## NOAA Technical Report NESDIS 124

Calibration of the Advanced Microwave Sounding Unit-A for NOAA-N'



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National Oceanic and Atmospheric Administration
National Environmental Satellite, Data, and Information Service

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Tsan Mo NOAA/NESDIS/STAR 5200 Auth Road Camp Springs, MD 20746

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#### U.S. DEPARTMENT OF COMMERCE

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## CALIBRATION OF THE ADVANCED MICROWAVE SOUNDING UNIT-A FOR NOAA-N'

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#### **ABSTRACT**

After successful launches of the Advanced Microwave Sounding Unit-A (AMSU-A) onboard the NOAA-K, -L, -M, -N, and the European Organization for the Exploitation of Meteorological Satellites (EUMETSAT) METeorological OPerational (METOP) satellite in 1998, 2000, 2002, 2005, and 2006, receptively, the sixth AMSU-A will be launched onboard the NOAA-N' in February 2009. This report concerns the analysis and evaluation of the thermal-vacuum chamber calibration data from the AMSU-A1 SN 107 and AMSU-A2 SN 109 flight models (FM) for NOAA-N'. The pre-launch calibration data were analyzed to evaluate the instrument performance, including calibration accuracy, nonlinearity, and temperature sensitivity. Great effort was taken to understand the instrument's radiometric performance as a function of instrument temperature. The calibration data provide a base for derivation of the calibration parameters input data sets (CPIDS) which will be incorporated into the operational calibration algorithms for producing the NOAA AMSU-A 1B data sets.

The nonlinearity parameters, which will be used to account the nonlinear contributions from an imperfect square-law detector, were determined from this data analysis. The existence and magnitude of nonlinearity in each channel were established and simulated with a quadratic formula for modeling the nonlinear contribution that was developed in the analysis of the NOAA-KLMN AMSU-A pre-launch calibration data. The model was characterized by a single parameter u, values of which were obtained for each channel via least-squares fit to the data. Quadratic corrections which would be expected from the on-orbit data after the launch of AMSU-A into space were simulated. In these simulations, the cosmic background radiance corresponding to a cold space temperature 2.73K was adopted as one of the two reference points of calibration. The largest simulated nonlinear correction is about 2 K. Experience learned from examining the previous AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps process these pre-launch calibration data. The calibration information presented in this report will be useful for post launch on-orbit verification of the AMSU-A instrument performance.

#### 1. INTRODUCTION

On 13 May 1998, the NOAA-K, which is designated NOAA-15 after the launch, was successfully launched into a circular, near-polar orbit with an altitude of 833-km above the Earth and an inclination angle of 98.7° to the Equator. NOAA-15 carries the first of a series of new microwave total-power radiometers, the Advanced Microwave Sounding Units (AMSU-A and AMSU-B), which provide 20 channels for atmospheric temperature and moisture soundings. NOAA-L/16, which was launched on September 28, 2000, carries the second AMSU-A. The NOAA-M/17 AMSU-A was successfully launched on June 24, 2002. NOAA-17 is the first NOAA satellite that has been launched into an orbit to cross the Equator at 1000 and 2200 local time. NOAA-N/18 was launched on May 20, 2005 and carried the first Microwave Humidity Sounder (MHS) which replaces AMSU-B for humidity sounding.

The METOP-A, the first of three satellites of the EUMETSAT Polar System (EPS), was launched on October 19, 2006, from Baikonur Cosmodrome in Kazakhstan, carries on board a set of NOAA provided instruments, including AMSU-A. The METOP series is part of the Initial Joint Polar-Orbiting Operational Satellite System (IJPS) constellation, along with the NOAA-N and -N' satellites. Under the IJPS, NOAA and EUMETSAT have agreed to provide instruments for each other's satellites.

The AMSU-A is divided into two physically separate modules, each of which operates and interfaces with the spacecraft independently. The AMSU-A1 module uses two independent antennaradiometer systems (A1-1 and A1-2) to provide 12 channels in the range 50.3 to 57.3 GHz oxygen band for retrieving the atmospheric temperatures from the Earth's surface to about 50 kilometers (or 1 mb), and another channel at 89 GHz. The AMSU-A2, which has 2 channels at 23.8 and 31.4 GHz, are used to identify precipitation and correct the effect of surface emissivity, atmospheric liquid water, and water vapor on temperature sounding. These window channels are also used to derive rain rate, sea ice concentration, and snow cover.

Table 1 lists some of the NOAA-N' AMSU-A main channel characteristics, including channel frequency, number of bands, 3-dB RF band width, radiometric temperature sensitivity (or NE $\Delta$ T), antenna beam efficiency, polarization, and field of view (FOV) of angular beam width for each channel. More detailed information on the AMSU-A radiometers is reported elsewhere [1]. Each of the AMSU-A antenna systems has a nominal FOV of 3.3° at the half-power points and covers a

cross-track scan of  $\pm 48^{\circ}20'$  (to beam centers) from the nadir direction with 30 Earth FOVs per scan line. Beam positions 1 and 30 are the outermost scan positions of the Earth views, while beam positions 15 and 16 (at  $\pm 1.67^{\circ}$  from nadir) straddle the nadir. Views of cold space and a black-body target at the end of a scan provide onboard calibration once every 8 seconds.

EUMETSAT provides the MHS for humidity sounding. It has two channels at 89 and 157 GHz, respectively, two channels around the 183 GHz water vapor line, and another channel at 190 GHz. Detailed description of the MHS radiometric performance has been given elsewhere [2].

All of the AMSU-A flight models were tested and calibrated in a thermal-vacuum (TV) chamber by the contractor. These pre-launch TV calibration data were evaluated and analyzed to derive the calibration parameters input data set (CPIDS) which is used in the NOAA operational calibration algorithm to produce the AMSU-A 1B data sets. Particularly, there is a small nonlinearity (of the order of 1 K), which cannot be evaluated by the two-point calibration but is determined from the pre-launch calibration data.

In this study, the TV test data from the AMSU-A1 SN 107 and AMSU-A2 SN 109 are analyzed and evaluated. The same procedure developed for NOAA-K AMSU-A calibration [1] is closely followed. Experience gained from examining the NOAA-15, -16, -17, and -18 AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps improve the analysis of the pre-launch calibration data.

In the following sections, the results from the data analysis are presented. Instrument performances evaluated in this analysis include calibration accuracy, nonlinearity, and radiometric temperature sensitivity (or NEΔT, the noise-equivalent temperature). Section 1 gives an introduction and section 2 presents a brief description of the TV chamber test data. The calibration algorithm is described in section 3. The results are presented in section 4. Conclusion and discussion are given in section 5. Tables of CPIDS for NOAA-N' AMSU-A are given in Appendices A. A brief description of the NOAA Level 1B data is given in Appendix B.

#### 2. DESCRIPTION OF CALIBRATION DATA

Aerojet (now Northrop Grumman), the AMSU-A contractor, took the calibration data in a TV test chamber using the full scan mode. In this mode, AMSU-A scans through 30 Earth field of views (FOVs), the cold target, and the internal warm blackbody target once every 8 seconds. It takes one sample at each Earth FOV and two samples at the cold and warm targets, respectively. Since the scene calibration target was fixed at FOV 6 (31° 40' from nadir), only radiometric counts from FOV 6 are used for calibration. Each antenna system looks at its own individual cold, warm, and scene targets. A series of Platinum Resistance Thermometers (PRTs) were used to monitor the temperatures of these calibration targets. The numbers of PRTs used to measure the physical temperatures of the scene, cold, and warm calibration targets in each antenna system are given in Table 2 which also lists the channels provided by each AMSU-A antenna system. Channels 9-14 have both primary and secondary phase locked loop oscillators (i.e. PLLO #1 and PLLO #2, respectively) built-in. The PLLO #2 will be used for backup if PLLO #1 fails in operation. Invar high-Q cavity stabilized local oscillators [3] are used in other channels.

The physical temperatures of scene and cold targets measured by individual PRTs were provided in Kelvin (K) on Aerojet's data packages. However, the data from the PRTs monitoring the warm blackbody targets are given in counts, which are proportional to the blackbody temperatures. One should note that the scene and cold targets used in the TV chamber will not be carried into space. The PRT counts from the warm blackbody targets must be converted to PRT temperatures. The normal approach of deriving the PRT temperatures from counts is a two-step process: (1) the resistance of each PRT in ohms is computed by a count-to-resistance look-up table provided by the manufacturer. In this study, we used a polynomial representation of the count-to-resistance relationship provided by Aerojet; and (2) the individual PRT temperature in degrees Celsius is obtained from an analytic PRT equation [4], which is described in Appendix A. However, the two steps can be compressed to a single step with negligible errors. This single step process, which has been used in the NOAA-KLMN and METOP-A calibrations, computes the PRT temperatures directly from the PRT counts, using a cubic polynomial

$$T_{Wk} = \sum_{j=0}^{3} f_{kj} C_k^j \tag{1}$$

where  $T_{Wk}$  and  $C_k$  represent the temperature and count of the PRT k. The polynomial coefficients,  $f_{kj}$ , are derived for each PRT in this study. Equation (1) also applies to 47 other housekeeping

Table 1. NOAA-N' AMSU-A1 SN 107 and AMSU-A2 SN 109 Channel Characteristics.

Channel		Frequency Hz)	No. of RF Bandwidth		NEΔT (K)		Beam <sup>b</sup>	Polarization (NA DVD)	FOV <sup>c</sup>
Number	Specification	Measured <sup>a</sup>	Bands	(MHz)	Spec.	Measured	Efficiency	( NADIR)	(deg.)
1	23800	23800.47	1	251.08	0.30	0.172	97%	V	3.49
2	31400	31401.37	1	160.08	0.30	0.170	98%	V	3.50
3	50300	50300.41	1	160.24	0.40	0.213	97%	$\mathbf{V}$	3.46
4	52800	52800.74	1	380.40	0.25	0.136	95%	V	3.39
5	53596 ± 115	53596.41 ± 115	2	167.86   167.86	0.25	0.160	96%	Н	3.39
6	54400	54400.05	1	379.94	0.25	0.136	97%	H	3.40
7	54940	54939.30	1	380.30	0.25	0.1650	97%	$\mathbf{V}$	3.41
8	55500	55500.69	1	310.12	0.25	0.172	95%	Н	3.37
9	fo = 57290.344	fo = 57290.332	1	310.12	0.25	0.178	96%	H	3.38
10	fo ± 217	fo ± 217	2	75.98   75.98	0.40	0.225		Н	
11	fo $\pm 322.2 \pm 48$	$fo \pm 322.2 \pm 48$	4	34.90 /35.29  35.29 /34.90	0.40	0.240		Н	
12	fo $\pm 322.2 \pm 22$	fo $\pm 322.2 \pm 22$	4	15.44 / 15.47   15.47 / 15.44	0.60	0.349		н	
13	fo $\pm$ 322.2 $\pm$ 10	fo $\pm$ 322.2 $\pm$ 10	4	7.82 / 7.85   7.85 / 7.82	0.80	0.506		Н	
14	fo $\pm$ 322.2 $\pm$ 4.5	$fo \pm 322.2 \pm 4.5$	4	2.95 / 2.96   2.96 / 2.95	1.20	0.829		Н	
15	89000	89010.00	1	1990.94	0.50	0.122	98%	V	3.26

<sup>&</sup>lt;sup>a</sup>At a temperature of 20° C for channels 1, 2, and 15; at 22° C for other channels. <sup>b</sup>Measured.

<sup>&</sup>lt;sup>c</sup>Specification is  $3.3^{\circ} \pm 10\%$ .

Table 2. Number of PRTs in each calibration target and channels provided by individual antenna systems. The last row gives the number of scans collected in the calibration for each system

T	Numb	per of PRTs in	Remarks	
Items	A2	A1-2	A1-1	
Warm Target	7	5	5	
Cold Target	11	7	7	
Scene Target	11	7	7	One bad PRT in A1-2
				Three bad PRTs in A1-1
Channels	1 & 2	3, 4, 5 & 8	6, 7, & 9-15	
PLLO#2	-	-	Ch. 9-14	Backup PLLO
N(Scan No.)	120	400 to 900	400 to 900	See Note

**Note**: For A1-1 & A1-2: N=400, 400, 550, 725, 725, and 900 when the scene target temperature at  $T_s$ =84, 130, 180, 230, 280, and 330K, respectively.

temperature sensors, such as the mixers, the IF amplifiers and the local oscillators. These  $f_{kj}$  coefficients for all PRTs and housekeeping sensors in NOAA-N' AMSU-A1 SN 107 and AMSU-A2 SN 109 are listed in Appendix A (Tables A-1 and A-2, respectively).

The mean internal blackbody temperature,  $T_W$ , is calculated from the individual PRT temperatures,

$$T_{W} = \frac{\sum_{k=1}^{m} W_{k} T_{Wk}}{\sum_{k=1}^{m} W_{k}} + \Delta T_{W}$$
 (2)

where m represents the number of PRTs for each antenna system (as listed in Table 2) and  $W_k$  is a weight assigned to each PRT k. The quantity  $\Delta T_W$  is a warm load correction factor, which is derived for each channel from the TV test data at three instrument temperatures (low, nominal, and high). The procedure for determining the  $\Delta T_W$  values has been described elsewhere [5]. For each AMSU-A antenna system, a special set of calibration data was taken to determine the  $\Delta T_W$  values, which are given in Appendix A (Table A-3). The  $W_k$  value, which equals 1 (0) if the PRT k is determined good (bad) before launch. Similarly, the mean temperatures of cold targets are defined in the same

way as in Equation (2), except without the term  $\Delta T_W$ .

The TV calibration data were taken at three instrument temperatures and the scene target was cycled through six temperatures 84, 130, 180, 230, 280, and 330K, respectively, at each instrument temperature. At each of the scene target temperatures, calibration data were acquired for enough number of scans to assure an effective temperature sensitivity less than 0.03K. The number of scans required depends upon the expected NE $\Delta$ T of channel 14 (which has the largest NE $\Delta$ T) and is therefore a function of the scene target temperature. Actual numbers of scans taken at individual temperatures are given in Table 2. The uncertainties in knowledge of brightness temperatures and measurement errors were obtained by Aerojet [6, 7]. These are given Table 3.

Table 3. Uncertainty in brightness temperatures and measurement errors.

Source of Error	AMSU-A	2 :Ch. 1 and 2	AMSU-A1: Ch. 3 - 15		
Source of Error	Bias (K)	Random (k)	Bias (K)	Random (k)	
Warm Target	-0.050	±0.122	-0.050	±0.122	
Cold Target	0.024	±0.105	0.024	±0.091	
Scene Target	0.002	±0.101	0.002	±0.090	

Antenna beam widths at all channels were also measured and the values for NOAA-N' AMSU-A are listed in Table 1. Antenna beam efficiency was calculated at each channel frequency and is listed in Table 1.

#### 3. CALIBRATION ALGORITHM

In this study, calculations of radiometric measurements and variables related to the calibration process are all performed in radiance with dimension of mW/(m²-sr-cm⁻¹), but the final results are presented in temperatures. All instruments flown on NOAA satellites produce measurements in radiance. Conversion between brightness temperature and radiance was performed using the full Planck function, instead of the Rayleigh-Jeans approximation. This also eliminates any possible conversion inaccuracy that may occur, particularly in the region of the space cosmic back-ground temperature ~2.73K, where the Rayleigh-Jeans approximation breaks down.

For each scan, the blackbody radiometric counts  $C_W$  are the averages of two samples of the internal blackbody. Similarly, the space radiometric counts  $C_C$  are the average of two samples of the space target. To reduce noise in the calibrations, the  $C_X$  (where X=W or C) for each scan line were convoluted over several neighboring scan lines according to the weighting function [1]

$$\overline{C}_X = \frac{\sum_{i=-n}^n W_i C_X(t_i)}{\sum_{i=-n}^n W_i}$$
(3)

where  $t_i$  (when  $i \neq 0$ ) represents the time of the scan lines just before or after the current scan line and  $t_0$  is the time of the current scan line. One can write  $t_i = t_0 + i\Delta t$ , where  $\Delta t = 8$  seconds for AMSU-A. The 2n+1 values are equally distributed about the scan line to be calibrated. Following the NOAA-KLM operational preprocessor software, the value of n=3 is chosen for all AMSU-A antenna systems. A set of triangular weights of 1, 2, 3, 4, 3, 2, and 1 is chosen for the weight factor  $W_i$  that appears in Equation (3) for the seven scans with i = -3, -2, -1, 0, 1, 2, and 3, respectively.

The calibration algorithm [1], which takes into account the nonlinear contributions due to an imperfect square-law detector, converts scene counts to radiance  $R_S$  as follows,

$$R_{S} = R_{W} + \left(R_{W} - R_{C}\right) \left(\frac{C_{S} - \overline{C}_{W}}{\overline{C}_{W} - \overline{C}_{C}}\right) + Q \tag{4}$$

where  $R_W$  and  $R_C$  are the radiance computed from the PRT blackbody temperature  $T_W$  and the PRT cold target temperature  $T_C$ , respectively, using the Planck function. The  $C_S$  is the radiometric count from the Earth scene target. The averaged blackbody and space counts,  $\overline{C}_W$  and  $\overline{C}_C$ , are defined by Equation (3).

The quantity Q, which represents the nonlinear contribution, is given by [1]

$$Q = u \left( R_W - R_C \right)^2 \frac{\left( C_S - \overline{C}_W \right) \left( C_S - \overline{C}_C \right)}{\left( \overline{C}_W - \overline{C}_C \right)^2} \tag{5}$$

where u is a free parameter, values of which are determined at three instrument temperatures (low, nominal, and high). After launch, the u values at the actual on-orbit instrument temperatures will be interpolated from these three values. For channels 9-14 (AMSU-A1-1), two sets of the u parameters

are provided; one set is for the primary PLLO#1 and the other one for the redundant PLLO#2.

The quantity  $R_S$  in Equation (4) represents the radiometric scene radiance of individual channels. For users of NOAA Level 1B data, a simplified formula, which converts  $C_S$  directly into  $R_S$ , is presented in Appendix B. It should be noted that the ratios in Equations (4) and (5) will eliminate the effect of any linear variation in the radiometric counts on  $R_S$ . The channel gain, G, is defined as

$$G = \frac{\overline{C}_W - \overline{C}_C}{R_W - R_C} \tag{6}$$

The quantity G varies with instrument temperature, which is defined as the RF Shelf temperature for each AMSU-A antenna system. For a fixed instrument temperature, G is approximately constant. The first two terms in Equation (4) constitute a linear two-point calibration equation, if the quadratic term is negligible. Let  $R_{SL}$  denote these two terms,

$$R_{SL} = R_W + \left(R_W - R_C\right) \left(\frac{C_S - \overline{C}_W}{\overline{C}_W - \overline{C}_C}\right) \tag{7}$$

The linear scene radiances  $R_{\rm SL}$  are calculated using the TV chamber test data. The results are discussed in the following section.

#### 4. RESULTS

### 4.1 Calibration Accuracy

The calibration accuracy,  $\Delta R$ , is defined as the difference between the scene PRT radiance  $R_{sprt}$  and the radiometric scene radiance. It is calculated from the equation

$$\Delta R = \frac{1}{N} \left[ \sum_{i=1}^{N} \left( R_{sprt} - R_{SL} \right)_{i} \right]$$
 (8)

where N is the number of scans at a specific scene temperatures (N= 120 for AMSU-A2 channels but ranging from 400 to 900 for AMSU-A1 channels).

Equation (7) would be a good representation of the microwave radiometric scene radiance for a perfect square-law detector. Any deviation of the quantity  $R_{SL}$  from the measured scene PRT radiance  $R_{sprt}$  indicates the presence of either nonlinearity in the radiometer system or some other potential source of contamination in the calibration data.

Figure 1 shows the calculated calibration accuracies in temperature,  $\Delta T$  (corresponding to  $\Delta R$ ), versus the scene PRT temperature for channels 1 through 15. The AMSU-A specification requires  $\Delta T = 2.0$  K for channels 1, 2, and 15, and  $\Delta T = 1.5$  K for all other channels. The results in Figure 1 show that the  $\Delta T$  values for the A1-1 channels do not meet the specification.

## 4.2 Nonlinearity

The  $\Delta T$  patterns (Figure 1) show clearly the nonlinearity patterns which can be represented by the quadratic formula Q defined in Equation (5). The nonlinearity of a channel is normally defined as the residuals from a least-squares fit of its scene PRT radiance as a linear function of the radiometric radiance  $R_{SL}$  (Equation 7). The nonlinearity is defined as the differences (or residuals) between the  $R_{sprt}$  and the best-fit results from a linear equation in the form LinFit = a + b  $R_{SL}$  (where a and b are the intercept and slope). These residuals, which are defined as the measured nonlinearity Q (=  $R_{sprt}$  - LinFit), are shown in Figure 2. The largest (absolute) Q value on each curve is defined as the measured nonlinearity that can be compared to those as defined in the AMSU-A specification.

The results in Figure 2 show that the maximum (absolute) Q values are about 0.6K. The AMSU-A specification requires Q = 0.5K for channels 1, 2, and 15; and Q = 0.375K for other channels. All results at AMSU-A1-1 channels in Figure 2 do not meet the specifications. All of the curves, except channel 2, in Figure 2 have two roots, representing solutions of a quadratic equation, which can be written in the form

$$Q = u(R_S - R_1)(R_S - R_2) \tag{9}$$

where  $R_1$  and  $R_2$  represent the two roots. One can obtain a similar equation from Equation (5) by replacing the counts by its individual radiance, since the radiometric counts are proportional to radiance in a first-order approximation. The resultant equation represents a different straight line intersecting the same curve at  $R_W$  and  $R_C$ . Once the parameter u is determined, it can be used with any pair of roots to calculate Q. We applied Equation (9) to fit the quadratic curves in Figure 2 to obtain the u values with the two roots extracted from each plot. Table 4 gives the best-fit u values at three instrument temperatures for individual channels.

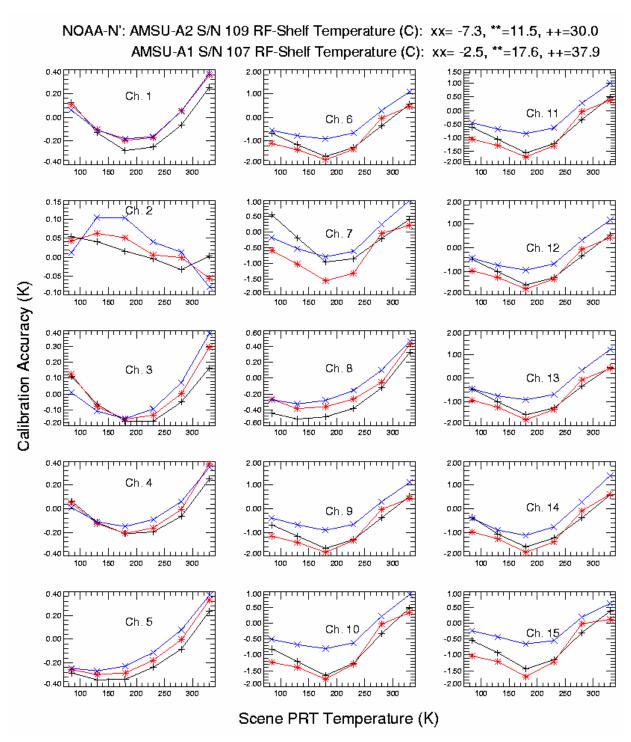


Figure 1. NOAA-N' AMSU-A: Calibration accuracies at three instrument temperatures.

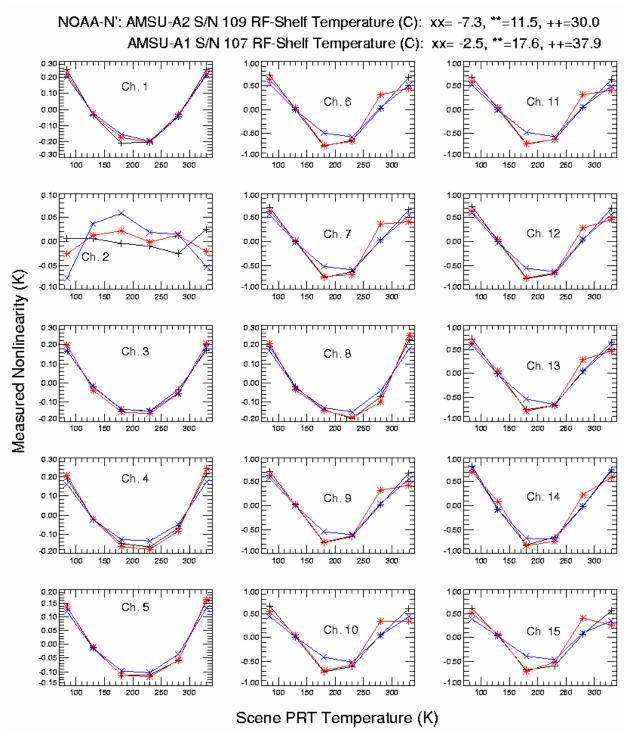


Figure 2. NOAA-N' AMSU-A: Pre-launch measured nonlinearities at three instrument temperatures.

## 4.3 Simulated Quadratic Corrections to On-Orbit Data

Once the u values are determined, Equation (9) can be used to simulate the quadratic contributions which are expected from on-orbit data. This can be accomplished by replacing the roots  $R_1$  and  $R_2$  with the radiance  $R_{2.73}$  (corresponding to cold space temperature 2.73K) and  $R_W$ , respectively. The simulated results are shown in Figure 3 for all channels at three instrument temperatures which are listed at the top of the plots. One should note that the instrument temperatures are different for the three AMSU-A antenna systems. The simulated quadratic contributions displayed in Figure 3 are larger than those shown in Figure 2. This is expected because the separation of the two reference calibration points is increased in the cases of simulations. Noticeable quadratic contributions appear in all channels, particularly the AMSU-A1-1 channels. The largest ones are approximately 2 K at several channels. It is important to note that the effect of instrument temperature on the quadratic contributions is nonlinear in general.

Similarly, Figure 4 shows the calculated results which are associated with the redundant PLLO#2 built into channels 9-14. The left column of Figure 4 displays the calibration accuracies,  $\Delta T$ , which are similar to Figure 1 and the middle column (corresponding to Figure 2) shows the residuals of the least-squares fit. The right column shows the simulated quadratic corrections, Q, which would occur in on-orbit data.

### 4.4 Temperature Sensitivity

The specifications of temperature sensitivity (or NE $\Delta$ T) for AMSU-A channels are listed in Table 1. It is defined as the minimum change of a scene brightness temperature that can be detected. In practice, it is calculated as the standard deviation of the radiometer output (in K), when an antenna system is viewing a scene target at 300K. The calculated NE $\Delta$ T values are shown in Figure 5 together with the AMSU-A specifications. These calculated NE $\Delta$ T values correspond to a scene temperature of 300K. All of the calculated NE $\Delta$ T values in Figure 5 are better than those given in the AMSU-A specification. Actually, all NE $\Delta$ T values measured at all temperatures are better than the specification. The calculated NE $\Delta$ T values at all scene target temperatures and three instrument temperatures are shown in Figure 6. In general, the NE $\Delta$ T value decreases as the instrument temperature becomes colder.

Table 4. NOAA-N' AMSU-A: Nonlinearity parameters u in dimension of (m<sup>2</sup>-sr-cm<sup>-1</sup>)/mW.

AMSU-A2 SN 109 channels:			AMSU-A1-	2 SN 107 c	hannels:		
Instrument Temp.(C)	Ch. 2	Ch. 2	Instrument Temp.(C)	Ch. 3	Ch. 4	Ch. 5	Ch. 8
-7.26	5.045037	-0.802970	-2.49	0.945506	0.807438	0.595301	0.762065
11.53	5.329711	-0.264490	17.62	1.066226	1.072694	0.694163	0.969826
30.04	5.953661	0.114607	37.92	0.913515	0.974329	0.685200	0.913523

AMSU-A1-1 SN 107 channels: PLLO #1

Instrumen									_
t	Ch. 6	Ch. 7	Ch. 9	Ch. 10	Ch. 11	Ch. 12	Ch. 13	Ch. 14	Ch. 15
Temp.(C)									
-2.52	2.394737	2.50256	2.33563	1.82341	2.06078	2.401584	2.43565	3.135952	0.608328
		3	3	3	4		4		
17.47	2.320407	2.19148	2.02943	1.70867	1.91176	2.127352	2.15587	2.563504	0.605858
		6	6	3	4		9		
37.88	3.247517	3.06423	2.91783	2.59963	2.68877	2.971152	2.85521	3.310253	0.982430
		8	3	6	5		6		

AMSU-A1-1 SN 107 channels: PLLO #2

Instrument Temp. (C)	Ch. 9	10	11	12	13	14
-2.67	2.559711	2.028035	2.361547	2.702107	2.663920	3.309312
17.47	1.984376	1.692755	1.870282	2.096040	2.053746	2.480539
37.60	2.541280	2.263376	2.355038	2.587566	2.527208	2.913724

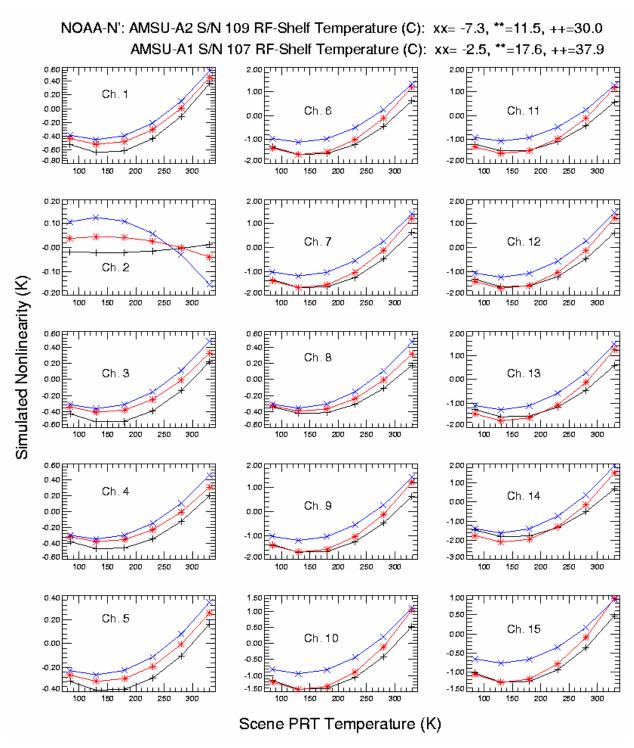


Figure 3. NOAA-N' AMSU-A: Simulated nonlinearities for on-orbit data.



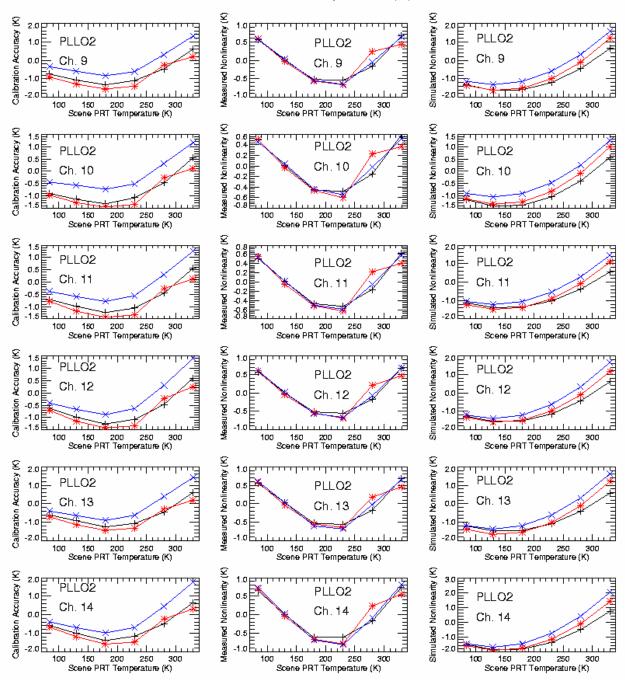


Figure 4. NOAA-N' AMSU-A: Calibration results with PLLO #2 at Channels 9-14, (a) calibration accuracy, (b) measured nonlinearity, and (c) simulated on-orbit nonlinearity.

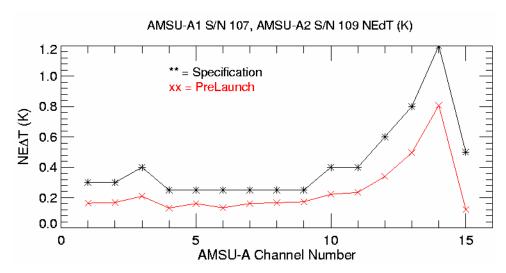


Figure 5. NOAA-N' AMSU-A: Comparison of the measured NEΔT values with specifications.

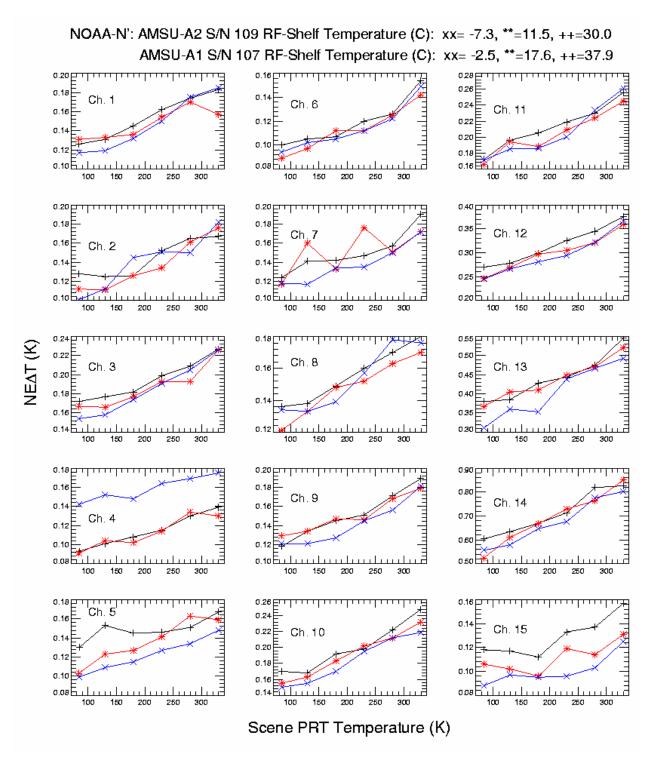


Figure 6. NOAA-N' AMSU-A: Measured NEΔT as a function of scene target temperatures at three instrument temperatures.

## 4.5 Radiometric Counts as a Function of Scene Target Temperature

Figure 7 shows the plots of TV scene radiometric counts versus the scene PRT temperatures for all channels. The radiometric counts increase linearly with scene PRT temperatures ranging from 84 to 330 K and good linear relationships exist between the radiometric counts and the scene PRT temperatures in the range 84 to 330 K. One can extrapolate these linear relationships to 0 K (corresponding to zero radiance) and obtain the intercepts for these plots. The intercept for each data point can also be computed from Equation (7) by setting  $R_{SL}$ = 0 and solve for  $C_{S}$ , which is denoted by  $C_{Sint}$  as

$$C_{\text{Sint}} = \overline{C}_W - G R_W \tag{10}$$

where G is the channel gain defined in Equation (6). The calculated  $C_{Sint}$  values at three instrument temperatures for each channel are shown in Figure 8. These are the mean values of all calculations performed with all available calibration data (from 120 to 900 scans as the scene temperature varies from 84 to 330 K). At each instrument temperature, the variation in the calculated  $C_{Sint}$  values is relatively small and all  $C_{Sint}$  values are positive. Figure 8 shows that the instrument temperature has a big impact on the magnitude of  $C_{Sint}$ , which increases as the instrument temperature decreases, except at channels 1, 2, and 8.

## 4.6 Channel gains

Values of the channel gains as defined in Equation (6) are also calculated from the TV calibration data. The gain values are converted into dimension of count/K. Figure 9 shows the calculated channel gains at three instrument temperatures, which are listed at the top of Figure 9. One should note that the abnormal behavior of channel 4 gain as a function of instrument temperature. Similarly, Figure 10 shows the calculated results of channel gain, radiometric counts, and intercept counts as a function of the scene PRT temperature when PLLO#2 was used.

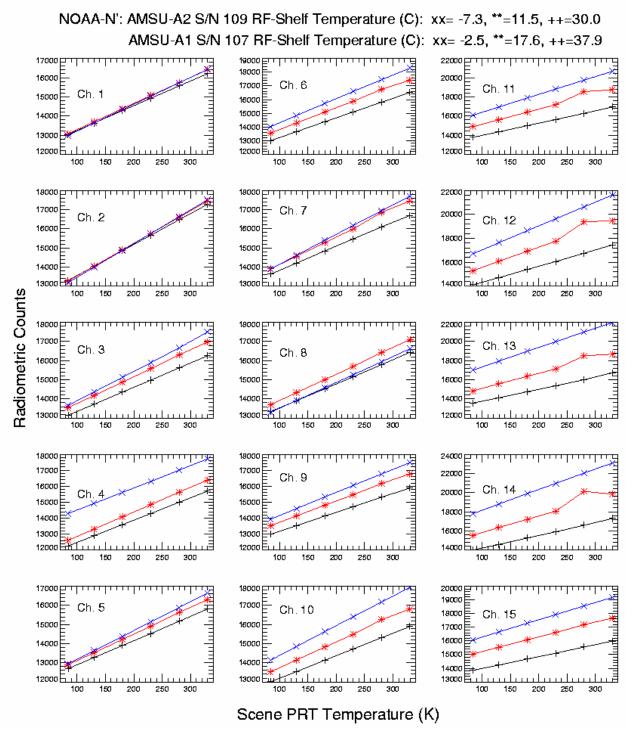


Figure 7. NOAA-N' AMSU-A: Radiometric counts versus Scene target temperature at three instrument temperatures.

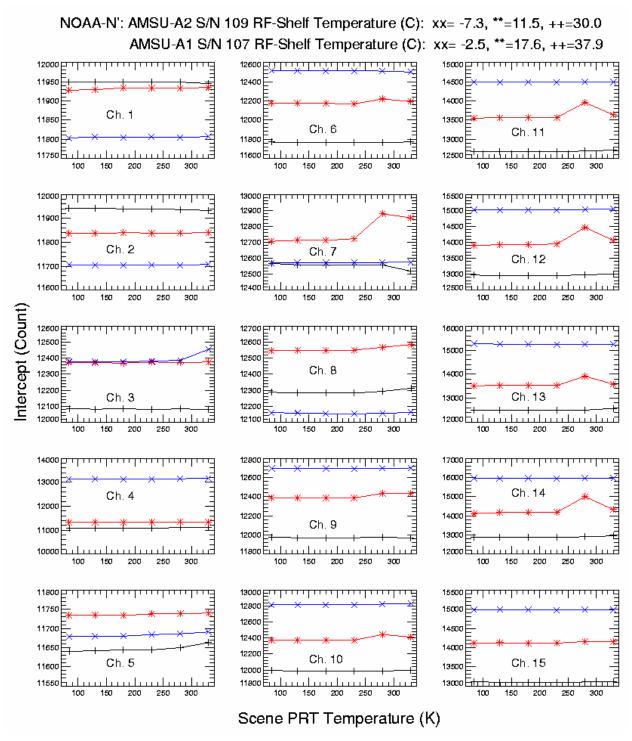


Figure 8. NOAA-N' AMSU-A: Intercept counts versus Scene target temperatures at three instrument temperatures.

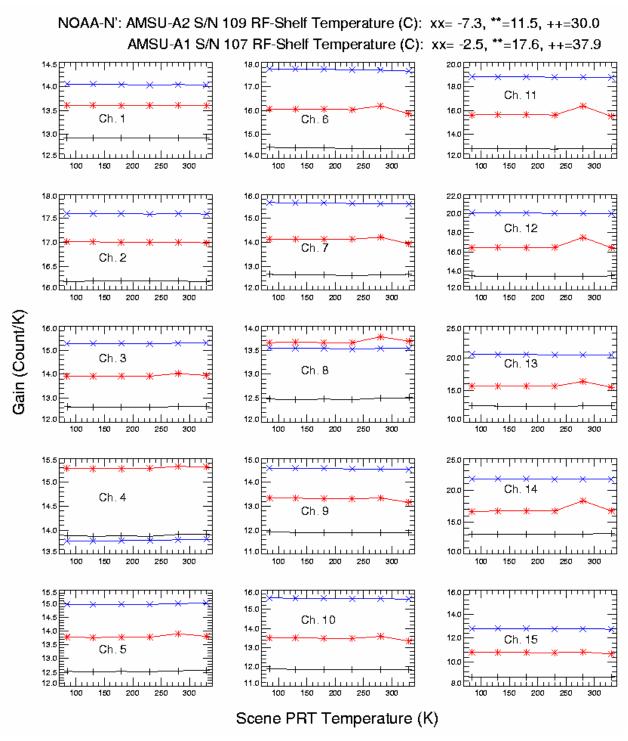


Figure 9. NOAA-N' AMSU-A: Channel gain versus scene target temperatures at three instrument temperatures.



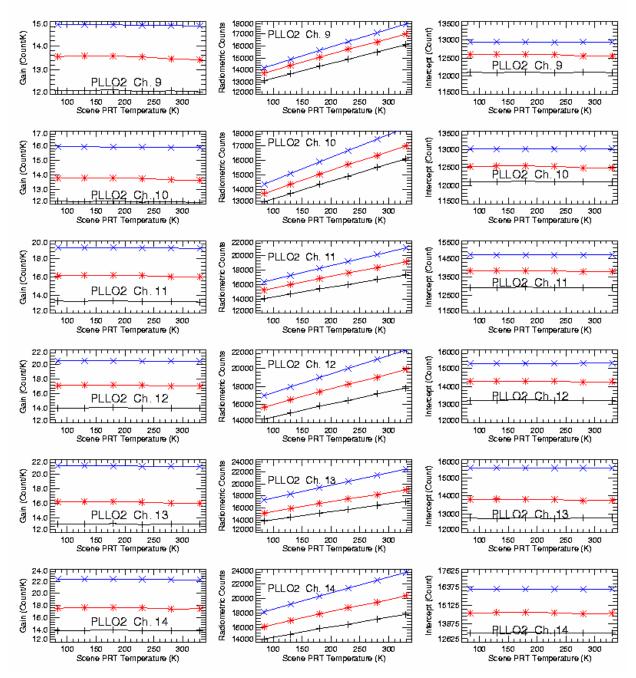


Figure 10. NOAA-N' AMSU-A: Calibration results as a function of scene target temperatures with PLLO #2 at three instrument temperatures. (a) gains, (b) radiometric counts, and (c) intercept counts

### 5. CONCLUSION AND DISCUSSION

The TV chamber calibration data for the NOAA-N' AMSU-A1 SN 107 and AMSU-A2 SN 109 were analyzed to derive the CPIDS which will be used in the NOAA operational calibration algorithms to produce the AMSU-A 1B data sets. The results show that the instruments meet the AMSU-A specification in calibration accuracy and temperature sensitivity, but the measured nonlinearities at the AMSU-A1-1 channels do not meet the AMSU-A specification.

The nonlinearity exists at all channels. A quadratic formula with a single parameter u is used to simulate these nonlinear contributions. The u values at three instrument temperatures were obtained from the pre-launch calibration data. Using the best-fit u values, the quadratic corrections which would be expected from the on-orbit data were simulated. In the simulations, the cold space radiance corresponding to 2.73K was adopted as one of the two reference calibration points (the other one is the internal blackbody temperature). The largest simulated nonlinear correction is about 2 K as shown in Figure 3.

Experience gained from examining the previous AMSU-A on-orbit data provides a better understanding of the AMSU-A performance in space and helps process the pre-launch calibration data. In general, the qualities of the calibration data are quite good. Particularly, the instrument temperatures were stabilized at pre-selected values with total variation less than  $\pm 0.5$ K during each test cycle. This renders the measured instrument nonlinearities more reliable. The results presented in this study confirm the good quality of the AMSU-A instruments for NOAA-N'. The calibration information presented in this report will be immensely useful for post launch on-orbit verification of the AMSU-A instrument performance.

Currently, AMSU-A SN 107 is under investigation of possible damage in an accident at Lockheed Martin. The analysis of the pre-launch calibration data will be redone if new calibration is required after any repair of damage.

### Acknowledgment

The contents of this manuscript are solely the opinions of the author and do not constitute a statement of policy, decision, or position on behalf of NOAA or the U. S. Government. Aerojet, (now Northrop Grumman) was the primary AMSU-A contractor for building the instruments.

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#### APPENDIX A

#### NOAA-N' CPIDS: COEFFICIENTS OF AMSU-A1 SN 107 AND AMSU-A2 SN 109

## A.1 Polynomial Coefficients for Converting PRT Counts into Temperatures

The two-step process for deriving the PRT temperatures from PRT counts  $C_k$  is briefly described here. First the count  $C_k$  from PRT k is converted into resistance  $r_k$  (in ohms) by a polynomial

$$r_k = \sum_{i=0}^{3} A_i C_k^i$$
 (A-1)

where the coefficients  $A_i$  for individual PRTs and temperature sensors were provided by Aerojet. Once the resistance  $r_k$  is known, then one can calculate the PRT temperature t (in Celsius) from the Callendar-Van Dusen equation [4], which is given by

$$\frac{r_t}{r_o} = 1 + \alpha \left[ t - \delta \left( \frac{t}{100} - 1 \right) \frac{t}{100} - \beta \left( \frac{t}{100} - 1 \right) \left( \frac{t}{100} \right)^3 \right]$$
 (A-2)

where:

 $r_t$  resistance (in ohms) at temperature t (°C) of the blackbody target

 $r_o$  resistance at t = 0°C (supplied by the manufacturer via Aerojet)

 $\beta$  0 for  $t > 0^{\circ}$ C, and 0.11 for  $t < 0^{\circ}$ 

 $\alpha$  and  $\delta$  are constants provided by the manufacturer via Aerojet.

Calculation shows that the error is negligible by setting  $\beta = 0$  in Equation (A-2). In this study, it is assumed  $\beta = 0$  in all cases, then Equation (A-2) is simplified into a quadratic equation in t. In such case, one can solve the quadratic equation for t in terms of  $r_t$ . Then the PRT temperature in degree Kelvin is obtained from t + 273.15. By this way, data sets of PRT temperatures versus counts for individual PRTs are computed. Then Equation (1) is applied to fit these data sets for obtaining the polynomial coefficients  $f_{kj}$  for individual PRTs and housekeeping sensors. These best-fit coefficients are listed in Tables A-1 and A-2, respectively, for AMSU-A1 SN 107 and AMSU-A2 SN 109. Test calculations show that these polynomials are very accurate in reproducing temperatures of all sensors with errors in order of 0.01 K.

Table A-1. NOAA-N' AMSU-A1 SN 107: Coefficients for converting PRT counts into temperatures.

PRT	$f_{k0}$	$f_{k1}$	$f_{k2}$	$f_{k3}$	Description
1	263.1118	1.742062E-03	3.497049E-09	1.447318E-14	Al-1 MOTOR
2	263.7931	1.733920E-03	3.858388E-09	4.108136E-15	A1-2 MOTOR
3	264.0873	1.772267E-03	1.969058E-09	3.782504E-14	Al-1 FEED HORN
4	263.7359	1.746087E-03	3.388829E-09	1.539593E-14	A1-2 FEED HORN
5	263.4033	1.731448E-03	4.052407E-09	6.281552E-15	Al-1 RFMUX
6	263.6212	1.736000E-03	3.594951E-09	1.264543E-14	Al-2 RFMUX
7	262.6824	1.728847E-03	4.019764E-09	5.986536E-15	Ch. 3 DRO
8	263.4037	1.735807E-03	3.885898E-09	8.667607E-15	CH. 4 DRO
9	262.7392	1.731672E-03	3.496970E-09	1.431707E-14	CH. 5 DRO
10	263.3888	1.740029E-03	3.520360E-09	1.251986E-14	CH. 6 DRO
11	263.6862	1.735802E-03	3.586685E-09	1.647744E-14	CH. 7 DRO
12	263.6868	1.744705E-03	3.305509E-09	1.818267E-14	CH. 8 DRO
13	263.6375	1.732038E-03	3.697993E-09	1.205412E-14	CH. 15 GDO
14	263.9067	1.734456E-03	3.429725E-09	1.544829E-14	CH.9 thru 14 PLO#2
15	263.3486	1.753624E-03	3.516755E-09	3.333381E-15	CH.9 thru 14 PLO#1
16	263.3486	1.753624E-03	3.516755E-09	3.333381E-15	NOT used (dummy)
17	263.2393	1.744402E-03	3.362316E-09	1.334880E-14	CH. 3 MIXERAF
18	263.6364	1.771411E-03	3.576661E-09	8.852903E-15	CH. 4 MIXER/IF
19	263.8027	1.735185E-03	3.816865E-09	4.611234E-15	CH. 5 MIXER/IF
20	263.9944	1.732857E-03	3.865460E-09	1.188194E-14	CH. 6 MIXER/IF
21	263.5726	1.734105E-03	4.104094E-09	4.784842E-15	CH. 7 MIXER/IF
22	263.1446	1.735692E-03	3.487495E-09	1.448047E-14	CH. 8 MIXERAF
23	263.3230	1.736698E-03	3.506417E-09	1.381465E-14	CH.9 thru 14 MIXERAF
24	263.8131	1.733783E-03	3.705516E-09	1.114634E-14	CH. 15 MIXER/IF
25	263.5378	1.746153E-03	3.526695E-09	1.533726E-14	CH.11 thru 14 IF AM
26	263.5392	1.735697E-03	3.888978E-09	1.279559E-14	CH. 9 IF AMP
27	263.1599	1.731394E-03	3.924602E-09	7.025278E-15	CH. 10 IF AMP
28	263.2114	1.735249E-03	3.437025E-09	1.482940E-14	CH. 11 IF AMP
29	263.5982	1.739884E-03	3.453134E-09	1.511657E-14	DC/DC CONVERTER
30	263.6745	1.740961E-03	3.504604E-09	1.020772E-14	CH. 13 IF AMP
31	263.5161	1.740639E-03	3.555526E-09	1.322503E-14	CH. 14 IF AMP
32	263.0495	1.754169E-03	3.485714E-09	9.320949E-15	CH. 12 IF AMP
33	263.7456	1.728914E-03	3.995046E-09	6.167204E-15	Al-1 RFSHELF
34	264.5479	1.742660E-03	3.626275E-09	1.254250E-14	A1-2 RFSHELF
35	263.4284	1.726987E-03	3.945960E-09	7.127579E-15	DETECTOR/PRE-AMP
36	254.8741	1.637893E-03	5.902597E-09	2.947544E-14	Al-1 WARM LOAD 1
37	254.4845	1.634828E-03	5.860831E-09	2.909161E-14	Al-1 WARM LOAD 2
38	254.5352	1.627381E-03	5.878626E-09	3.107089E-14	Al-1 WARM LOAD 3
39	254.1389	1.638958E-03	5.777645E-09	2.745335E-14	Al-1 WARM LOAD4
40	254.6330	1.636434E-03	5.860082E-09	2.914760E-14	A1-1 WARM LD CENTER
41	254.6251	1.637361E-03	5.846084E-09	3.029967E-14	A1-2 WARM LOAD 1
42	254.6393	1.637848E-03	5.829585E-09	2.888881E-14	A1-2 WARM LOAD 2
43	254.1602	1.630533E-03	5.839149E-09	3.010925E-14	A1-2 WARM LOAD 3
44	255.0383	1.634408E-03	5.912973E-09	2.908720E-14	A1-2 WARM LOAD 4
45	254.3306	1.632722E-03	5.889366E-09	2.895453E-14	A1-2 WARM LD CENTER

Table A-2. NOAA-N' AMSU-A2 SN 109: Polynomial coefficients for converting PRT counts into temperatures.

PRT	$f_{k0}$	$f_{k1}$	$f_{k2}$	$f_{k3}$	Description
1	263.3287	1.742668E-03	4.311887E-09	2.025034E-16	Scan Motor
2	263.3765	1.752967E-03	3.748854E-09	1.258927E-14	Feedhorn
3	263.5512	1.739217E-03	3.818752E-09	1.353860E-14	RF Diplexer
4	263.6189	1.749884E-03	3.816858E-09	1.067548E-14	Mixer/IF CH1
5	263.9921	1.751035E-03	3.631815E-09	1.485421E-14	Mixer/IF CH2
6	263.7419	1.753762E-03	3.693388E-09	1.127053E-14	CH 1 DRO
7	263.4610	1.747977E-03	4.016104E-09	7.197988E-15	CH 2 DRO
8	263.4379	1.763693E-03	3.358171E-09	1.355759E-14	Compensat Motor
9	263.9190	1.751008E-03	4.282710E-09	2.331043E-15	Sub Reflector
10	262.9899	1.747058E-03	3.705516E-09	1.233494E-14	DC/DC Converter
11	263.5386	1.745205E-03	4.009926E-09	9.340984E-15	RF Shelf
12	263.2827	1.749743E-03	4.089112E-09	6.352310E-15	Det. Pre-Amp
13	254.4454	1.646443E-03	5.907230E-09	3.059857E-14	Warm Load Ctr
14	254.3621	1.658168E-03	5.933477E-09	2.513103E-14	Warm Load #1
15	254.8100	1.652669E-03	5.849476E-09	3.025275E-14	Warm Load #2
16	254.1330	1.648472E-03	5.857347E-09	3.302980E-14	Warm Load #3
17	254.9246	1.648978E-03	5.884300E-09	3.126444E-14	Warm Load #4
18	254.7082	1.641230E-03	5.955578E-09	3.155298E-14	Warm Load #5
19	254.6145	1.648466E-03	5.895623E-09	3.138678E-14	Warm Load #6

### A.2 Warm Load Correction

The in-flight warm load correction (WLC) was calculated according to a formula developed by Aerojet [6-7]. For each AMSU-A antenna system, a special set of calibration data were acquired by setting the temperature of its variable scene target equal to that of the internal blackbody (warm) target. The physical temperature  $T_W$  of the internal blackbody target was determined from the PRT counts as described in Section 2. The radiometric temperature  $T_{wrad}$  of the blackbody (in-flight warm load) was calculated (for each scan in the data set) by the formula

$$T_{wrad} = T_{sprt} + \left(T_{sprt} - T_C\right) \left(\frac{C_W - C_S}{C_S - C_C}\right)$$
(A-3)

where:

 $T_{sprt}$  PRT temperature of the variable scene target,

 $T_C$  PRT temperature of the cold target,

 $C_W$  the average of two radiometric counts from the warm target,

 $C_C$  the average of two radiometric counts from the cold target, and

 $C_S$  radiometric counts from the variable scene target.

One should note that temperatures,  $T_{sprt}$  and  $T_C$ , from the scene and cold targets are used as the two reference calibration points in Equation (A-3) to calculate the radiometric temperature of the warm target. The in-flight warm load correction factor  $\Delta T_W$  was computed from the formula,

$$\Delta T_W = \frac{1}{N} \left[ \sum_{i=1}^{N} \left( T_{wrad} - T_W \right)_i \right] \tag{A-4}$$

where N represents the number of scans in a data set. The  $\Delta T_W$  values at three instrument (RF Shelf) temperatures for each AMSU-A antenna system are listed in Table A-3. For AMSU-A1, the  $\Delta T_W$  values for both PLLO#1 and PPLO#2 were calculated and listed. These  $\Delta T_W$  values will be used in both NOAA-N' AMSU-A operational calibration algorithms.

## **A.3** Nonlinearity parameters

The nonlinearity is discussed in Section 4.2 and the values of the nonlinearity parameter u are listed in Table 4.

Table A-3. NOAA-N' AMSU-A warm load corrections (K) at three instrument temperatures.

AMSU-A2 SN 109 Channels

AMSU-A1-2 SN 107 Channels

Instrument Temp.(C)	Ch. 1	Ch. 2	Instrument Temp.(C)	Ch. 3	Ch. 4	Ch. 5	Ch.8
-7.26	0.142	0.002	-2.49	-0.026	-0.002	-0.027	-0.087
11.53	0.062	0.008	17.62	0.093	0.009	0.005	-0.031
30.04	0.040	-0.029	37.92	0.041	-0.066	-0.066	-0.187

### AMSU-A1-1 SN 107 Channels: PLLO#1:

Instrument Temp.(C)	Ch. 6	Ch. 7	Ch. 9	Ch. 10	Ch. 11	Ch. 12	Ch. 13	Ch. 14	Ch. 15
-2.52	0.338	0.339	0.324	0.375	0.349	0.321	0.309	0.379	0.348
17.47	0.442	0.416	0.403	0.405	0.419	0.411	0.426	0.403	0.396
37.88	0.499	0.538	0.506	0.481	0.487	0.498	0.511	0.459	0.470

### AMSU-A1-1 SN 107 Channels: PLLO#2:

Instrument Temp.(C)	Ch. 9	Ch. 10	Ch. 11	Ch. 12	Ch. 13	Ch. 14
-2.67	0.281	0.282	0.288	0.272	0.243	0.209
17.47	0.722	0.717	0.731	0.692	0.703	0.721
37.60	0.278	0.313	0.283	0.281	0.280	0.281

#### A.4 Correction to In-orbit Cold Space Calibration

For on-orbit cold space calibration, there is an uncertainty due to antenna side lobe interference with the Earth limb and spacecraft. The contribution from this uncertainty should be added to the cold space cosmic background temperature of 2.73K. Therefore, the "effective" cold space temperature  $T_{EC}$  can be represented by

$$T_{EC} = 2.73 + \Delta T_C \tag{A-5}$$

where  $\Delta T_C$  represents the contribution from the antenna side-lobe interference with the Earth limb and spacecraft. Estimates of  $\Delta T_C$  for individual channels of NOAA-N' AMSU-A are made from the measured antenna pattern data [8]. These estimated  $\Delta T_C$  values for the four available cold calibration positions are listed in Table A-4.

Table A-4. NOAA-N' AMSU-A cold bias  $\Delta T_C$ .

Pos. ID	Anglea	Ch. 1	Ch. 2	Ch. 3	Ch.4	Ch. 5	Ch. 6	Ch. 7	Ch. 8	Ch. 9	Ch. 15
1	83.333	1.34	0.48	1.83	2.19	1.47	1.32	1.69	1.73	1.57	0.61
2	81.667	1.33	0.45	1.67	2.03	1.38	1.20	1.57	1.59	1.45	0.57
3	80.000	1.33	0.44	1.60	1.94	1.34	1.14	1.5	1.52	1.38	0.55
4	76.667	1.33	0.44	1.50	1.84	1.28	1.08	1.42	1.45	1.31	0.53

<sup>&</sup>lt;sup>a</sup> Measured from nadir.

### A.5 Limit of Blackbody and Cold Counts Variation

For each scan, the blackbody counts  $C_W$  is the average of two samples. If the two samples of the blackbody differ by more than a pre-set limit of blackbody count variation  $\Delta C_W$ , the data in the scan will not be used. The  $\Delta C_W$  values for individual channels are listed in Table A-5. These  $\Delta C_W$  values, which equal approximately  $3\sigma$  (where  $\sigma$  is the standard deviation of the internal blackbody counts), are calculated from the TV calibration data. Similarly, the cold count is the same.

Table A-5. NOAA-N': Error Limits of Warm and cold radiometric counts between samples of same scan line.

Limit	Ch.1	2	3	4	5	6	7	8	9	10	11	12	13	14	15
Warm	18	18	18	18	18	18	18	18	18	18	24	24	30	60	22
Cold	18	18	18	18	18	18	18	18	18	18	24	24	30	60	22

## A.6 Pre-launch Determined Weight Factors $w_k$ for the Internal Blackbody PRTs

The weight factors  $w_k$  (see Equation 2) assigned to individual PRTs in the internal blackbody targets are listed in Table A-6. All NOAA-N' AMSU-A PRTs are good.

Table A-6. Pre-launch determined weight factors w<sub>k</sub> assigned to NOAA-N' AMSU-A PRTs in blackbody targets.

Antenna System	$\mathbf{w}_1$	$\mathbf{w}_2$	W <sub>3</sub>	$W_4$	W <sub>5</sub>	$W_6$	$\mathbf{w}_7$
AMSU-A2	1	1	1	1	1	1	1
AMSU-A1-1	1	1	1	1	1		
AMSU-A1-2	1	1	1	1	1		

## A.7 Conversion Coefficients of Analog Data

AMSU-A instrument has an analog telemetry bus to monitor key temperatures and voltages through the spacecraft. The resolutions of the analog telemetry received on the ground is 20 mV for one part in 256. To convert the analog data into physical quantities y, one must multiply the analog values by 0.02V (20 mV) to obtain the measured output, x, in volts and then uses the conversion equation,

$$y = B + M x \tag{A-6}$$

where the values of B and M are given in Tables A-7 and A-8 for NOAA-N' AMSU-A2 SN 109 and AMSU-A1 SN 107, respectively.

Table A-7. NOAA-N' AMSU-A2 SN 109: Analog data conversion coefficients.

UIIS	Description	у	M	В
Ref.				
1	Scanner Motor Temperature	K	68.027	0
2	Comp Motor Temperature	K	68.027	0
3	R.F. Shelf Temperature	K	68.027	0
4	Warm Load Temperature	K	68.027	0
5	Comp. Motor Current (average)	mA	46.6	0
6	Ant Drive Motor Current (ave)	mA	46.6	0
7	Signal Processing +15Vdc	V	4.315	0
8	Antenna Drive +15Vdc	v	4.315	0
9	Signal Processing -1 5Vdc	V	2.504	-22.562
10	Antenna Drive - 15Vdc	V	2.504	-22.562
11	Mixer/IF Amplifier +10 Vdc	V	2.889	0
12	Signal Processing +5Vdc	V	1.667	0
13	Antenna Drive +5Vdc	V	1.667	0
14	Local Oscillator +lOVdc. (ch.1)	V	2.861	0
15	Local Oscillator +lOVdc (ch.2)	V	2.861	0

Pin#	Description	y	M	В
3	Al-1 Scanner Motor Temperature	K	68.027	0
22	A1-2 Scanner Motor Temperature	K	68.027	0
2	A1-1 RF Shelf Temperature	K	68.027	0
21	Al-2 RF Shelf Temperature	K	68.027	0
4	Al-1 Warm Load Temperature	K	68.027	0
23	Al-2 Warm Load Temperature	K	68.027	0
8	Al-1 Antenna Drive Motor Current	mA	23.3	0
27	Al-2 Antenna Drive Motor Current	mA	23.3	0
11	Signal Processing (+15 VDC)	v	4.315	0
9	Antenna Drive (+15 VDC)	v	4.315	0
29	Signal Processing (-15 VDC)	v	2.504	-22.562
28	Antenna Drive (-15 VDC)	v	2.504	-22.562
34	Receiver Amplifiers (+8 VDC)	v	2.50	0
12	Signal Processing (+5 VDC)	v	1.667	0
10	Antenna Drive (+5 VDC)	v	1.667	0
17	Receiver Mixer/IF (+10 VDC)	v	2.889	0
16	Phase Lock Loop Ch 9/14 (+15 VDC)	v	4.315	0
33	Phase Lock Loop Ch 9/14 (-15 VDC)	v	2.504	-22.562
13	Ch 3 L.O. Voltage (50.3 GHz)	v	2.861	0
30	Ch 4 L.O. Voltage (52.8 GHz)	v	2.861	0
14	Ch 5 L.O. Voltage (53.596 GHz)	v	2.861	0
31	Ch 6 L.O. Voltage (54.4 GHz)	v	2.861	0
15	Ch 7 L.O. Voltage (54.94 GHz)	v	2.861	0
32	Ch 8 L.O. Voltage (55.5 GHz)	v	2.861	0
25	PLLO Primary Lock Detect (PLO #1)	V	1.0	0
6	PLLO Redundant Lock Detect (PLO #2)	V	1.0	0
_18	Ch 15 L.O. Voltage (89.0 GHz)	V	4.315	0

#### **APPENDIX B**

#### NOAA POLAR ORBITER LEVEL 1B DATA

The NOAA Polar Orbiter Level 1B data are raw data that have been quality controlled and assembled into discrete data sets, to which Earth location and calibration information are appended but not applied. For simplification of application, the conversion of scene counts  $C_S$  into scene radiance  $R_S$  is accomplished by writing Equation (4) as

$$R_S = a_0 + a_1 C_S + a_2 C_S^2 \tag{B-1}$$

where the calibration coefficients  $a_i$  ( where i = 0, 1, and 2), which are output in the 1B data sets for individual scans, are expressed in terms of averaged calibration counts (over 7 scans) and calibration target temperatures. Derivation of these coefficients is described in the *NOAA KLM User's Guide*. In Equation (B-1), there is a quadratic term which represents the nonlinearities in the measurements [1]. Users, who prefer scene temperature instead of radiance, can make the simple conversion,

$$T_{\rm S} = B^{-1}(R_{\rm S}) \tag{B-2}$$

where  $B^{-1}(R_S)$  is the inverse of the Planck function for radiance  $R_S$ . The  $T_S$  is the converted antenna scene temperature.

The above calibration process applies to all AMSU-A measurements from NOAA-15, -16, -17, 18, and -N'.

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- NESDIS 122 JCSDA Community Radiative Transfer Model (CRTM). Yong Han, Paul van Delst, Quanhua Liu, Fuzhong Weng, Banghua Yan, Russ Treadon, and John Derber, December 2005.
- NESDIS 123 Comparing Two Sets of Noisy Measurements. Lawrence E. Flynn, November 2006.

## NOAA SCIENTIFIC AND TECHNICAL PUBLICATIONS

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