	B-17 B-17 B-17 B-17 B-17 B-17	A11 A12 Btca 2Clcsca 2C2csca 2C3cs 2C4cs	4 10 19 40 89 179 249	1.03 1.28 0.84 0.87 0.027 1.16 2.21	1.84 1.72 1.46 0.11 -0.37 0.38	0.54 0.86 0.63 0.36 0.03 0.51	0.029 0.040 0.047 0.098 0.024 0.036 0.032	-1.48 -2.15 -2.52 -2.45 3.63 -2.97 -1.56	7.39 16.7	0.27 0.32 0.21 0.087 0.008 0.19 0.17	0.35 0.54 0.33 0.24 0.007 0.26 0.63	0.27 0.35 0.28 0.13 0.008 0.26 0.46	0 0.18 0.11 0.087 0.003 0.09 0.17	0 0 0.40 0.43 0.003 0.08 0.27
							Fan 6,	505 ka						
51 52 53 54 55	B-3 B-3 B-3 B-3 B-3	A Btca Clcsca 2C2cs 2C3cs	4 24 42 62 +67				  				  		  	  
56 57 58 59 60 61	B-18a B-18a B-18a B-18a B-18a B-18a	All Al2 Btca 2Clcsca 2C2cs 2C3cs	2 5 10 37 46 +86	0.44 0.58 0.48 1.79 0.70 4.11	0.47 0.62 -0.21 0.32 0		0.008 0.016 0.018 0.033 0.007 0.038	-0.39 -0.87 -1.79 -1.15 0.51 -5.54	0.0119 0.0031 0.0023 15.5 1.99 5.80	0.084 0.12 0.21 0.079 0.19 0.71	0.15 0.23 0.17 0.13 0.13 0.63	0.090 0.16 0.19 0.034 0.17 1.13	0 0.081 0.053 0.079 0.087 0.23	0 0.053 0.82 0.080 0.30
62 63 64 65 66	B-18b B-18b B-18b B-18b B-18b B-18b	A Btca 2Clcsca 2C2cs 2C3cs 2C4cs	6 22 51 75 130 +255	1.17 3.20 2.46 6.97 3.31 13.3	1.30 2.54 1.20 5.25 0.29 3.00	0.30 1.88 2.13 4.59 0.22 1.93	0.021 0.099 0.091 0.052 0.114 0.150	-1.29 -5.50 -3.37 -2.74 -4.27 -8.51		0.25* 0.98 0.82 0.79 0.47 2.37	0.45* 1.16 0.75 0.45 0.47 2.17	0.27* 1.05 1.33 2.48 0.67 3.65	0* 0.43 0.51 0.23 0.16 0.69	0* 0 0 1.69 0.27 0.99
							Fan 7,	585 ka						
68 69 70 71 72	B-4 B-4 B-4 B-4 B-4	A B21tcacs B22tcsca 2C1csca 2C2cs	12		  		  		  					
73 74 75 76 77 78 79	B-19 B-19 B-19 B-19 B-19 B-19	A B21tca B22tcsca 2C1csca 2C2cs 2C3cs 3C4cs	6 14 25 40 85 137 +262	0.94 -0.32 0.69 0.92 0.64 6.43 9.25	-0.03 -0.93 1.00 2.81	-0.03	0.011 0.023 0.084	-2.01 -0.95 -3.72 -6.19 -5.42	0.509 7.13 13.9	0.31 -0.09 0.063 0.23 0.14 1.48 1.55	0.25 -0.08 0.053 0.23 0.11 0.84 1.23	0.38 -0.21 0.17 0.52 0.23 2.06 2.72	0.082 -0.057 0.037 0.076 0.033 0.26 0.39	0 -0.033 0.013 0.22 0.30 0.63 0.58

 $<sup>^{1}</sup>$  Pedogenic weights, estimated parent material amount subtracted.  $\star$  Estimated by comparison to A horizons of soil B-18a.

# Supplementary Table 7, Part 2. Horizon weights of major oxides plus zirconium

[--, not measured. Analyst: M. C. Reheis, U.S. Geological Survey]

Methods

Horizon weights for element oxides are calculated using the first term in equation 2 given in Supplementary Table 7, Part 1.

	Sample		Basal depth	Weig	iht (g/c	m <sup>2</sup> /hori	zon co	lumn) of	oxide	in less	-than-2m	m fracti	on
No.		Horizon	(cm)	SiO <sub>2</sub>	A1 <sub>2</sub> 0 <sub>3</sub>	Fe <sub>2</sub> 0 <sub>3</sub>	Mg0	Ca0	Na <sub>2</sub> 0	K <sub>2</sub> 0	Ti02	Mn0	Zr0 <sub>2</sub>
						Fan 1	, 5 ka						
1 2 3 4	B-21 B-21 B-21 B-21	A 2Btca 2Cca 2Cn	6 13 33 73	1.8 1.3 1.4 1.8	0.24 0.21 0.17 0.20	0.089 0.076 0.072 0.10	0.10 0.11 0.53 0.72	0.30 0.44 1.9 2.5	0.033 0.022 0.024 0.033	0.054 0.051 0.054 0.071	0.013 0.010 0.0077 0.0084	0.0017 0.0013 0.0020 0.0028	0.001 0.001 0.000 0.000
						Fan 2,	65 ka						
5 6 7 8 9 10	B-5 B-5 B-5 B-5 B-5	A B2tca B3tca 2C1ca 2C2ca 2C3ca	3 11 20 31 46 +71	   									
11 12 13 14 15 16	B-15 B-15 B-15 B-15 B-15 B-15	A B1 B2tca 2B3tca 2Cca 2Cn	4 12 24 45 65 +165	2.0 6.7 8.9 2.3 2.4	0.27 0.98 1.4 0.28 0.29 1.4	0.10 0.36 0.50 0.12 0.13 0.64	0.10 0.30 0.49 0.39 0.69 3.4	0.26 0.75 1.4 1.6 2.2	0.034 0.12 0.16 0.035 0.042 0.21	0.060 0.22 0.28 0.068 0.081 0.40	0.016 0.053 0.069 0.016 0.015 0.071	0.0019 0.0060 0.0074 0.0027 0.0034 0.019	0.002 0.006 0.008 0.001 0.001 0.005
						Fan 3,	100 ka						
17 18 19 20 21 22 23	B-6A B-6A B-6A B-6A B-6A B-6A	A B2tca 2B3tca 2Cca 2C1csca 2C2cs base of	4 29 46 88 181 +242 fan 3										
24 25 26 27 28 29 30 31	B-14 B-14 B-14 B-14 B-14 B-14 B-14	A11 A12 B1 B21tca B22tca 2B3tca 2Cca 2Cnca	4 8 17 28 44 70 123 193	3.3 3.1 7.0 8.5 11. 3.5 5.6 8.5	0.40 0.43 1.0 1.3 1.7 0.43 0.72	0.14 0.15 0.37 0.46 0.61 0.18 0.30 0.55	0.17 0.15 0.31 0.37 0.47 0.54 2.0 2.4	0.45 0.34 0.75 0.95 1.1 2.0 5.9	0.064 0.055 0.13 0.17 0.21 0.070 0.13 0.18	0.090 0.095 0.24 0.28 0.36 0.090 0.19 0.28	0.022 0.022 0.053 0.068 0.091 0.022 0.033 0.059	0.0031 0.0029 0.0058 0.0066 0.0092 0.0037 0.0082 0.023	0.003 0.005 0.006 0.008 0.003 0.003
						Fan 4,	315 ka						
32 33 34 35 36 37 38	B-16 B-16 B-16 B-16 B-16 B-16	A11 A12 Bt Btca Ccacs 2C1csca 2C2cs	4 9 12 28 75 137 +237	1.9 1.8 0.70 3.7 6.2 3.9	0.28 0.29 0.10 0.55 0.83 0.50 2.3	0.10 0.10 0.039 0.21 0.38 0.21 0.94	0.14 0.12 0.081 0.43 2.9 1.4	0.50 0.49 0.43 2.3 16. 5.6 5.9	0.034 0.030 0.012 0.062 0.11 0.082 0.25	0.064 0.062 0.020 0.11 0.17 0.98 0.40	0.015 0.015 0.0055 0.029 0.038 0.025 0.13	0.0018 0.0016 0.0005 0.0027 0.0085 0.0035 0.029	0.001 0.001 0.002 0.003 0.003 0.002

	Sample		Basal depth	Weight (g/cm <sup>2</sup> /horizon column) of oxide in less-than-2mm fraction									
No.		Hori zon	(cm)	Si0 <sub>2</sub>	A1 <sub>2</sub> 0 <sub>3</sub>	Fe <sub>2</sub> 0 <sub>3</sub>	Mg0	Ca0	Na <sub>2</sub> 0	K <sub>2</sub> 0	Ti02	Mn0	Zr0 <sub>2</sub>
						Fan 5,	410 ka						
39	B-7	A	6										
40	B-7	C1csca	30										
41	B-7	2C2csca	50										
42	B-7	2C3csca	83										
43	B-7	2C4cs	+198										
44	B-17	A11	4	3.2	0.51	0.18	0.21	0.63	0.06	0.12	0.027	0.0033	0.002
45	B-17	A12	10	3.8	0.62	0.21	0.28	1.0	0.069	0.14	0.030	0.0033	0.002
46	B-17	Btca	19	2.9	0.48	0.17	0.48	1.8	0.047	0.096	0.022	0.0025	0.001
47	B-17	2Clcsca	40	1.6	0.21	0.082	0.41	5.3	0.034	0.043	0.0089	0.0011	0.000
48	B-17	2C2csca	89	4.2	0.54	0.29	2.6	7.3	0.061	0.13	0.032	0.0064	0.002
49	B-17	203cs	179	10.	1.5	0.64	3.6	11.	0.17	0.32	0.081	0.014	0.006
50	B-17	2C4cs	249	7.1	0.96	0.39	2.3	7.5	0.11	0.21	0.050	0.0082	0.004
				, • •				, • 5					
						Fan 6,	505 ka						
51	B-3	Α	4										
52	B-3	Btca	24										
53	B-3	Closca	42										
54	B-3	2C2cs	62										
55	B-3	203cs	+67										
56	B-18A	A11	2	1.3	0.17	0.066	0.079	0.32	0.021	0.039	0.0092	0.0017	0.001
57	B-18A	A12	5	1.9	0.28	0.11	0.14	0.59	0.030	0.058	0.015	0.0026	0.001
58	B-18A	Btca	10										
59	B-18A	2Clcsca	37										
60	B-18A	202cs	46										
61	B-18A	2C3cs	+86										
62	B-18B	Α	6	3.3	0.48	0.18	0.23	0.94	0.55	0.10	0.025	0.0045	0.002
63	B-18B	Btca	22	8.5	1.3	0.49	0.59	1.7	0.13	0.24	0.064	0.0096	0.005
64	B-18B	2Clcsca	51	7.0	0.85	0.37	0.95	6.9	0.11	0.18	0.050	0.0091	0.005
65	B-18B	2C2cs	75	8.2	1.4	0.53	2.5	7.5	0.16	0.27	0.059	0.0085	0.004
66	B-18B	2C3cs	130	14.	1.6	1.4	2.9	9.9	0.23	0.37	0.088	0.085	0.008
67	B-18B	2C4cs	+255	32.	3.8	2.0	4.1	16.	0.51	0.95	0.21	0.074	0.020
						Fan 7,	585 ka						
68	B-4	A	6										
69	B-4	R21tcacs	` 12										
70	B-4	B22tcsca	30										
71	B-4	2C1csca	55										
72	B-4	2C2cs	+75										
73	B-19	A	6	3.7	0.54	0.22	0.20	0.68	0.053	0.11	0.028	0.0068	0.002
74	B-19	B21tca	14	2.3	0.34	0.17	0.13	0.65	0.033	0.065	0.028	0.0068	0.002
7 <del>5</del>	B-19	B22tcsca	25	1.2	0.10	0.041	0.059	0.46	0.0091		0.017	0.0013	0.001
76	B-19	2C1csca	40	2.6	0.10	0.15	0.53	3.1	0.033	0.021	0.0036	0.0013	0.001
70 77	B-19	202cs	85	11.	1.5	0.69			0.033	0.081	0.017	0.0084	0.001
77 78	B-19 B-19	2020s	137	17.	2.1		1.5	6.5 6.7	0.12		0.078	0.027	0.010
70 79	B-19	3C4cs	262	23.	3.0	1.1 1.7	2.1 1.9	6.7 9.5	0.28	0.45 0.70	0.17	0.047	0.010
, ,	ローエフ	UUTU3	LU L	LJ.	J.U	1.0/	1.7	7.0	0.40	0.70	0.17	0.007	0.013

