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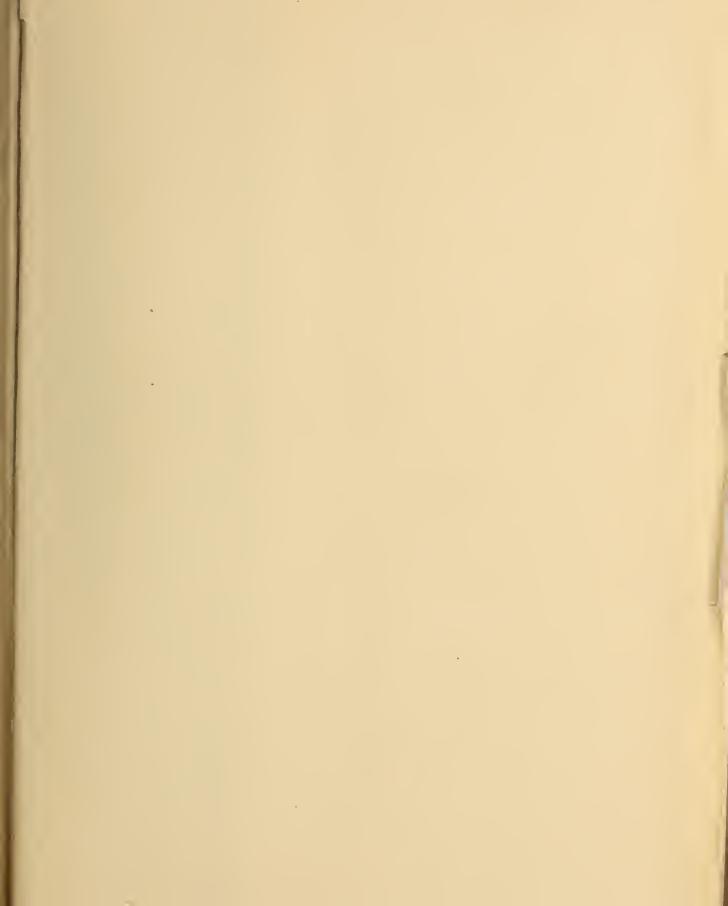
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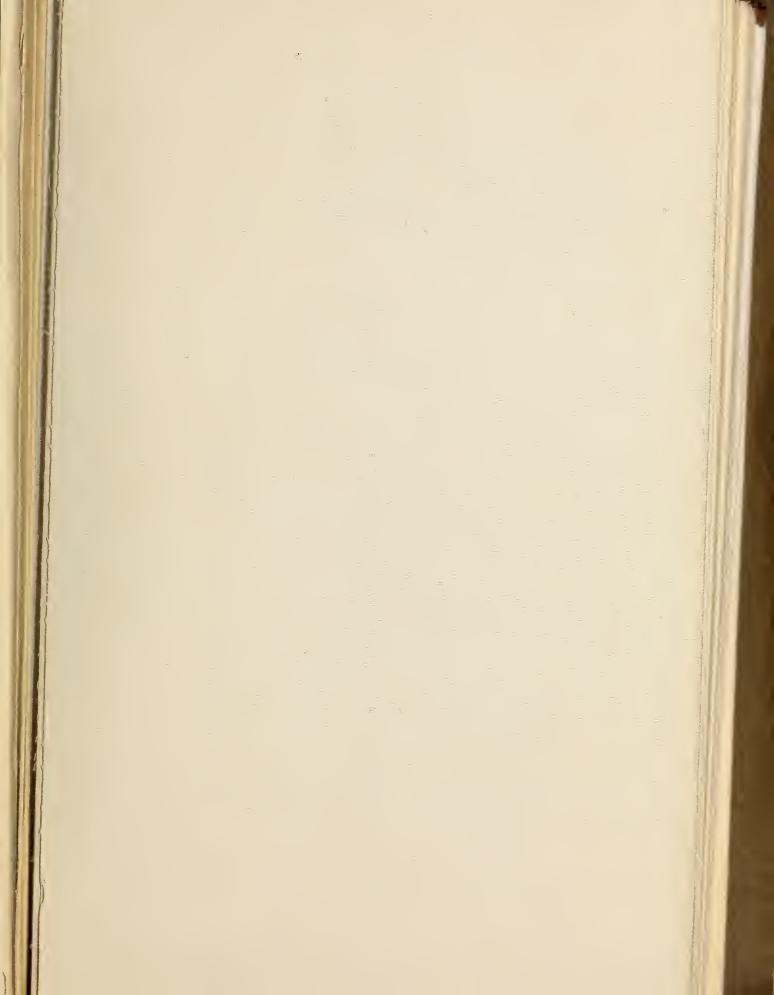
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TABLES AND GRAPHS FOR FACIL ITATING THE COMPUTATION OF SPECTRAL ENERGY DISTRIBUTTION BY PLANCK'S FORMULA

CONSISTING OF SEVEN SHEETS (FIVE CHARTS) AS FOLLOWS: TEXT (SH EET 1). GRAPHS: 1,000 TO 5,000°K. (SHEET 2, 3, 4; CHARTS 1, 2, 3); S,000 TO 8,000°K. (SHEET 5; CHART 4); 8,000 TO 24,000°K. (SHEET 6; CHART 5). TABLES (SHEET 7).

By M. KATHERINE FREHAFER, Associate Ph ysirist, and CHESTER L. SNOW, Draftsman¹ Baraut of Stanlards

ABSTRACT

Very frequently it has been need body" in the visible spectral restary to compute the energy distribution of a "black

 $E_{\lambda} = \frac{C_1 \lambda^{-4}}{\frac{C_1}{14}}$

Such computation consumes much time and labor, in consequence, the following short cuts in the way of tables and graphs have been devised. Using the equation in the form

$$\frac{E_{\lambda}}{E_{R}} = \left(\frac{A}{\lambda \theta}\right)^{\dagger} \left(e^{\tilde{A}} - \epsilon\right) \left(e^{\tilde{b}} - \epsilon\right)^{-1}$$

a table has been made with values of $\lambda \theta$ and corresponding values of $\frac{E_{\alpha}}{E_{\alpha}}$. For any pirce temperature the wave lengths in the visible and particly in the utra-violet and inferred regions, corresponding to these values of relative energy, can readfly be atomic from the products M. Hence the spectral energy care can be pilted at a torse of the spectra of the spect

Two sets of tables have also been made up, each for al temperatures, in which the relative energy values compared directly from Plance's formation are recorded for the avare length as 0 to 720 mm or 10 threads in 1000 mmb, and recorded for the energy value equal to 200 mm or 1000 mmb, the other equal to 100 at 560 mm. The tempera-ture range is form 1000 to 36000 K, (ccentray, value) and 560 mm. The tempera-ture range is form 1000 to 36000 K, (ccentray, value) value or 100 mmb, the energy star the relative energy has been plated (accuracy, within a signer energy star also been energy has been plated (accuracy, within a signer energy) are output also been temperature, and one complete energy forms for each of the 33 wave lengths, from these graphs, the energy distribution for the visible region can be obtained in a few minutes for any temperature whatever in the range abave specified.

I. INTRODUCTION

In connection with the work of the color laboratory, very frequent use is made of the spectral energy distribution at various temperatures of "black body" radiation, as represented by Planck's formula:

$$E_{\lambda} = \frac{C_1 \lambda^{-3}}{\frac{C_1}{e^{\lambda \beta} - 1}}$$

(1)

 E_s is the radiant energy for any wave length λ_i at any temperature θ (degrees absolute); C_i and C_j are constants, and e is the base of the natural logarithms. As a rule it is desired to know the relative values of E at a given temperature for a range of 320 nm

throughout the visible spectrum; that is, 400 to 720 mµ, at interdironglouit the visible spectrum; that is, 400 to 720 ma, at inter-vals of 10 ma. Moreover, these values must conform to a scale such that at some particular wave length *E* will always be the same for all temperatures, for instance, equal to 100 at wave length 560 ma.³ When computed directly from the formula, this proves to be a long and tedious process; therefore, certain short ents have been devised. Tables have been made and curves plotted which greatly facilitate this work. It was thought worth while to publish these for the benefit of others who make similar use of this formula. The following reference tables have been found of great assistance: F. W. Newman and J. W. D. Glaicher, Tables of Exponential Functions, C. E. VanOrstrand, Tables of the Exponential Function.²

II. EXPLANATION OF TABLES

1. TABLE I

Instead of using Planck's equation in the form given above, it was combined ⁶ with Wien's displacement law connecting the temporature with the wave length at which the maximum energy occurs; that is,

$$\lambda_m \theta = A_1$$
 a constant
gy is given by

$$\int_{a} \frac{C_1 \lambda_m^{-4}}{e^{\lambda_m^0} - 1}$$

Ē Substituting from (2) $\lambda_m = \frac{A}{R}$, equation (3) takes the form

(5)

Dividing (1) by (4) we obtain

This maximum ene

$$\frac{E_{\lambda}}{E_{m}} = \left(\frac{A}{\lambda \theta}\right)^{s} \left(e^{\frac{C_{\beta}}{\lambda}} - t\right) \left(e^{\frac{C_{\beta}}{\lambda \theta}} - 1\right)^{-1}$$

Au inspection of this equation will show that the right-hand side is a function of $\lambda\theta$ alone. The values of the constants used in the computations are:

A = 2,890 micron degrees $C_2 = 14.350$ micron degrees.⁴

Table 1 has been arranged with a wide range of values of $\lambda \theta$ and corresponding values of $\frac{E_{\lambda}}{E_{m}}$ computed to four significant figures.

¹ The authors with its exhausticity that indefendence is Harry J. Kerglettet up vor annumeration. Support in N. Sector and Sector is a sector of the sector of landscaley databases and landscale provided in the form which seek sector and landscale provided in the form. Work, Area, Sci. 3, No. 11, Janese, Isag. ¹ Politational fitth by four works, the sector of landscale provided in the form wavegood of the L. Devic, and S. Sector and Sector a ish is acknowledge (beir indebtedness to Harry) Keepsu for his shie assistance with the

Use of Table 1.—From the products $\lambda\theta$ the wave lengths can be readily computed for any term verture θ . (These wave lengths are in microns, and must be mult tiplied by 1,000 to be expressed in millimicrons.) The values of relative energy at these wave lengths are already known from (he table; bence the distribution curve for this temperature can at once be plotted. The next

step is to transform the values $\frac{E_{\mu}}{E_{m}}$ so that at 560 mµ the energy will be equal to 100. This is accomplished by reading from the curve the energy coordinate for 56 or mµ, taking its reciprocal (times 100) and multiplying the other c soordinates by this factor. The spectral energy curve is thus we ll defined and will pass through the required point at 560 mµ.

the required point at 560 nµ. In which we have a structure for the theorem of theorem of the theorem of theorem of the theorem of theorem of theorem of theorem of theorem

 $\lambda \theta_i$ (2) plotting of λ against $\frac{E_{\lambda}}{E_m}$ given in the table; (3) reading from

this curve the values of $\frac{E_{\lambda}}{E_{m}}$ at in itervals of 10 mµ; and (4) multiplying these values by the reciproc al of the value at 560 m μ , and the

plying these values by the reciprod product by 100. This process of computation, while more easily and somewhat more rapidly performed than di-formula, still consumes a consid-formula, still consumes a consid-tion over, without extreme care in accuracy obtainable may not be use of Table 1 can not be recom in the plotting of the curves, the greater than 0.6 per cent. The greater speed or great accuracy, under the conditions in the introduction; but it is very useful to give readily the shape of the energy distribu-tion curve and the relative value s of the energy when the above conditions need not be met. More reover, with the use of Table 1, the curve can be extended into ' the ultra-violet and infra-red

These tables have been prep ared directly from Wicn's and Planck's formulas, with the extreption only of the values for $1,000^{\circ}$ K., which were calculated directly from Table 1. Wicn's formula for 2,200 to 28,000° K. Since the energy values given in these tables are all relative to the value chosen for wave length 560 (or 590) mg the constant C, chosen to ensideration. G is taken equal to 14,330 micro in degrees. The values in these tables are as follows:

tables are as follows:

1000 10 210001 0	19 Steps of 100 .
2,000 to 3,000, t	iy steps of 200°.
3,000 to 4,000, t	vy steps of 250°.
4,000 to 7,000, t	by steps of 500°.
7,000 to 10,000, b	>y steps of 1,000°.
10,000 to 28,000, t	by steps of 2,000°.
	n made equal to 100 at 590 mµ fo
temperature; in Table 2B t	he value 100 has been chosen a

III. EXPLANATIC ON OF GRAPHS

The plotting of the graphs we is suggested by the existence of Tables 2A and 2B. The energy d istribution is often required to be known for some temperature intermediate to those given, and a series of curves based on these t ables sbould be very suitable to yield these data. This led to a s crice of isochromatic curves, one for each wave length at intervals of from μ from μ to to π to π . 1,000 to 28,000° K

560

1,000 to 28,000° K. The plotting of these eurves in it with some difficulties for the following reasons: (i) It was de sired to obtain an accuracy of plotting and of reading to 0.33 p er cent. (a) The energy values vary from 0.007 (approximately) to 5,000; in other words, the largest value is about 90,000 t interest the samelest value. This increases several changes of si (a) The isochromatics crowd competitives 4,000 to 7,000°. The word these considerations the value of no at 500 ms, The obsets by the use of a suitable fr vary the determined by the obtained from the contrast of a suitable fr vary the largest value of no at 500 ms. The vary the contrast of a suitable fr vary WE largest. The latter permitted the

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that is, $\frac{E_{\lambda}}{E_{\lambda 0}} = f(t) \neq (\lambda \delta)$.

above (0.33 per cent) should apply except for $E_{\rm h}$ less than 3.5 and possibly also on very steep parts or near the line $\theta = 1,000^{\circ}$ K. The temperature can be estimated to within 1° for the range 1,000 to $8,000^{\circ}$, and to within 5° for the range 8,000 to $28,000^{\circ}$. The curves have been tested by comparing values furnished by the graphs (for given temperatures other than those recorded in Table 2B) with the values computed for these temperatures.

IV. PRACTICAL APPLICATION

It has been shown ' that a number of the ordinary illuminants have a spectral energy distribution that approximates closely that of a Planckian radiator. This property is valuable in that the unknown spectral energy distribution of an illuminant can very readily be determined if it can be matched in color with a so-called "black body." The temperature to which the "black body" is brought in order to produce color match with the illuminant is defined as the color temperature of the latter. Since this quan-tity is not difficult to determine in laboratories having suitable equipment, the energy distribution in the visible region of such illuminant can at once be recorded by reading from the charts the energy values for the given temperature. the energy values for the given temperature.

V. SUMMARY

If it is required to know the spectral distribution of a Planckian radiator at any temperature θ throughout the visible and indra-dutra-violet and indra-ref regions, Table taffords a ready means of computing it with very little labor. But if the distribution curve is required to pass through a certain point (for instance, too at 560 mg), and if the values of energy are desired for 10 mg inter-vals, further computation is necessary. Tables 2A and 2B have been computed giving these values for 3S temperatures in the interval 1,000 to 35,000 K. In order to obtain directly the spectral distribution for any temperature whatever in this interval, at wave lengths 10 mg apart, and having the value to 0a t500 mg, a series of isochromatic curves has been constructed. From these curves, which are plotted to a large scale, the energy values can be read to an accuracy invadite mitting at the ora. The measurement we choose the spectral plotted to a large scale, the energy values can be read to an accuracy If it is required to know the spectral distribution of a Planckian

usually within 0.33 per cent, the temperature values within 0.1 per cent.

VI. APPENDIX .

The question may arise as to what variation will be caused in the relative values of $E_{\rm c}$ by the adoption of a value of $C_{\rm c}$ different from 14,350; for example, Coblentz's most recent value "for this constant is 14,320; the International Critical Tables has adopted 14,330; and further research may eventually cause further, though probably very small, revision of these latest values.

An inspection of the equation, $E_{\lambda} = \frac{C_{1}\lambda^{-4}}{c^{14}}$ shows that the

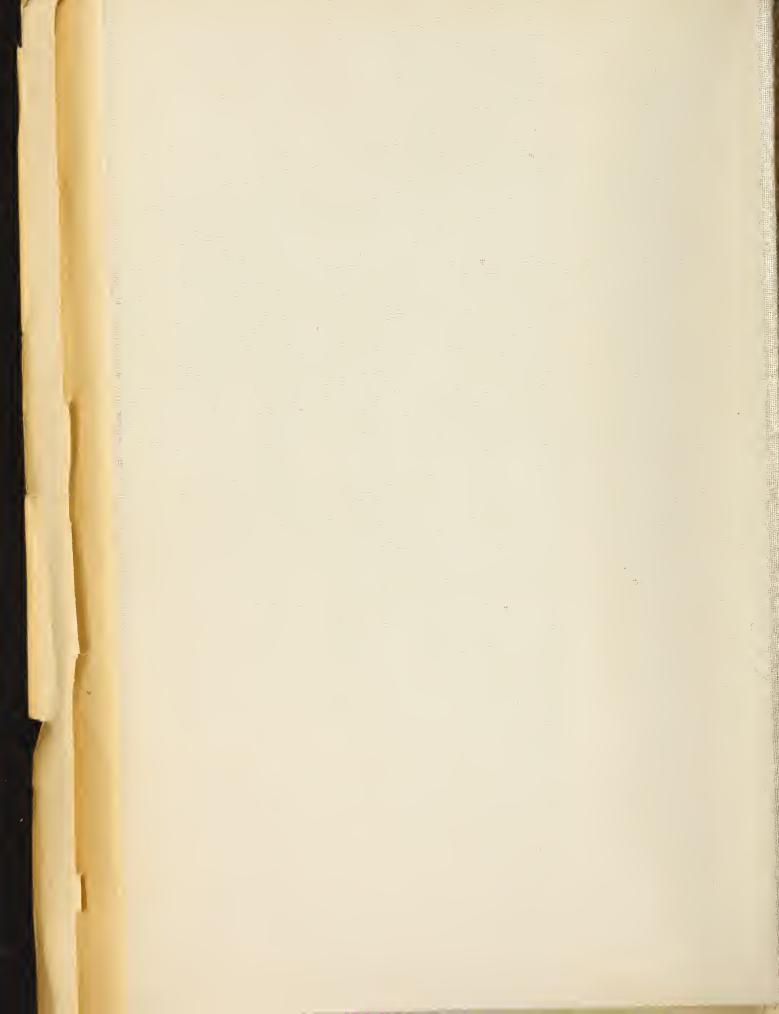
effect of a variation of C_s with θ constant is the same as, but opposite in sense to, a variation of θ with C_s constant. So far as the graphs are concerned, in which values of E_s are plotted against values of θ for the various wave lengths, we have, value of C_s which it may be desired to use. The method is simply this: Compute the percentage difference between 14,350 and the value to be used. Call this percentage difference p. Multiply the temperature by $1 \pm \beta 100$ and read values of E_s from the graphs at this corrected temperature. If C_s is decreased, θ must be increased, and vice varia.

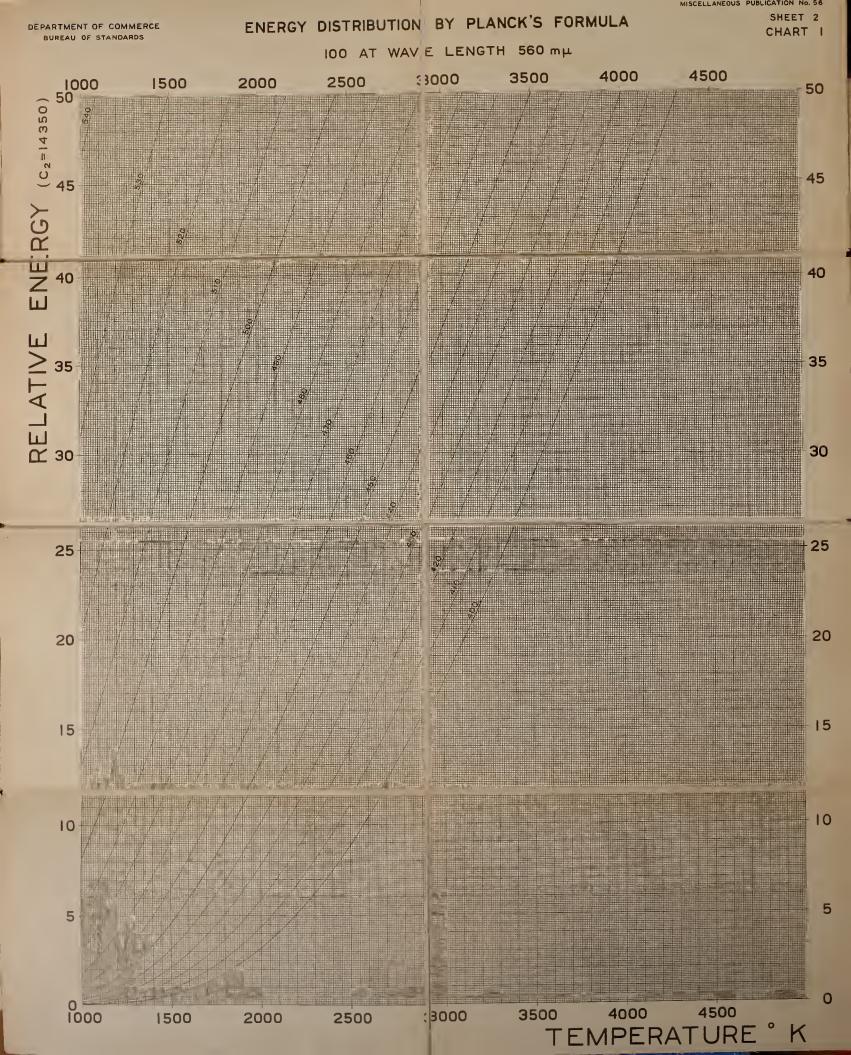
at this corrected temperature. If C_p is decreased, θ must be in-ereased, and vice versa. While true values of E_p may thus be obtained for any value of C_p , and as readily as for the value 14,350, the difference in E_p obtained by using 14,330 or 14,330 is so small as to be negligible for most work for which these tables and graphs will be used. In the first place, the difference in temperature necessary to make the correction on the graphs is about 0.2 per cent, the order of magni-

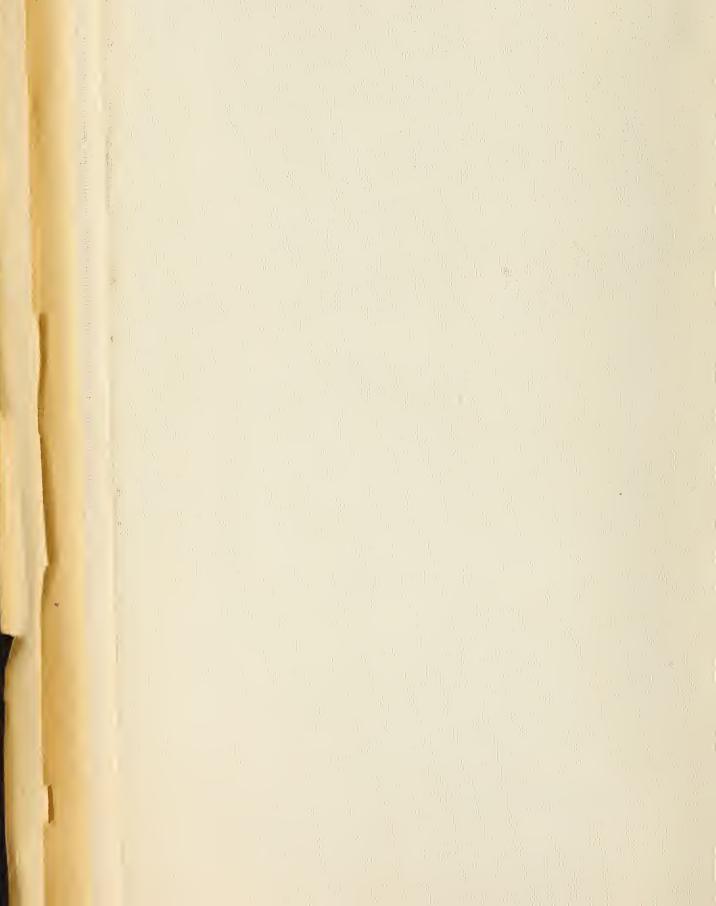
correction on the graphs is about 0.2 per cent, the order of magnitude of differences in color temperature determinations by experi-oneed observes. In the second place, the differences in values of E_i relative to that at 500 me caused by this 0.2 per cent change in E_i is but a few tenths of 1 per cent, except at the extreme wave lengths for temperatures below about 1,500°, where they may be concepted by a second place, they may be considered by the second place, they may be considered by the second place, between spectral energy measurements and values of energy com-puted for the corresponding color temperature may be several per cent in the end regions of the visible spectrum. While, therefore, true values of E_i may be readily obtained from the error involved in using values of E_b based on $C_b = 14,350$ -the tis, those given by the tables or read from the graphs for the temperature correction—will be negligible.

⁴ Hydr and Flerviche, The quality of light from an illuminary is inflated by its color if J. Panki, Inte, JSJ, pp. 357-358. "Proc. Measurement of the Gale Competitured the Mar-Manizal Light Sources by the Attechni di Rokardy. Discours, B. S. Se, Panes. B. S. Se, Panes. "The stathary with as achieved by the augmention of Dr. K. S. Gahan is constraint with the B. S. Son, Panes As and Sources, rates

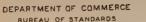
regions of the spectrum 2. TABLES 2 A AND 2B







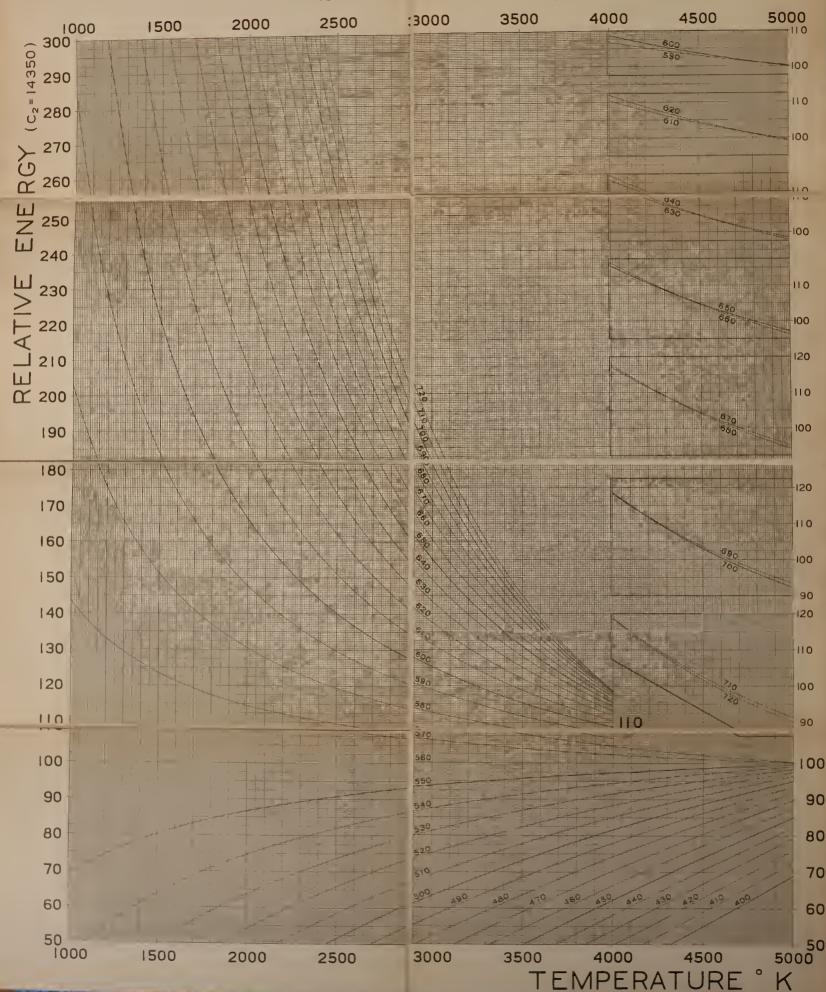
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ENERGY DISTRIBUTION BY PLANCK'S FORMULA

MISCELLANEOUS PUBLICATION NO. 56 SHEET 3 CHART 2

100 AT WAVE LENGTH 560 mm



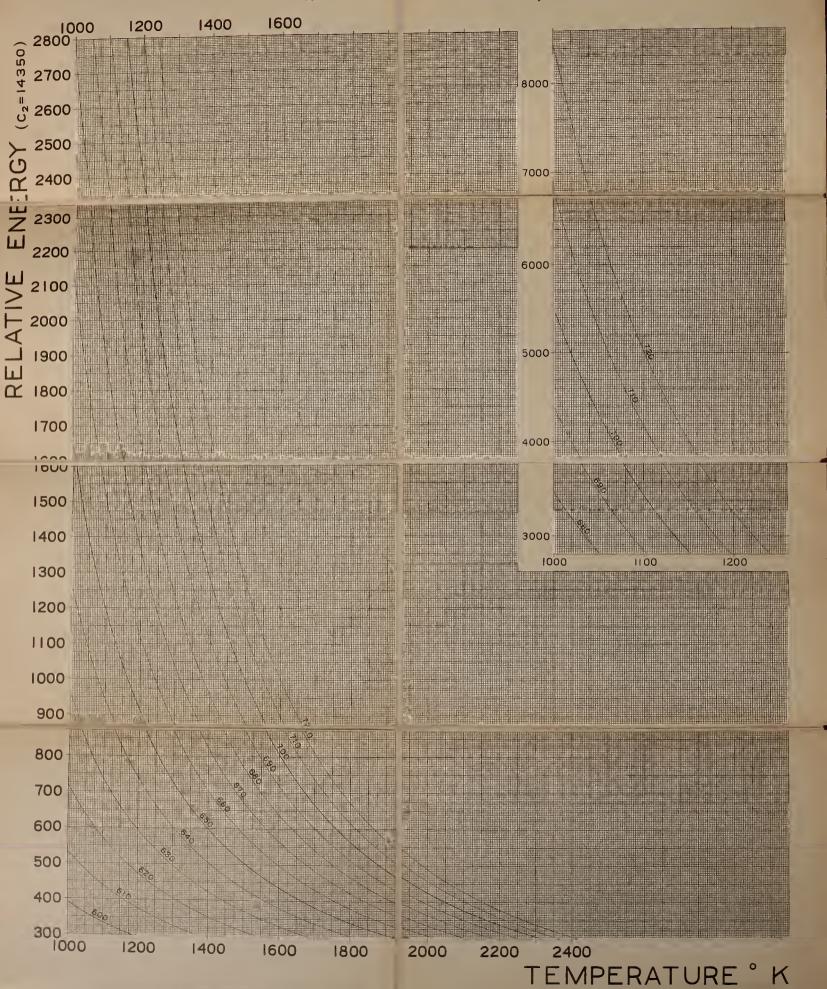


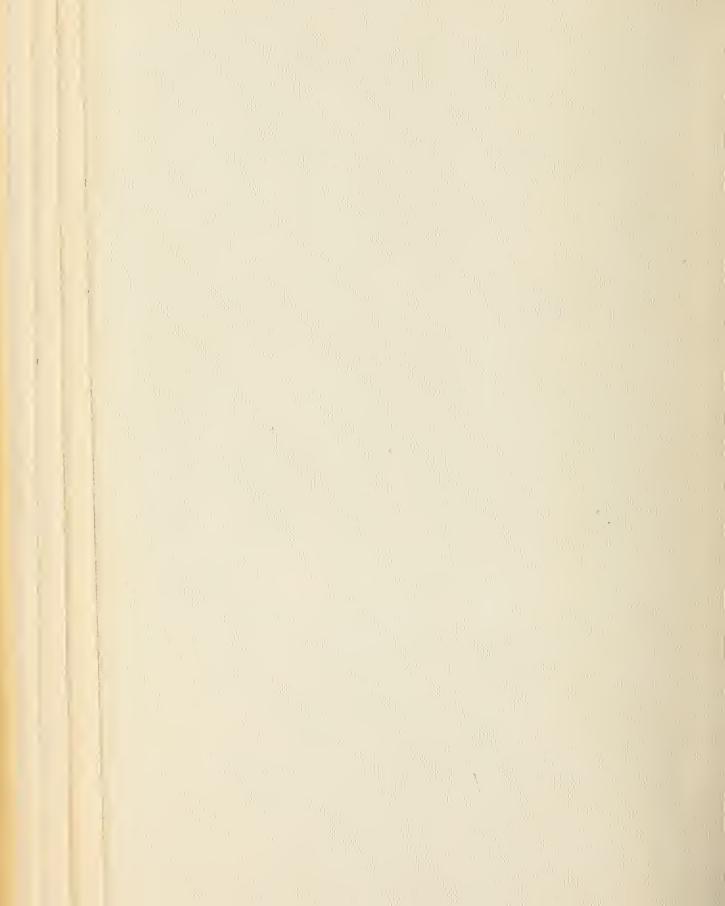
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DEPARTMENT OF COMMERCE BUREAU OF STANDARDS

ENERGY DISTRIBUTIC IN BY PLANCK'S FORMULA

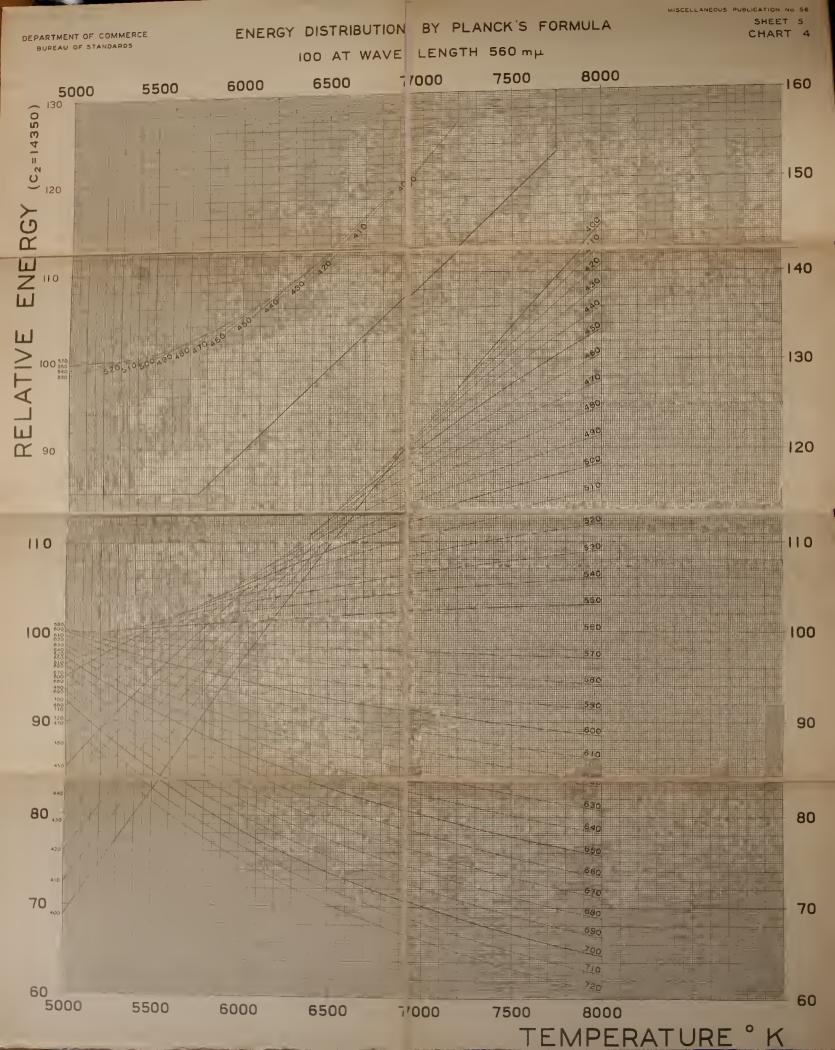
100 AT WAV E LENGTH 560 mp



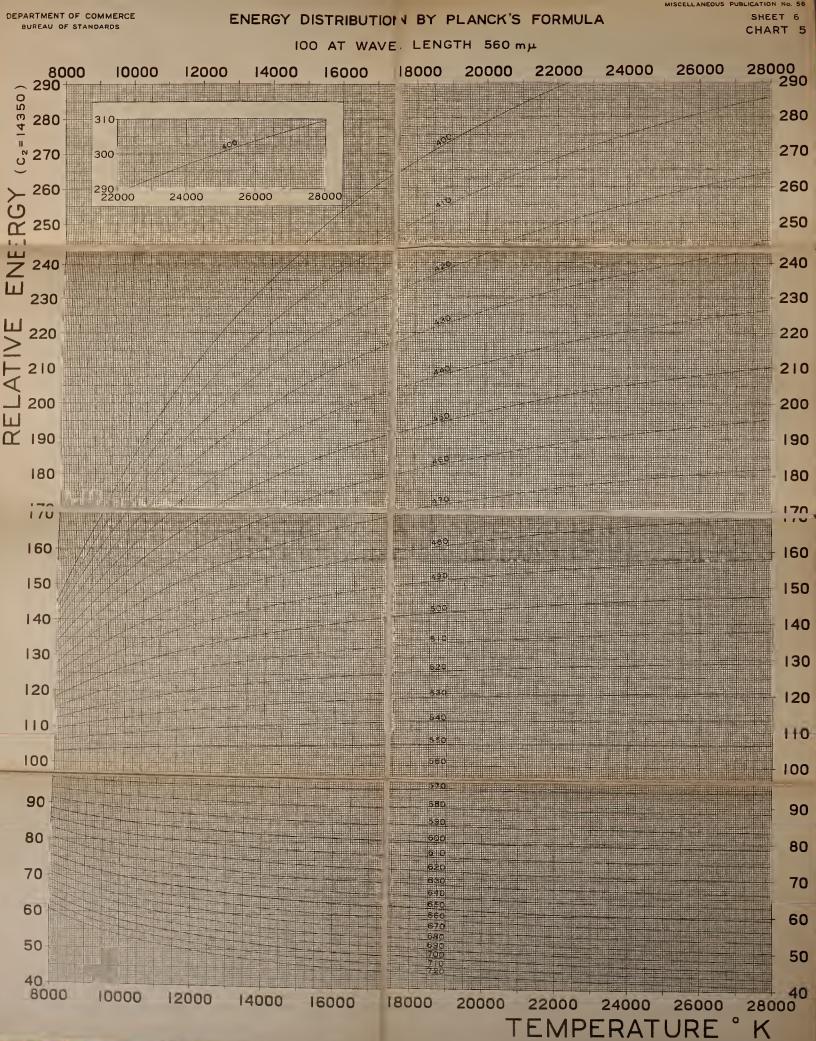


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M. Katherine Frehafer Chester L. Snow

TABLES AND GRAPHS FOR FACIL .ITATING THE COMPUTATION OF SPECTRAL ENERGY DISTRIBUTION BY PLANCK'S FORMULA

MISCELLANEOUS PUBLICATION No. 54 SHEET 7 TABLES

WASHINGTON : GOTERNMENT PRINTING OFFICE : 1025

TABLE 1- $\frac{E_{\lambda}}{ \mathcal{E}_{\infty}^{m} } \left(\frac{\Lambda}{\lambda^{3}}\right)^{c} \left(e^{\frac{C_{\lambda}}{\Lambda}}-1\right) \left(e^{\frac{C_{\lambda}}{\lambda^{2}}}-1\right)^{-1}$	TABLE 2A Continued	TABLE 2B-Continued Relifie values of EA for the following temperatures, #in degrees basicle
[A=2,050 micron degrees, Cy=14,350 micros degrees, c ⁶ →1−142,37]	A in millimiterons Relative raises of EA for till to following temperatures, 9 in degrees absolute 2,000° 2,000° 2,000° 2,000° 3,00° 4,000° 4,000° 5,000°	λ in millimitrans 1,700° 1,000° 1,000° 2,000° 2,000° 2,000° 2,000° 2,000°
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TABLE 2A.—Energy Distribution of Planckian Radiator [Computed free formink R ₁ , $\frac{C_{1}}{C_{1}}$ in which C_{1} =14,388 [A factor has been used for each temperature that makes R ₂ =100 st were length 590 mpl [A factor has been used for each temperature that makes R ₂ =100 st were length 590 mpl [A factor has been used for each temperature that makes R ₂ =100 st were length 590 mpl [A factor has been used for each temperature that makes R ₂ =100 st were length 590 mpl	GD 95 70 32.70 32.97 33.87 32.31 <td>60</td>	60
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